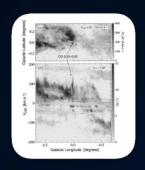
Automatic identification of High-Velocity com act clouds

Sekito Tokuyama, Tomoharu Oka, Shunya Takekawa, Masaya Yamada, Yuhei iwata, and Shiho Tsujimoto

Department of Physics, Faculty of Science and Technology, Keio University



High-velocity compact clouds (HVCCs) found in the central molecular zone of our Galaxy are a population of molecular clouds, which have compact appearance (d<10 pc) and large velocity width($\Delta V > 50 \text{ km s}^{-1}$).

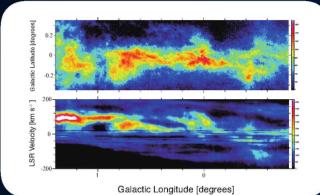


An example of HVCC CO 0.02-0.02 Oka et al. (1999)

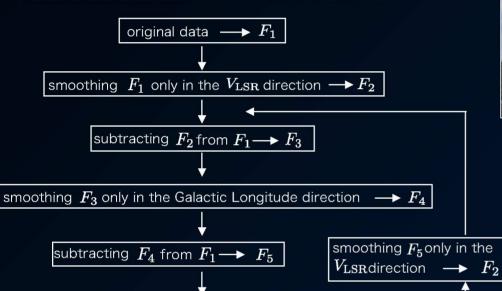


- Nobeyama Radio Observatory 45 m Telescope
- January 19-29, 2011
- $^{12}CO(J=1-0) (\nu=115 \text{ GHz})$

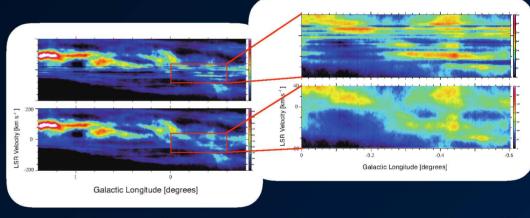
original data



1. Pressing method (Sofue 1995)



Result of Pressing method



We succeeded to reduce the local gas contamination!

2. Unsharp mask

YES

• subtracting F_7 from $F_6 \longrightarrow F_8$

Disk gas clouds have been removed?

End, Result= $\,F_5$

ullet F_8 is the result of Unsharp mask

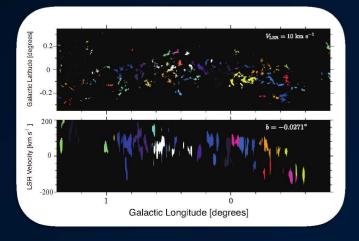
resul

Compact, broad-velocity-width features became prominent.

• smoothing F_5 which is the result of Pressing method in the $V_{ m LSR}$ direction \longrightarrow F_6 • smoothing F_6 in the Galactic Longitude and Latitude $\longrightarrow F_7$

NO

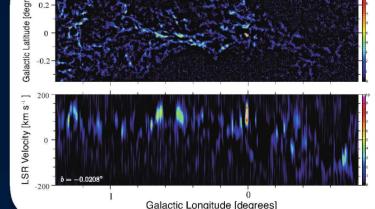
3. Modified CLUMPFIND



- We identified HVCCs candidates.
- Color shows identification numbers.

115 HVCC

candidates are Identified!



FUTURE

With $^{12}CO(J=3-2)$, $C^{18}O(J=1-0)$ and $^{13}CO(J=1-0)$, we are going to make a complete HVCC list and perform statistical analysis!