

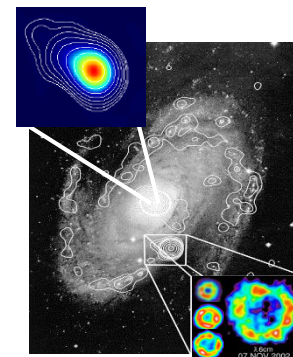
Activities of Shanghai VLBI correlator

Wu Jiang, Zhi-Qiang Shen, Ru-Sen Lu
and correlator group at Shanghai Astronomical
Observatory (SHAO)

2019/05/20

Introduction

□ Correlator is the very important part of VLBI technology.



e-VLBI/diskpacks

ST1



ST...



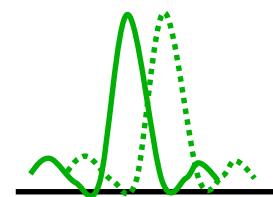
STn



Hardware Corr



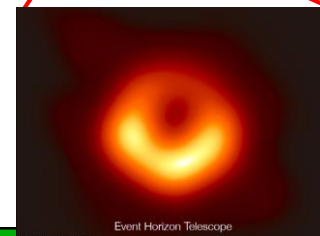
Correlator



Fringe!



Software Corr



TByte

GByte

MByte
Page 2

Introduction

□ **DiFX correlator** is developed by Adam Deller/Walter Briske (2007) and maintained by a global correlator team. Widely and formally used at EHT, MPIfR, LBA, VLBA, RadioAstro, IVS etc. for both astrophysics and geodesy.



Introduction

□ SHAO's DiFX correlator platform



Build in Dec. 2014

3 Mark6 units (2019)



Introduction

□ Deployment

Software

DiFX2.2/2.3/2.4.1/**2.5.2**, HOPS3.8/3.10/3.12/3.18/**3.20** etc.

Hardware

Head nodes **2 head nodes**, 20 CPU, each head node manages 10 computing nodes

Computing nodes **400 cores**, 20 computing nodes, 20 CPU each node, Intel Xeon E5-2660 v3 2.6GHz, 64GB RAM.

Networks **56Gb infiniband** network for cluster and RAID, 10Gb/1Gb Ethernet for cluster, RAID and Mark5.

Storage system **432 /1052 TB**, RAID 6 and parallel file system.

Mark5/6 **7 units**, 1 Mark5A, 2 Mark5B, 1Mark5B+, 3Mark6

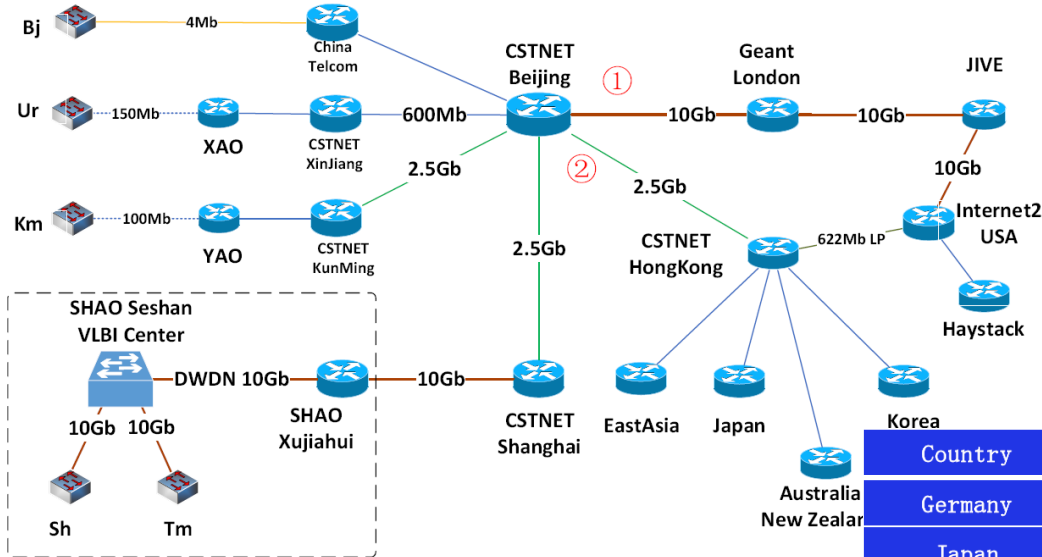
Achieve a speed of 10 stations, 1Gbps/station. More stations allowable at a relative low correlation speed.



Introduction

❖ Via CSTNET(Chinese Science&Technology Network, CAS)

❖ IPv6 supported in CVN



e-VLBI network conditions

Almost a real time connection.

Country	Institutes	Network B/W	To/From Shanghai
Germany	Bonn MPRIFR	900Mb/s	800Mb/s
Japan	NICT	10Gb/s	1Gb/s
Japan	GSI	10Gb/s	1Gb/s
South Korea	NGII	1Gb/s	800Mb/s
Italy	IRA	10Gb/s	800Mb/s
South Africa	Hartebeesthoek	10Gb/s	550Mb/s
Australia	University of Tasmania	10Gb/s	800Mb/s
New Zealand	Auckland University of Technology	10Gb/s	800Mb/s
Brazil	IPNE	1Gb/s	200Mb/s
Netherlands	JIVE	10Gb/s	1Gb+/s
Russia	IAA	1Gb/s	500Mb/s
USA	Haystack	1Gb/s	600Mb/s
Malaysia	University of Malaya	100Mb/s	90Mb/s

Operations and applications

1. Flexible: Support Mark4, Mark5B, VDIF format as input, provide Mark4 and FITS-IDI outputs.
2. Pulsar binning/gating, zoom maser line, multi-phase center correlation.
3. e-connected.

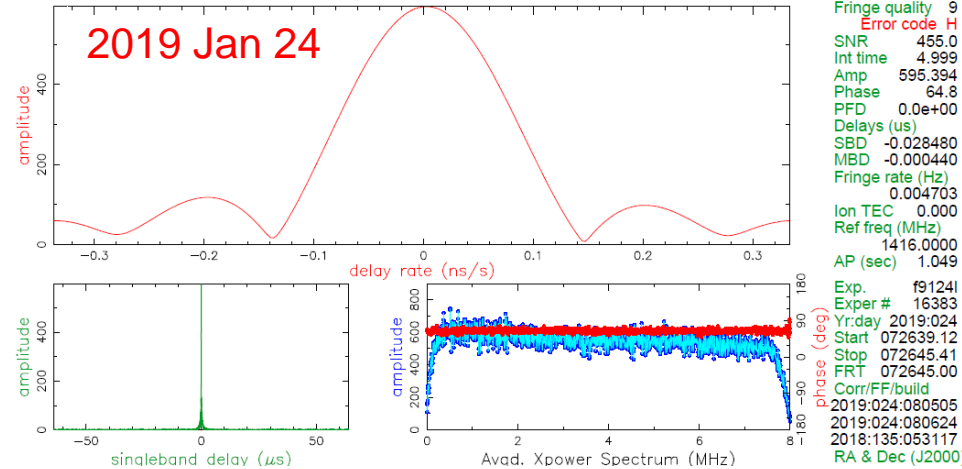
□ Routine operations

1. **Astrophysics**, CVN observations, east Asia VLBI, joint observations.
2. **Geodesy**, serve as an IVS correlator since 2015. domestic geodetic observations and joint VLBI astrometry programs.
3. Other test experiments, **data exchange buffer** etc.

Operations and applications

Mk4/DiFX fourfit 3.18 rev 2251

3C454_3.0H281P, No0100, LW
FAST500 - TIANMA65, fgroup L, pol LL



CVN Chinese VLBI Network

Observing Band:

- Bj S/X
- Km S/X, C
- Sh L, S/X, C(6cm)
- T6 L, S/X, C, X/Ka, K, Q
- Ur L, S/X, C(6cm), K

Backend:

CDAS/DBBC, Mark5B, up to 2Gbps

Baseline Length:

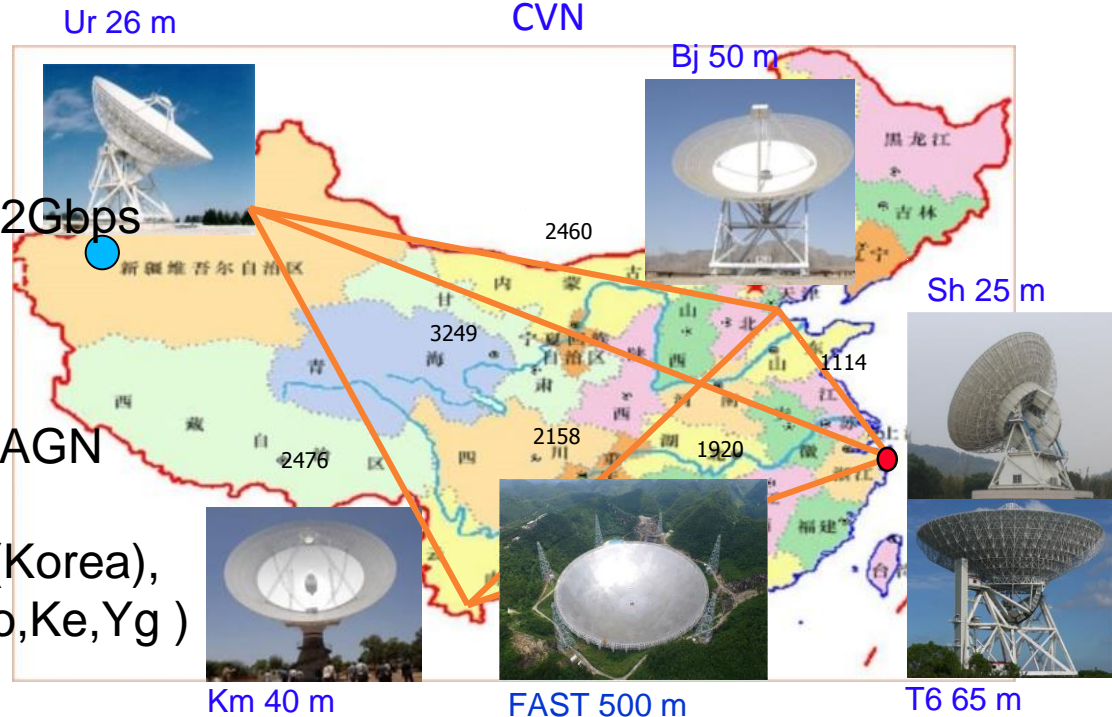
~1100km-3250km

Scientific programs:

Geodesy, astrometry, pulsar, AGN

Joint Observations:

Kashima34 (Japan), Sejong (Korea),
Australian telescopes (Hb, Ho, Ke, Yg)



Operations and applications

EAVN

Observing Band:

C(6.7GHz), K, Q

Recording rate:

1Gbps

2Gbps(future)

Stations(common):

CVN(3),

KVN(3),

VERA(4)

Baseline Length:

~200km-5500km

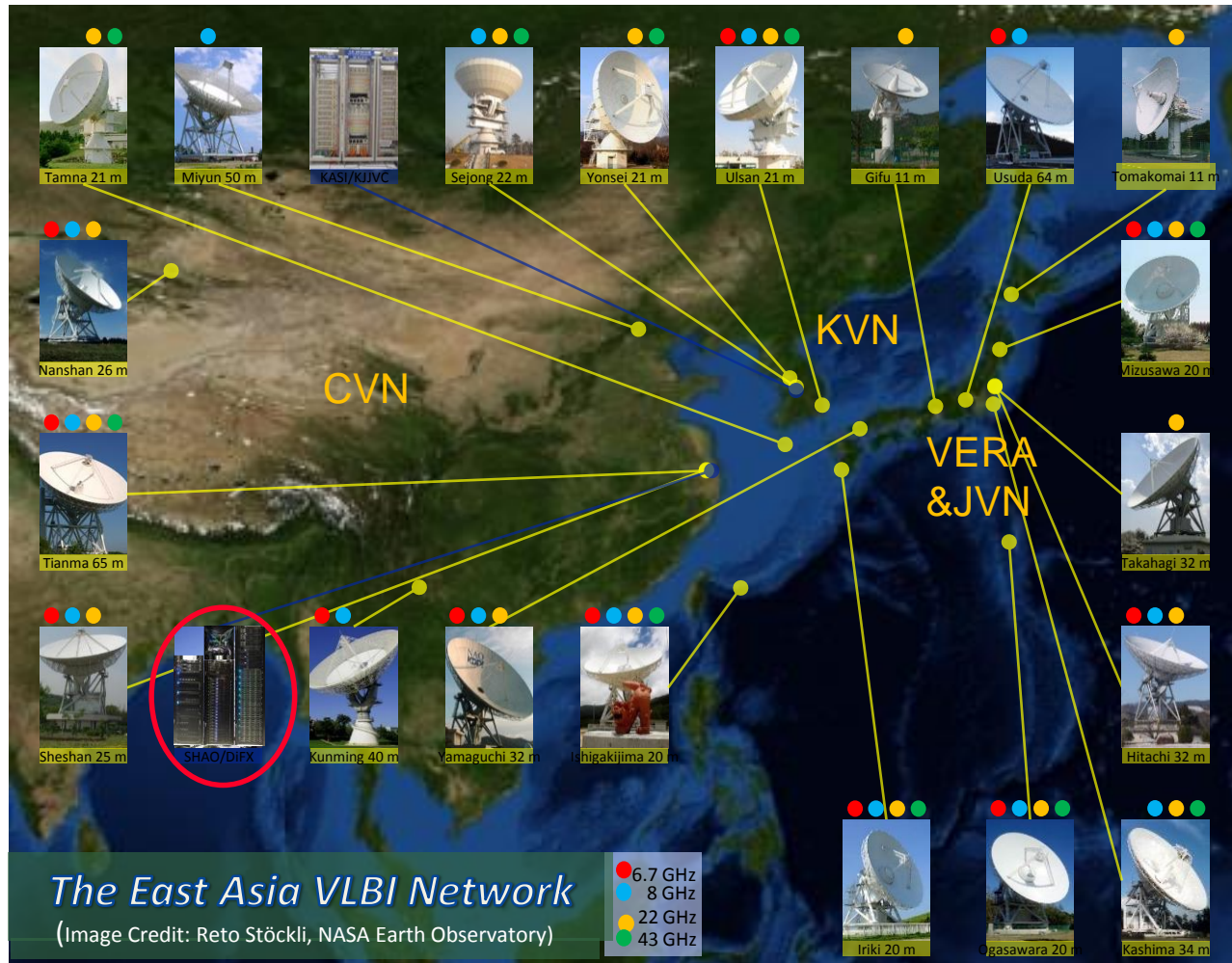
Science work groups:

AGN

Astrometry

Evolved star

Star formation



Early fringe tests for EAVN.

EAVN open-use started in 2018!

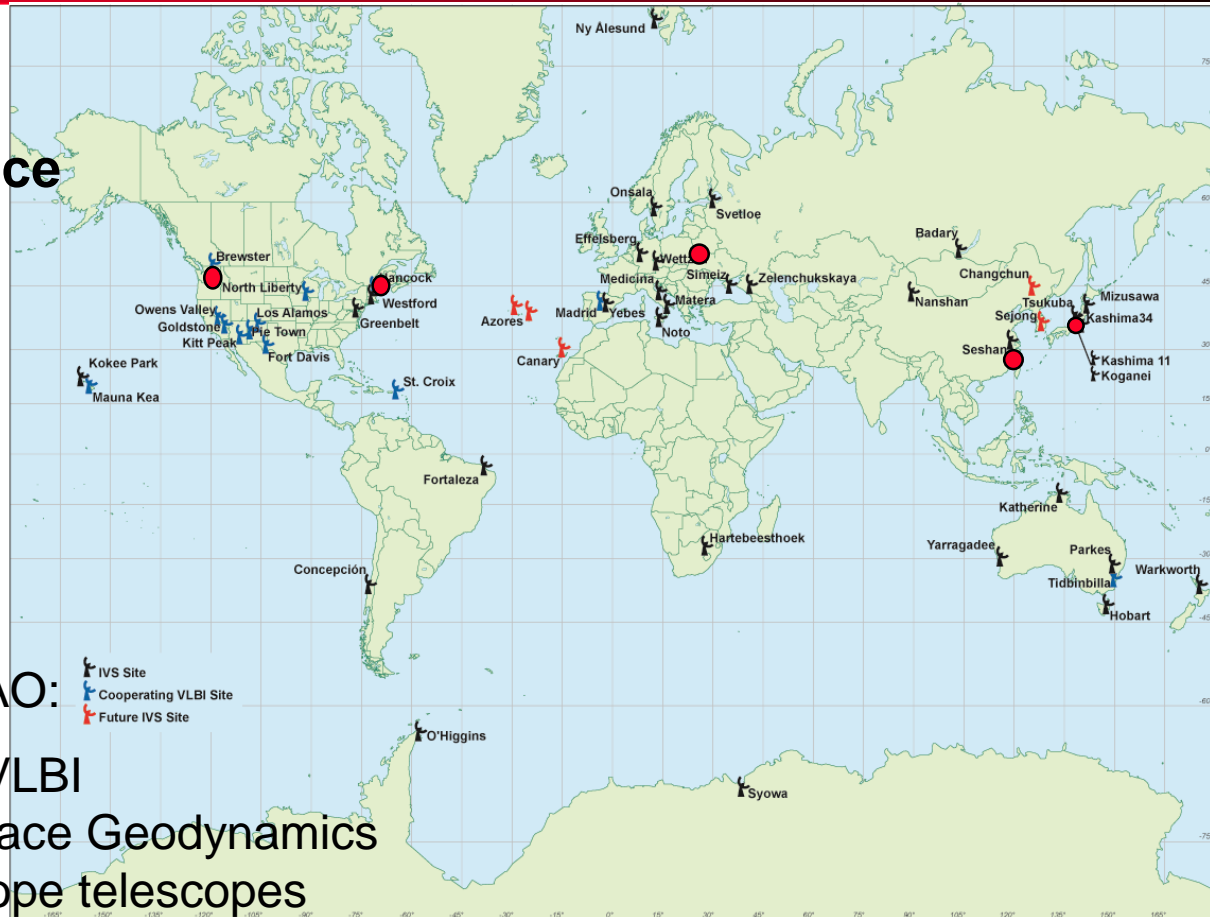
Operations and applications

IVS International VLBI service

1/5 of the main IVS correlator,
Bonn (German)
Haystack (USA)
GSI(Japan)
Shanghai(China)
Washington(USA)

Session series correlated at SHAO:

AOV Asia Oceania VLBI
APSG Asia-Pacific Space Geodynamics
AUA* Austral AuScope telescopes
CRF Celestial reference frame
CRDS CRF deep-south
R&D Research and development



Operations and applications

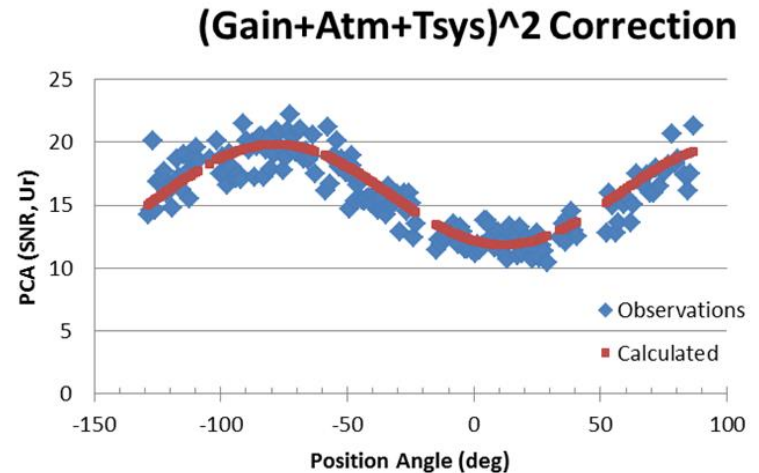
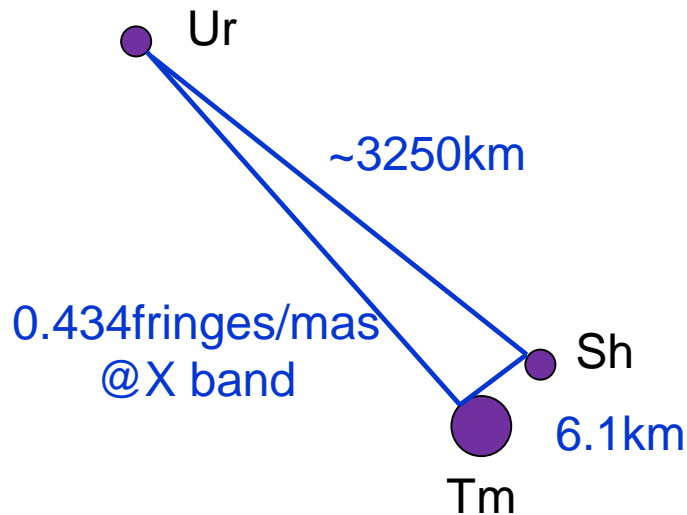
□ Routine operations

	CVN	Joint	IVS
2015	10 Continuum, Pulsar, Geodesy	10 EAVN(FT), Astrometry	10 AOV, APSG, CRF
2016	20 T6 Rx test, Continuum, Geodesy	8 EAVN(FT), FT, Astrometry	26 AOV, APSG, AUA, AUG, CRDS, CRF, R&D
2017	Astrometry, Geodesy, Pulsar	EAVN, FT	31 AOV, APSG, AUA, CRDS, CRF, R&D



Scientific Program

AGN



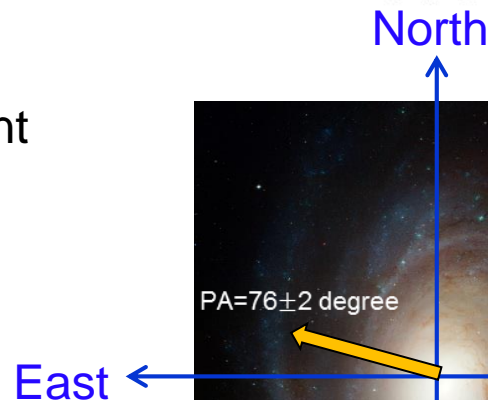
Pseudo Closure Amplitude(PCA) measurement

$$PCA_{Ur} = \frac{\rho_{SU} \cdot \rho_{TU}}{\rho_{ST}} \approx S_o \cdot r^2 \cdot S_{Ur}$$

So : Source Flux, S_{UR} : Station Sensitivity

small dish – calibration assistant

big dish – sensitivity booster



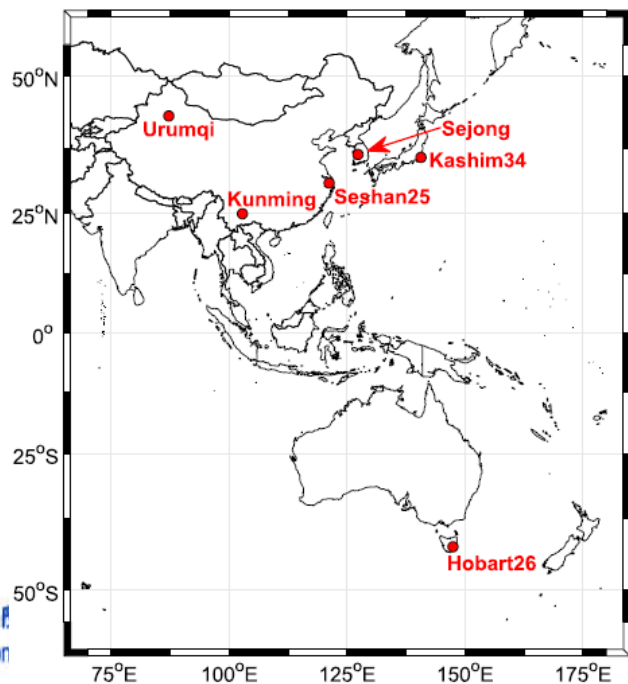
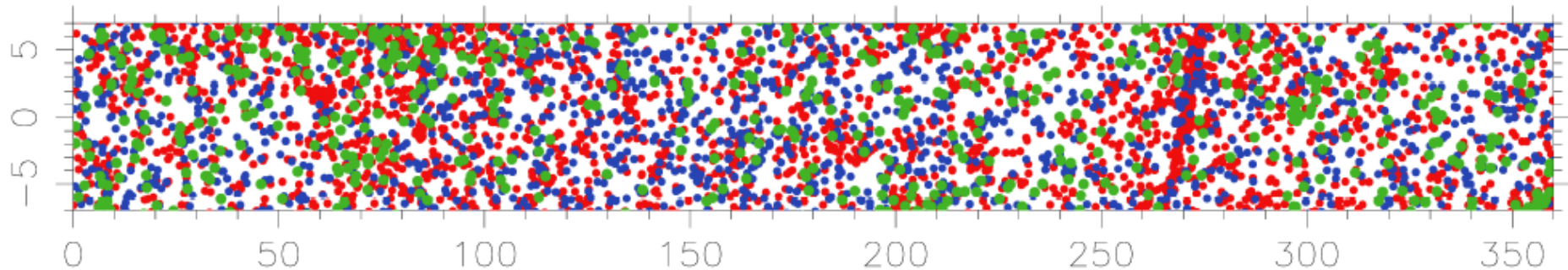
M81



Scientific Program

Ecliptic plane survey

Among 3321 sources (red+green), 556 sources detected (green). (Shu F C, session2-3)

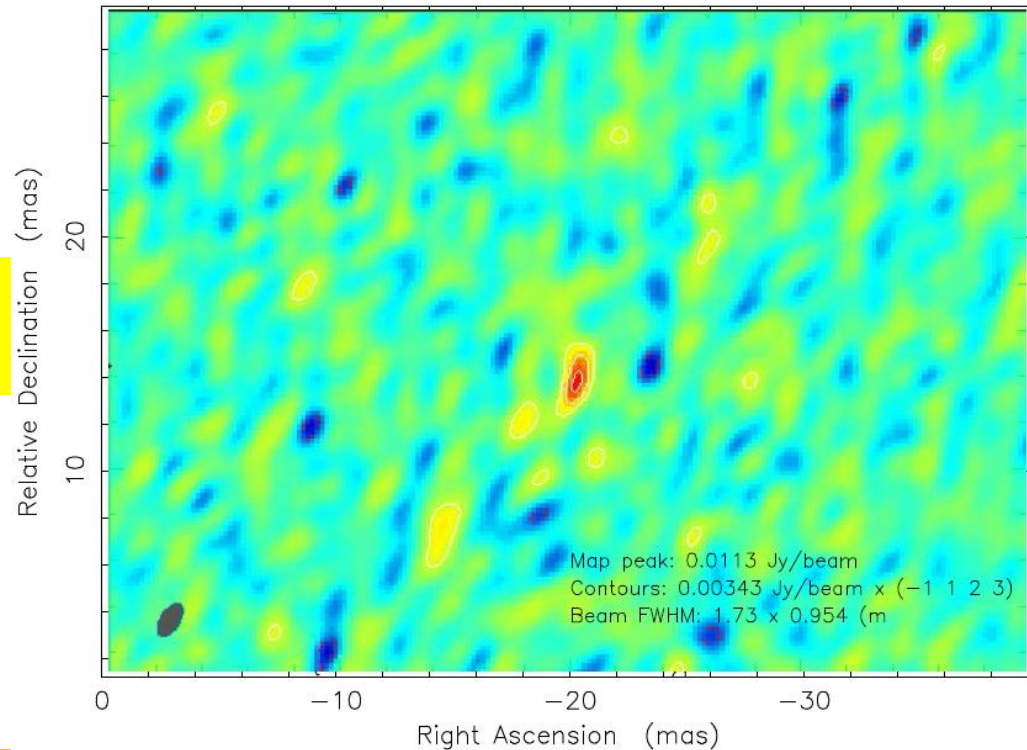
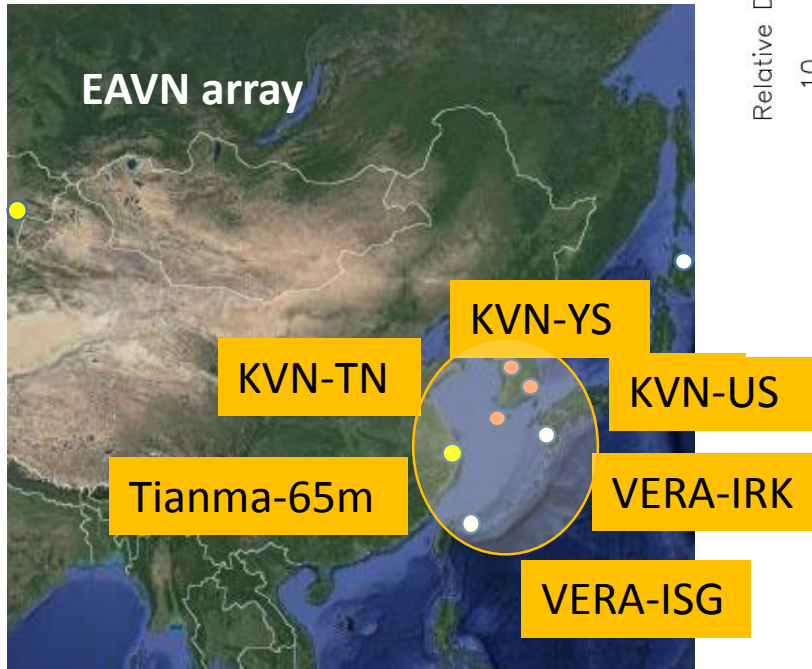


Hybrid recording and correlation

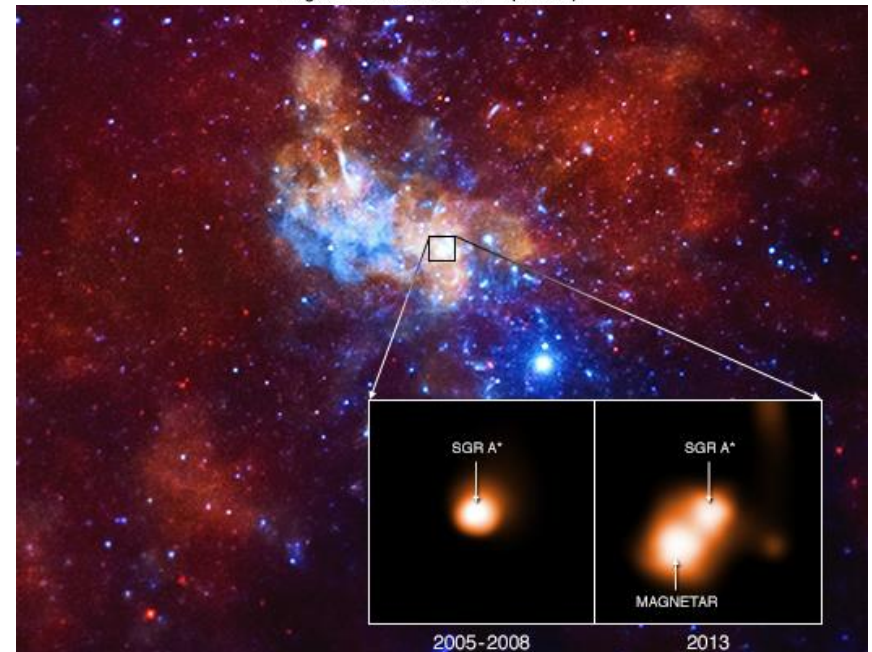
ShKbKmUr $16IF \cdot 32MHz \cdot 2bit = 2Gbps$
Kv(Sejong) $16IF \cdot 32MHz \cdot 1bit = 1Gbps$
Ho(Hobart26) $16IF \cdot 16MHz \cdot 2bit = 1Gbps$

GC magnetar study

In beam (2.4 arcsec) phase-referencing imaging of magnetar

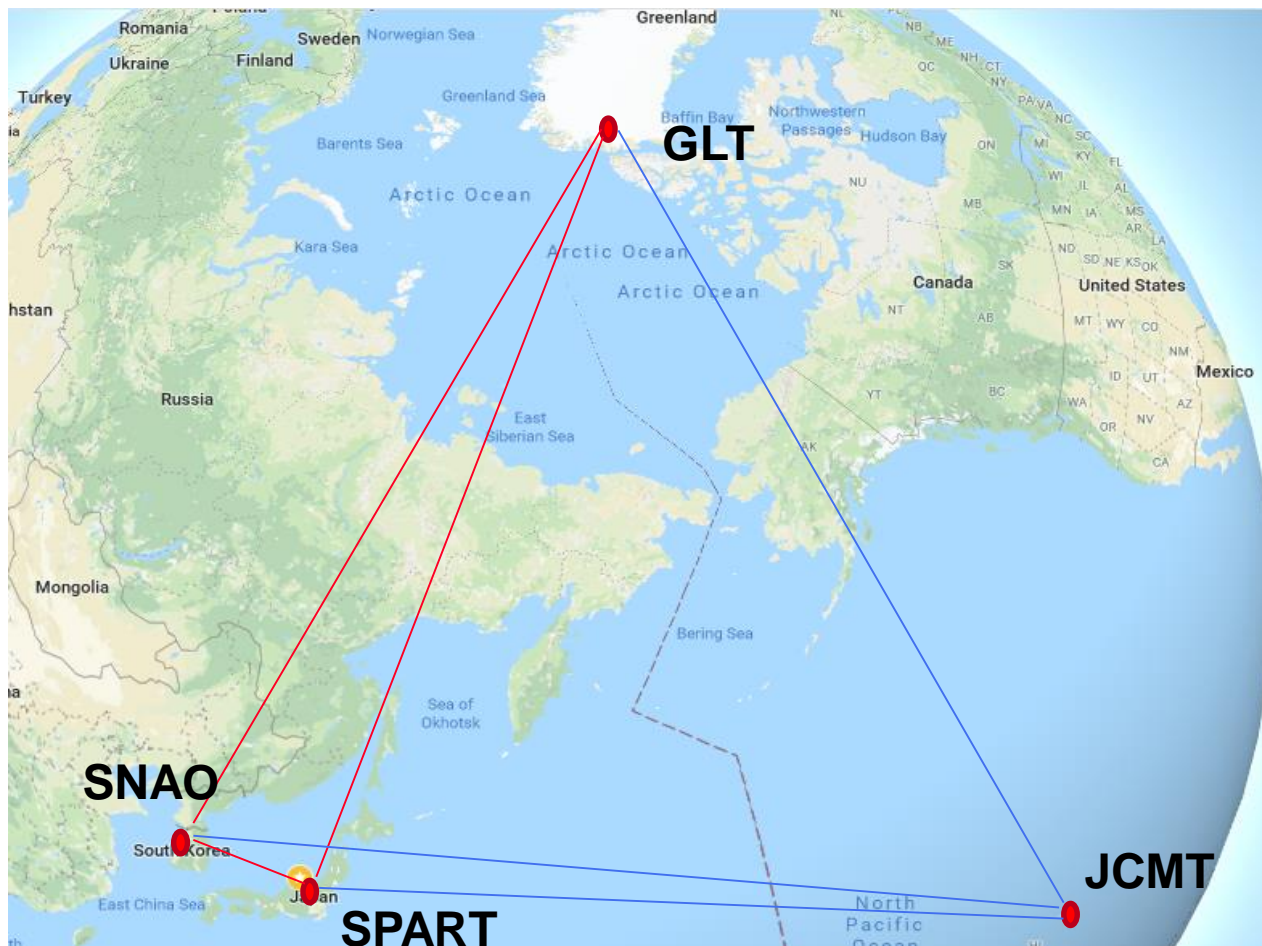


A 43G detection with EAVN, multi epoch observations in total on SgrA*/Magnetar with EAVN in 2017/2018/2019 EHT campaign.



EA 1mm VLBI

The **first EA 1mm VLBI** test observations (2019.03.17-19)! The data are collecting now.



Discussion: related to EAO Sub-mm Futures

The correlator at SHAO definitely can server for both test and scientific EA sub-mm VLBI observations!

Any requirements for the correlator?

For future sub-mm VLBI

More stations, higher frequency, new technology...

Science driven programs...

Connect to the outside region...

Thank you for your attention!
谢谢!

