

A night photograph of a traditional Chinese building with a star trail background. The building has multiple stories with balconies and is illuminated from within. The sky is dark blue with numerous white streaks representing star trails. The text is overlaid in yellow.

# The Future of the EAO JCMT-Transient Survey

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# Era of Transient Surveys (ASAS-SN, PTF, Gaia, LSST)

Accretion bursts in youngest stages of stellar growth:  
Not detectable in optical, near-IR surveys



The East Asian Observatory JCMT-Transient Survey:  
the first long-term sub-mm monitoring program

# Transient Team (~80 people total)

## Coordinators

Gregory Herczeg (co-PI; PKU/China)

Doug Johnstone (co-PI; NRC/Canada)

Jeong-Eun Lee (KHU/Korea)

Yuri Aikawa (Tsukuba/Japan)

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Vivien Chen (NTU/Taiwan)

Jennifer Hatchell (Exeter/UK)



Yong-hee Lee, PhD student,  
Kyung-Hee Univ.

Data PhDs: **Steve Mairs (Victoria), Hyunju Yoo (Chungnam/KASI)**, Bhavna Lalchand (NCU), Yong-hee Lee (Kyung-hee)

Modeling PhDs: **Ben MacFarlane (UCL)**, Giseon Baek (Kyung-hee)

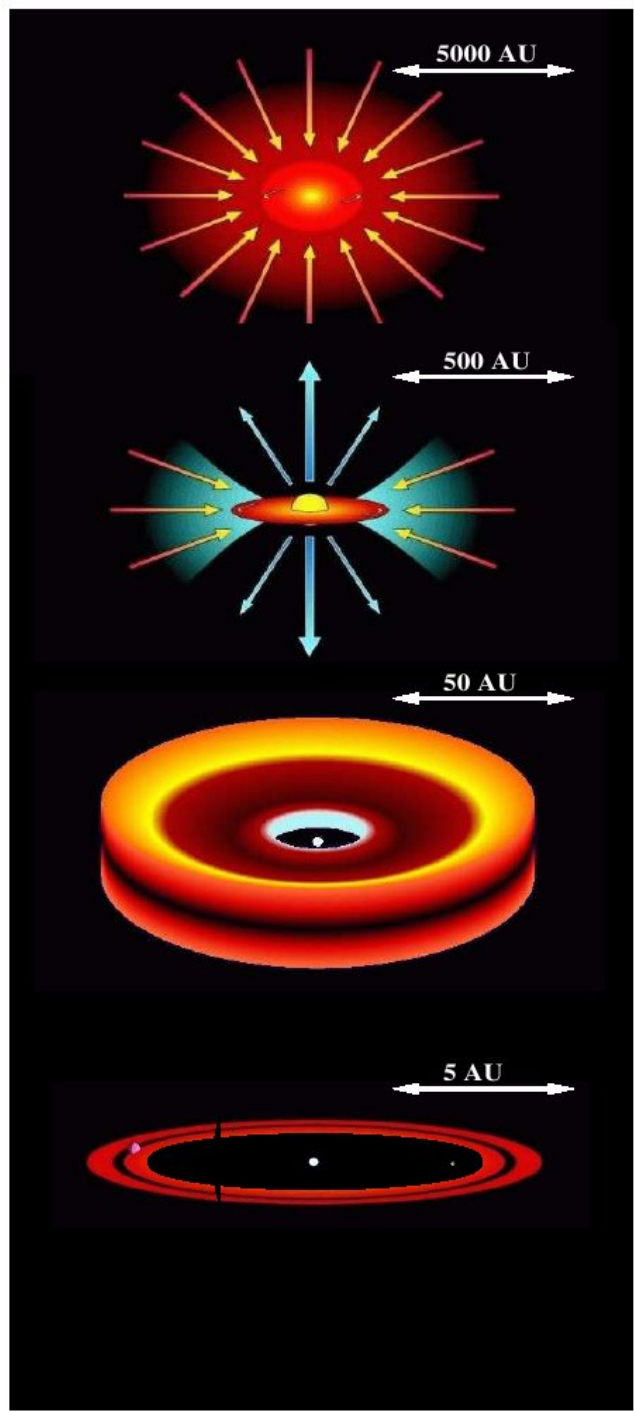
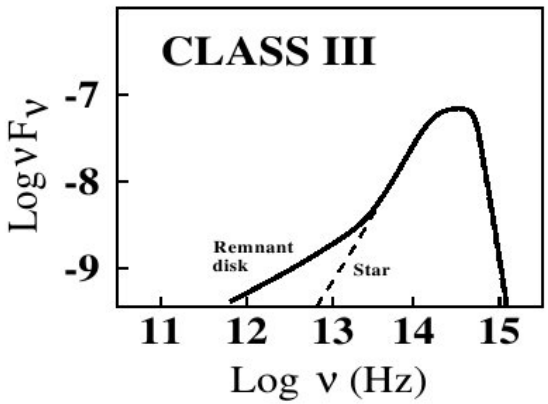
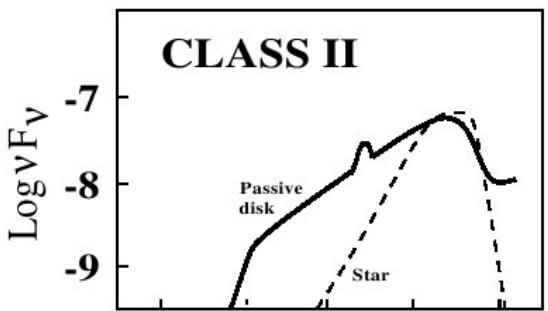
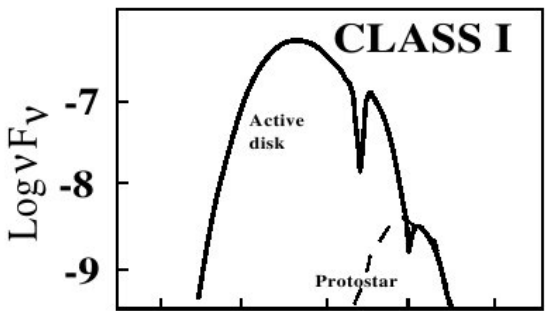
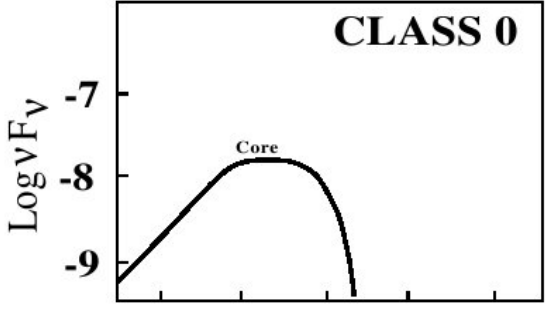
Related PhDs: Logan Francis, Sunkyung Park

Other significant contributors: Wen-Ping Chen, Sung-ju Kang, Seokho Lee, Graham Bell, Sarah Graves, Carlos Contreras-Pena, Tim Naylor, Miju Kang, Oscar Morata, Dimitris Stamatellos, Miju Kang, Aleks Scholz, Shu-Ichiro Inutsuka, Jan Forbrich, Helen Kirk

Friends of the program: Watson Varricatt (UKIRT), Klaus Hodapp (UH), Patrick Sheehan & John Tobin (NRAO)



# Protostellar Evolution

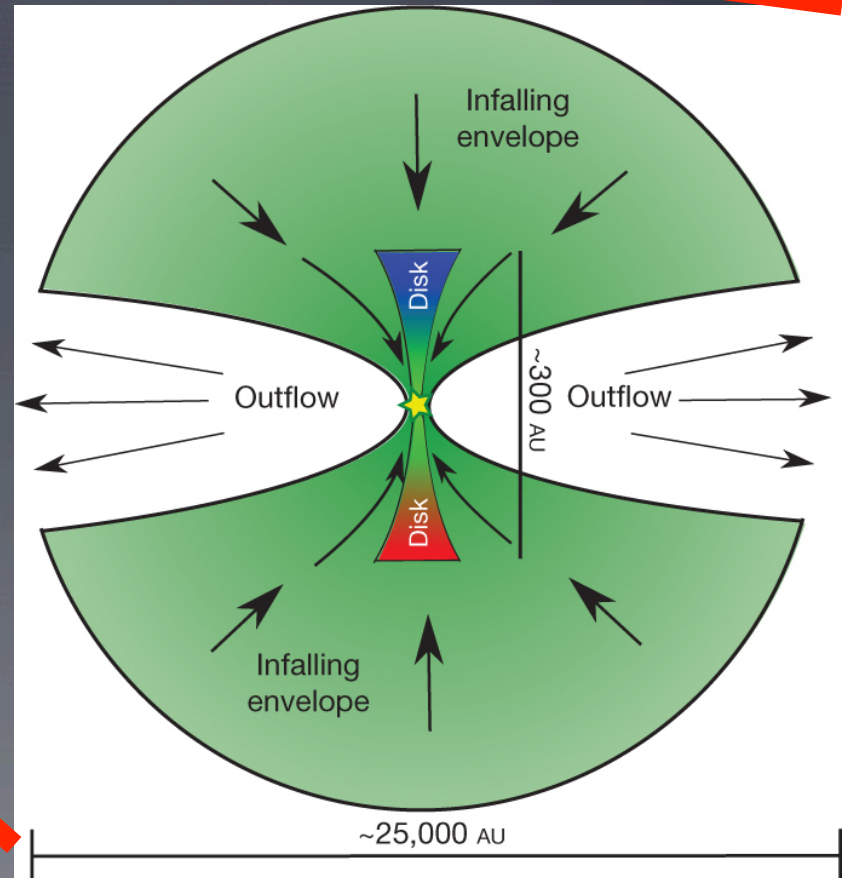
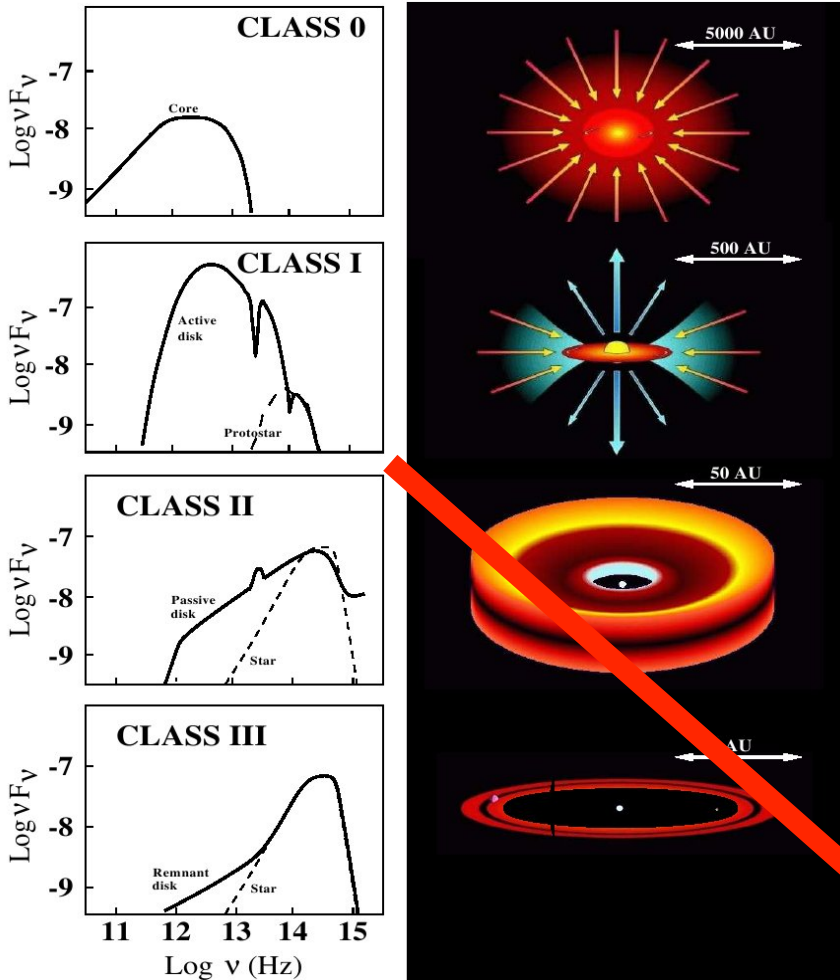


Cartoon from van Boekel 2005

# Stars grow during protostellar phase

$$L_{\text{tot}} = L_{\text{acc}} + L_{\text{phot}}$$

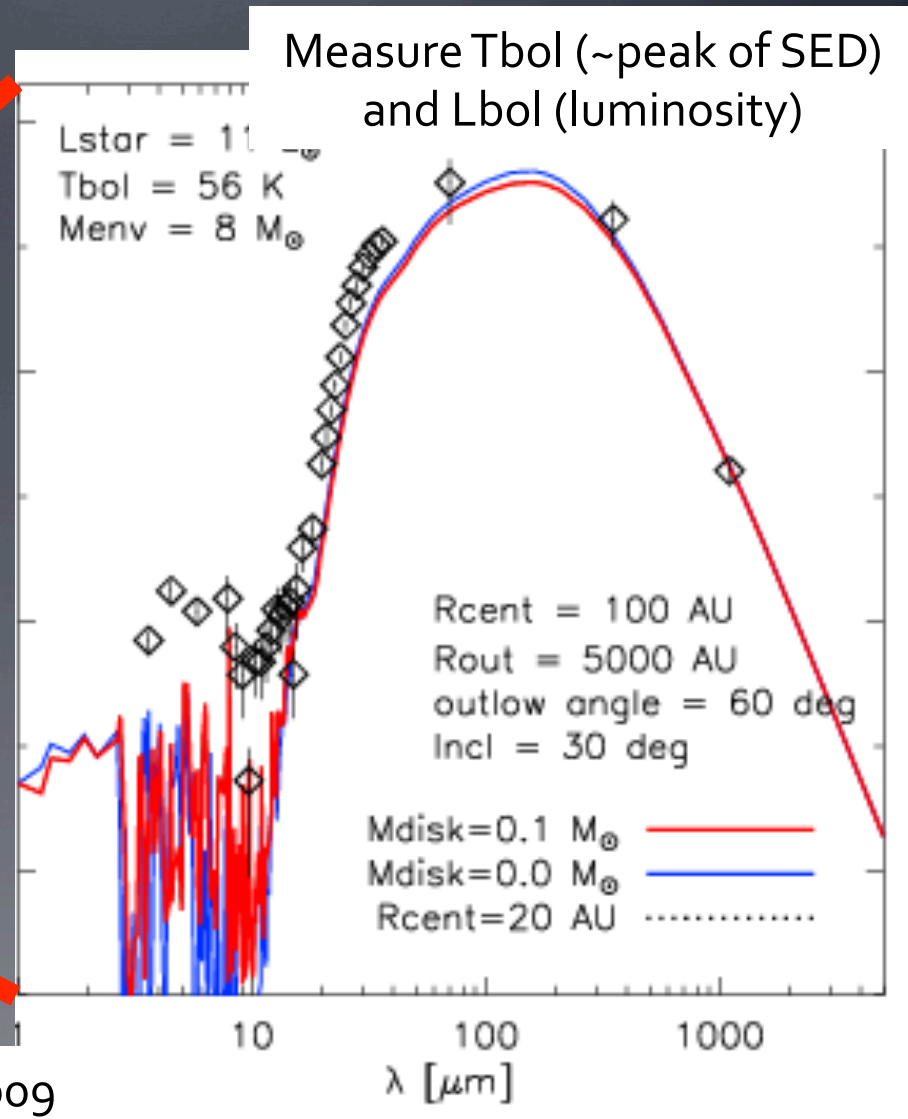
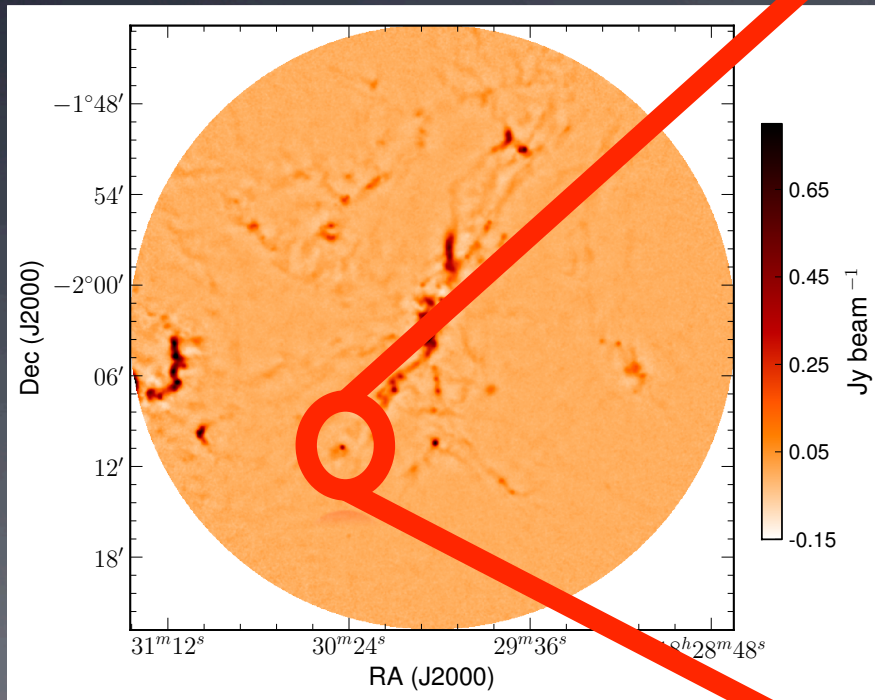
Scattered by dust



Cartoon from Isella 2006

Cartoon from Tobin+2012

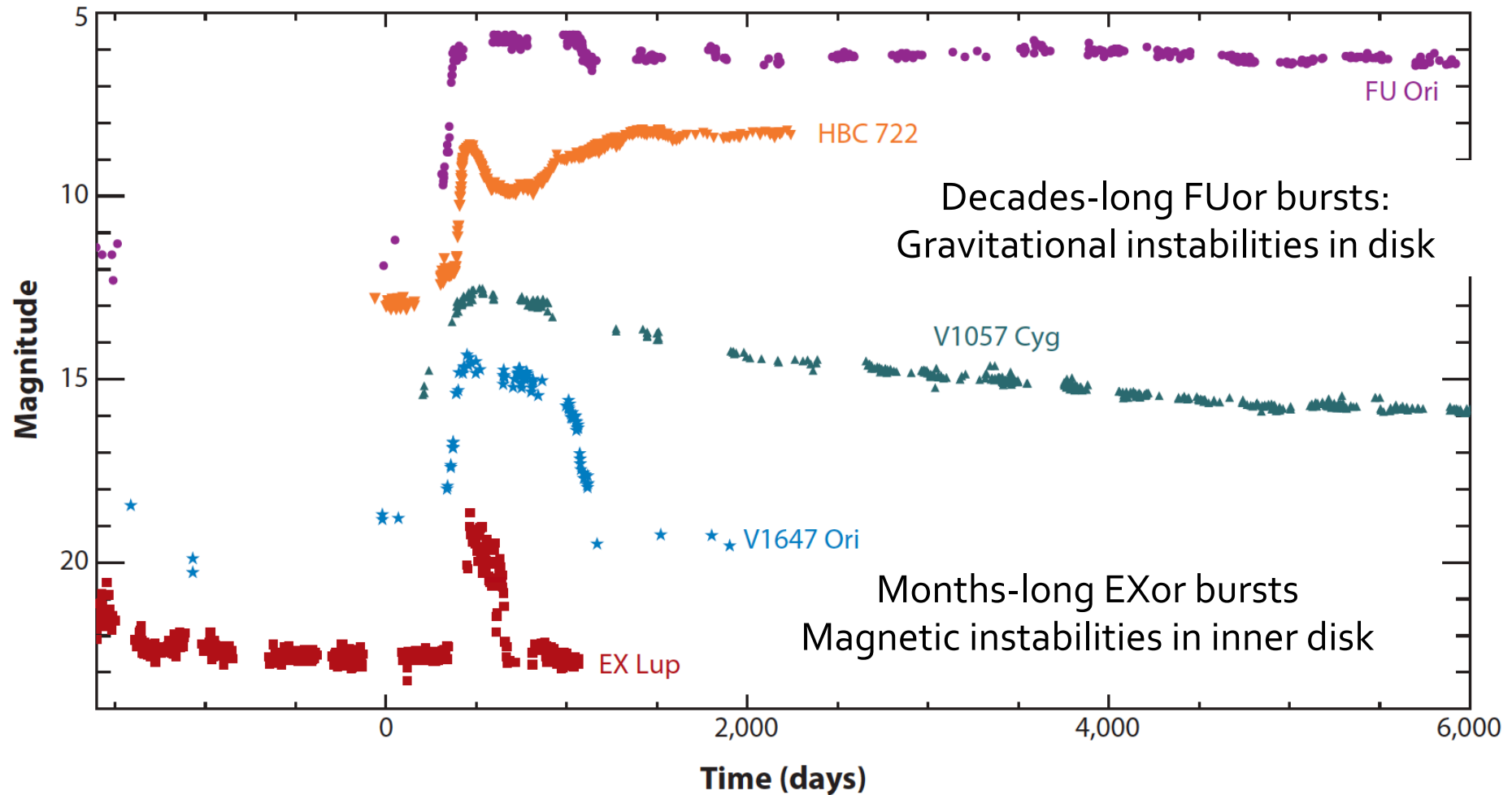
# Spectral Energy Distribution



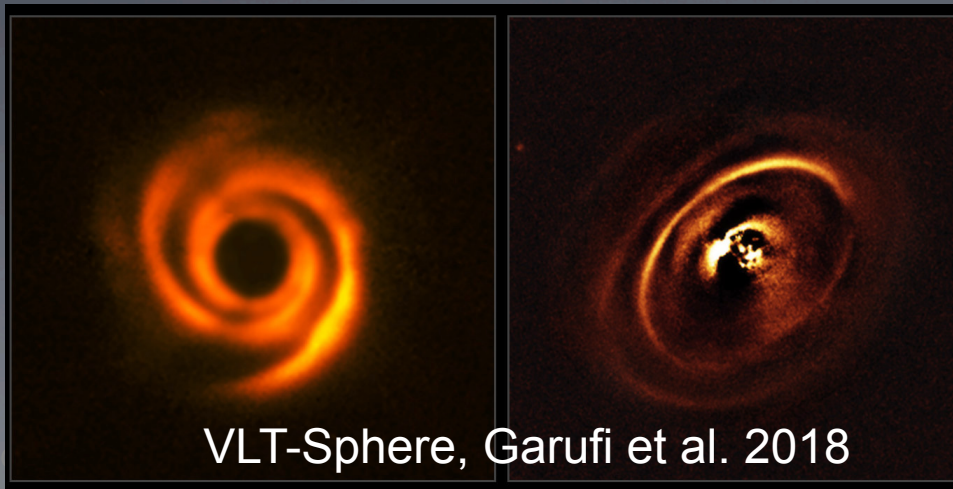
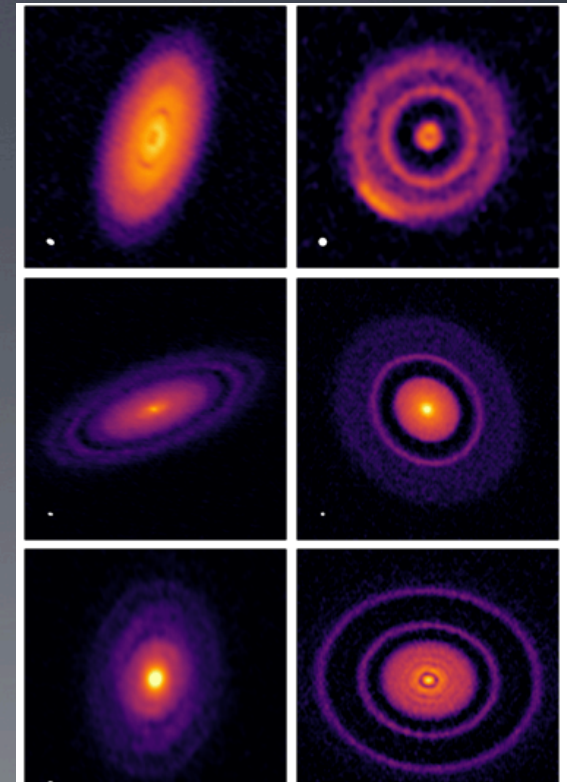
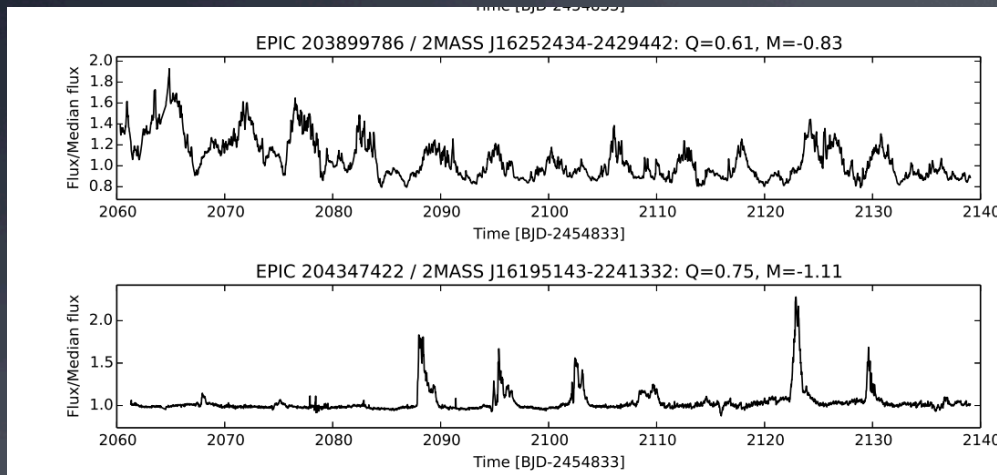
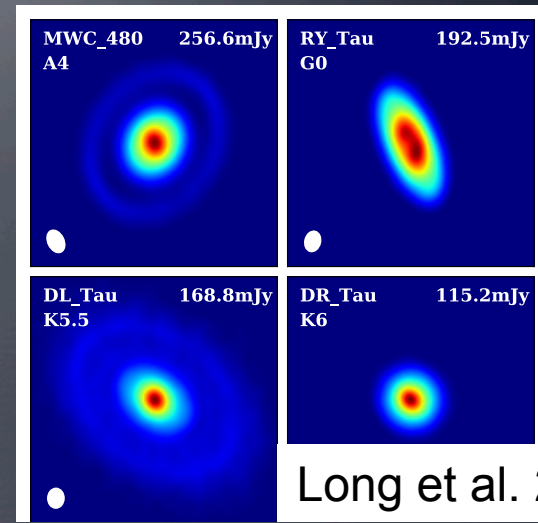
Enoch+2009

# Disk instabilities and YSO variability

(adapted from Kospal+2011)



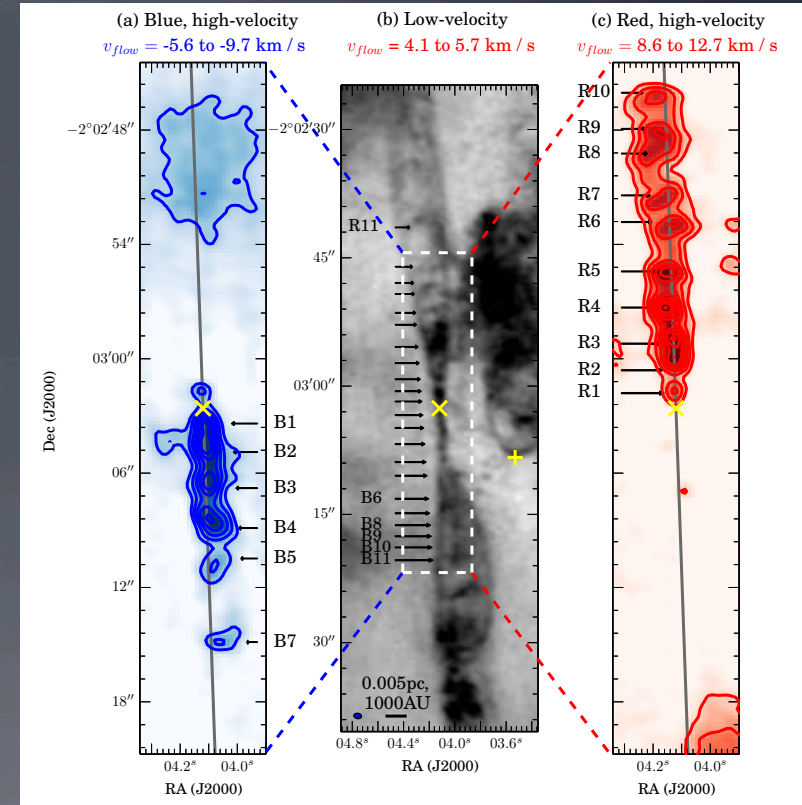
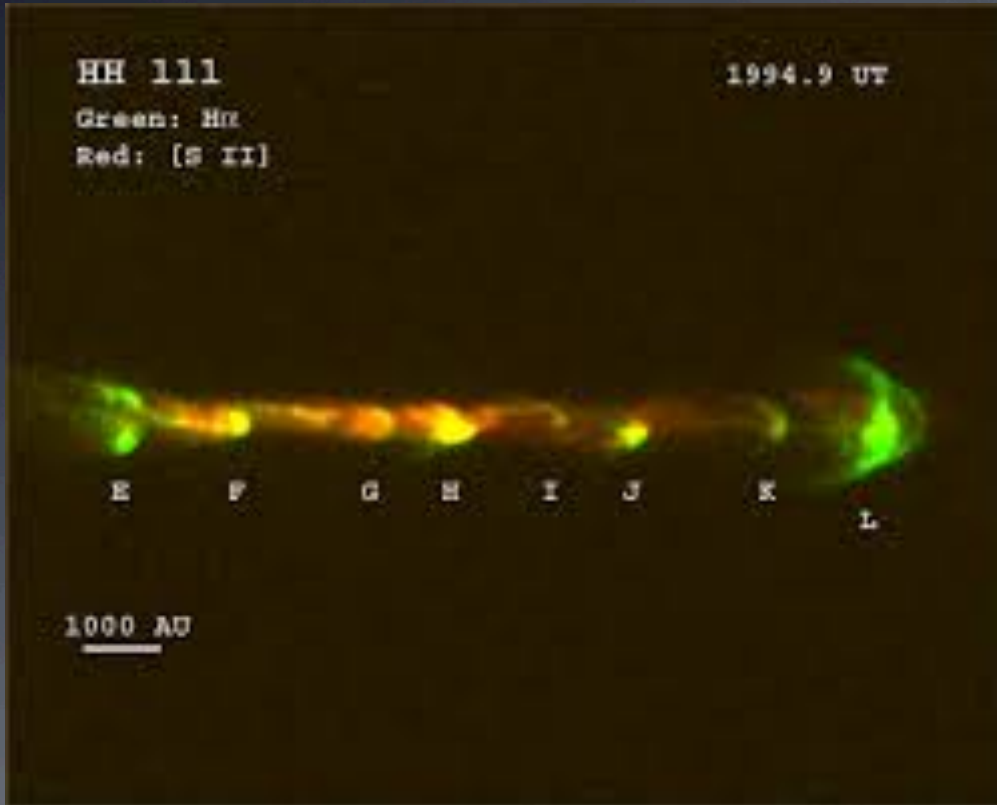
# Disk instabilities and substructures



Andrews, Huang, et al. 2018



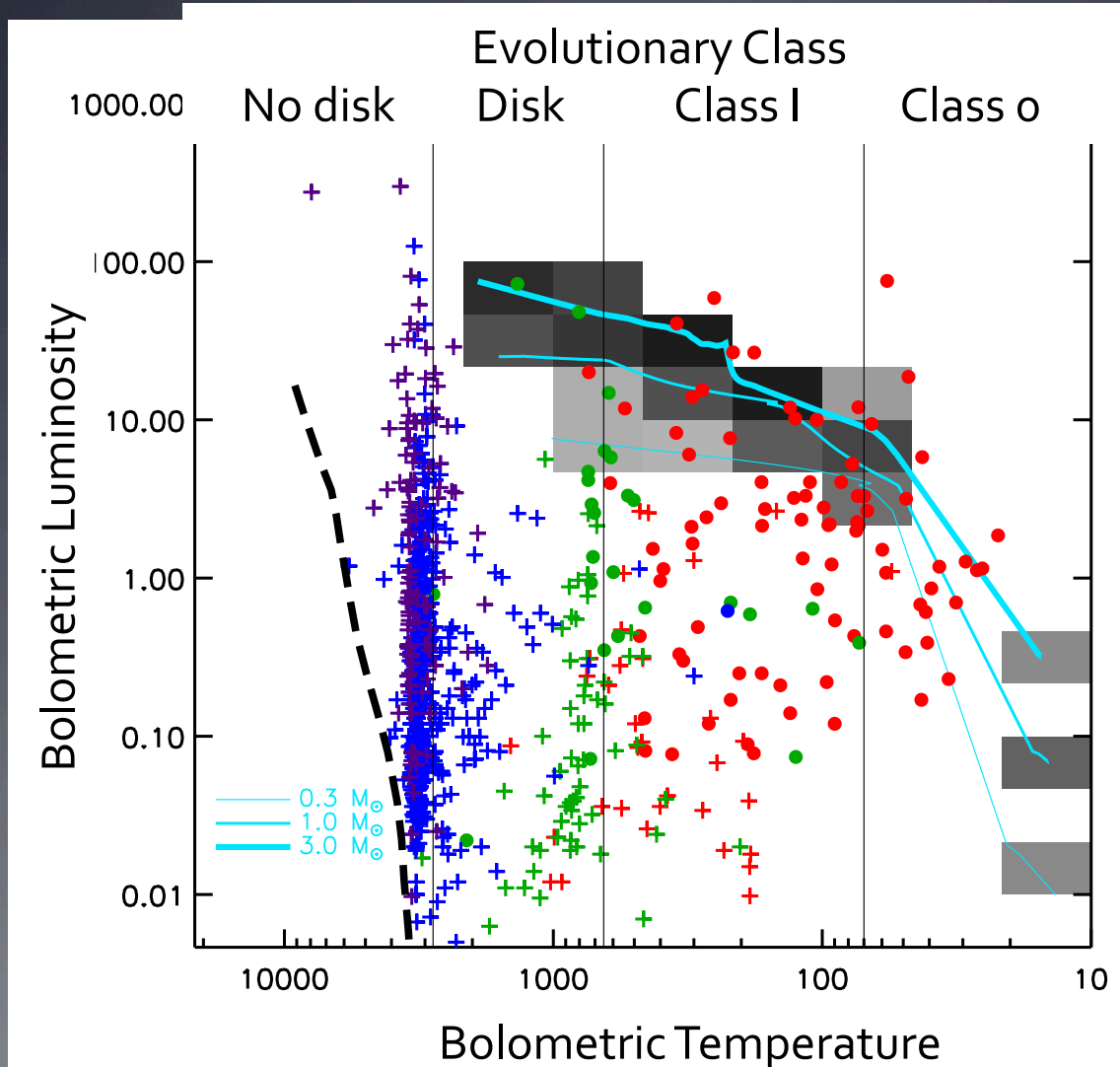
# Periodic jet shocks and relationship to accretion?



Jet shocks of YSOs  
(e.g., Reipurth 1989; Hartigan+2011; Plunkett+2015)

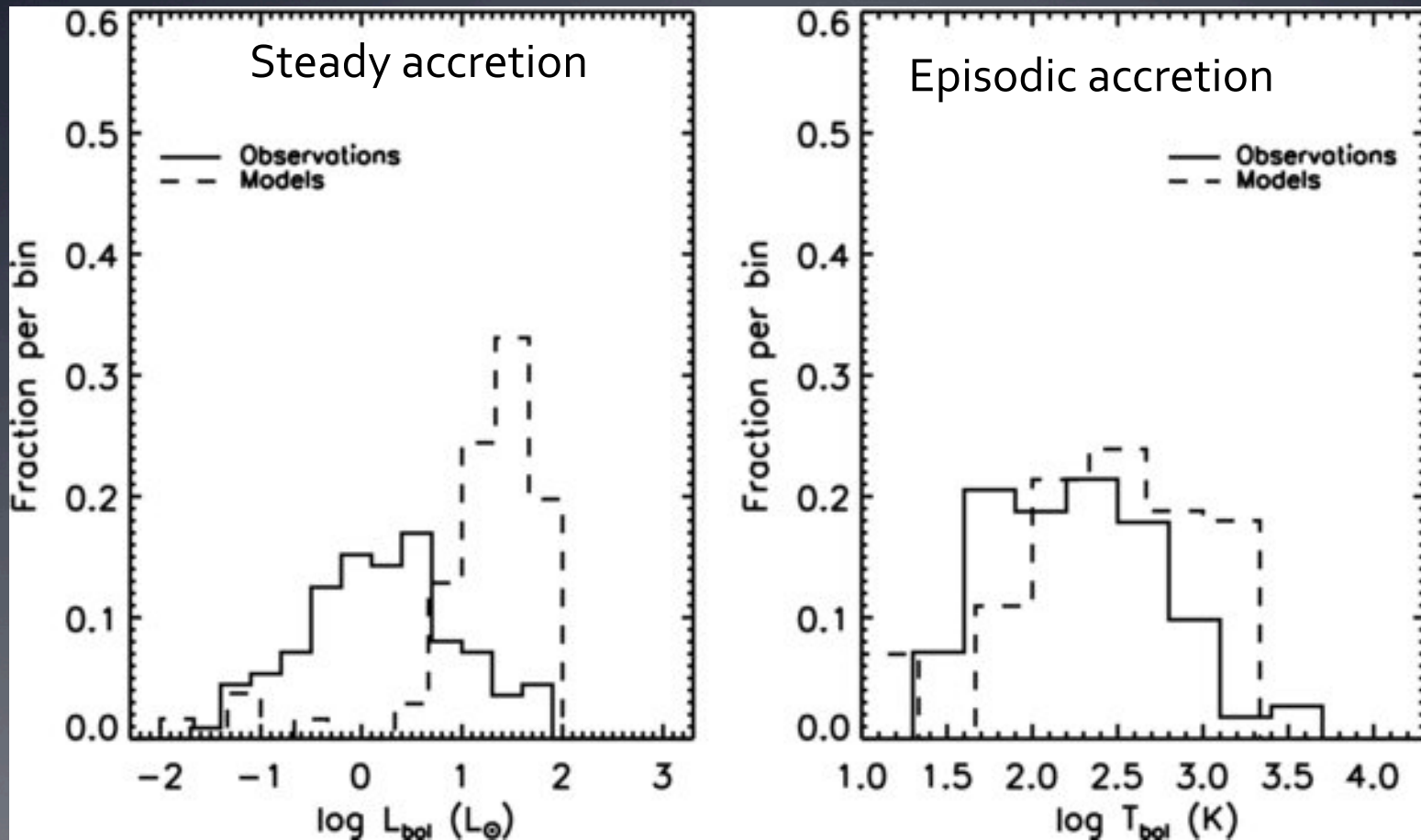
# Luminosity Problem

(Kenyon et al. 1990; Dunham et al. 2009)



# Episodic bursts of accretion

(Kenyon et al. 1990; Dunham, Evans, et al. 2009)



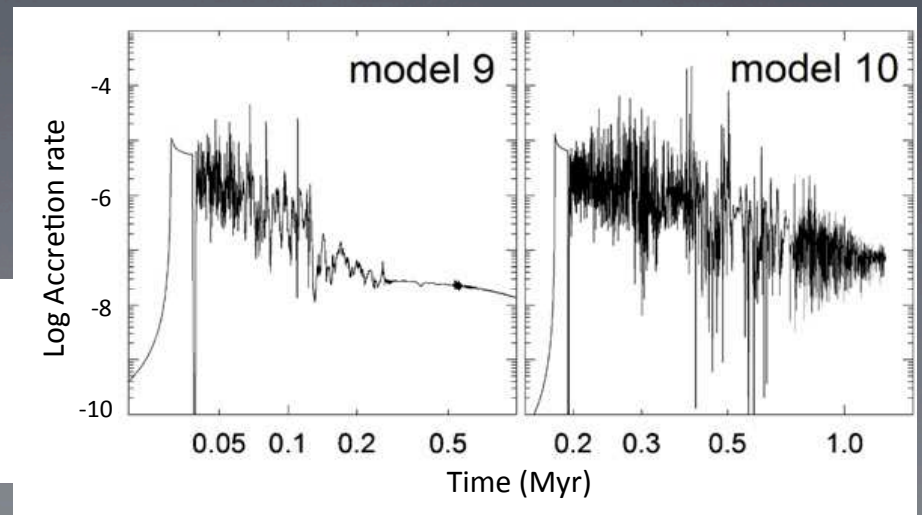
Time dependence needed; episodic accretion is likely (but not only) solution (e.g., Offner & McKee; see review by Hartmann, Herczeg, & Calvet 2016).

# Causes of outbursts

- Disk instabilities (universal)
  - Gravitational, magneto-rotational, thermal
- Binarity (e.g. Bonnell & Bastien 1992, Reipurth 2000)
- Magnetospheric instabilities (D'Angelo & Spruit 2010, Armitage 2016)
- Tidal disruption of planets or alien weaponry (Herczeg+2016)

Frequency: 1 of  $10^4$  stars from optical (Hillenbrand & Findeisen 2015; Contreras-Pena 2019)

Models from Dunham & Vorobyov (2012)  
See also, e.g., Zhu+, Bae+, Stamatellos+,  
Vorobyov+, Machida+, others



# Namakanui Detector and the Transient Survey

- Depth: more calibrators, especially disks
  - ToO easier for bursts identified in optical/IR surveys
- Field of view: more objects! More calibrators! Higher cadence?
- Faster coverage?
  - Do rapid changes in atmosphere limit calibration?
- 450 microns? Faster = much more powerful
- Better able to achieve main science goals
  - Measure sub-mm variability as a probe of accretion instabilities in very young disks
  - Evaluate variability (episodic accretion) as explanation of luminosity problem
- Challenge: Confusion limit in young regions

# Other applications related to Transient

- The variable sub-mm sky: extinction can be ignored!
  - Supernova? TDEs?
  - AGN/Quasars??
  - Stellar flares???? (coordinated for multi-wavelength opacity)
  - eta Car outbursts?
- Protoplanetary disks (primordial, debris) and proto-brown dwarfs
  - Confusion-limited in youngest regions
- Filaments
- Polarimetry: many other talks

# The immediate future of Transient

- Initial survey is finished in January
  - Several papers in preparation
  - Changes on 1-2 year timescales: we need longer-term lightcurves
- We hope to continue the survey
  - Similar fields/cadence
  - Possible changes (e.g., a high-mass SF region?)
  - Need help for some improvements in primary science; also in ancillary science (e.g., disks in our very deep coadded maps)
  - Team meeting in August 2019 at NCU, Taiwan

**If you would like to help guide the future of the survey, contact Doug, Steve, or me ([gherczeg1@gmail.com](mailto:gherczeg1@gmail.com))**