

# The distribution of magnetic field strength in Orion A region

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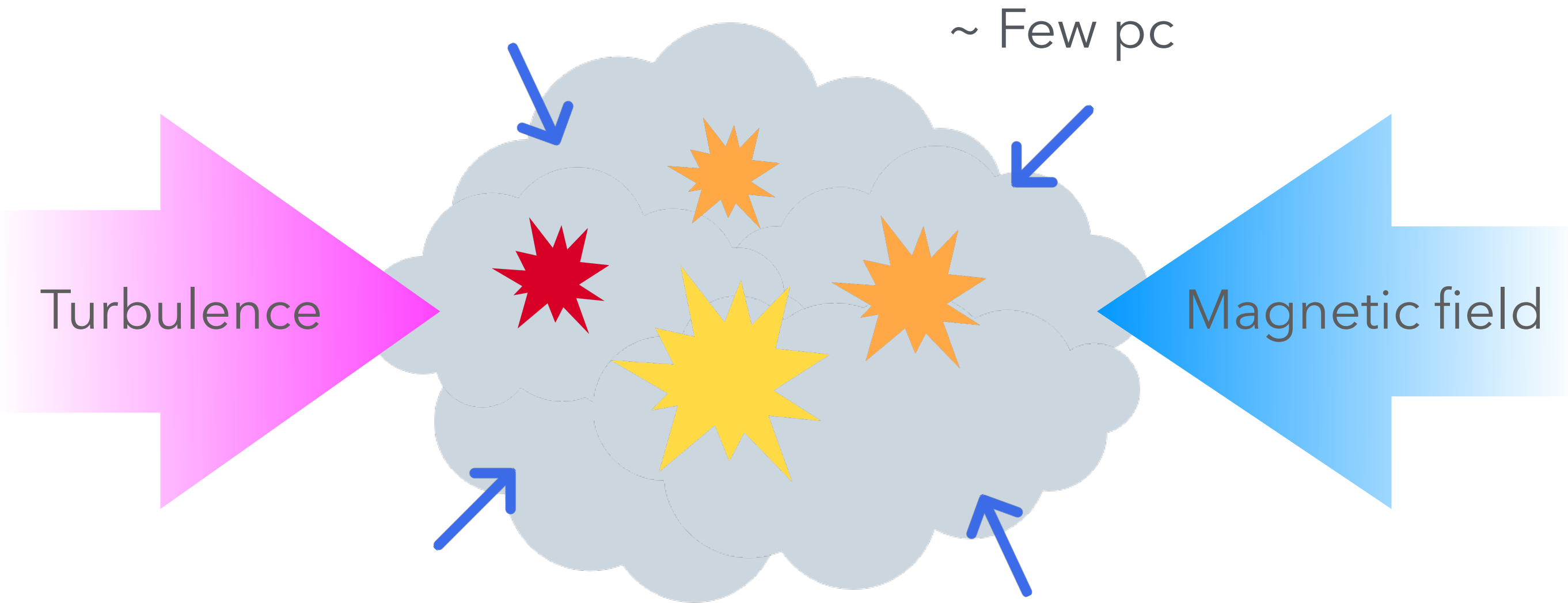
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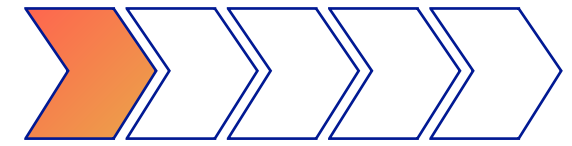
SUMMARY

# INTRODUCTION

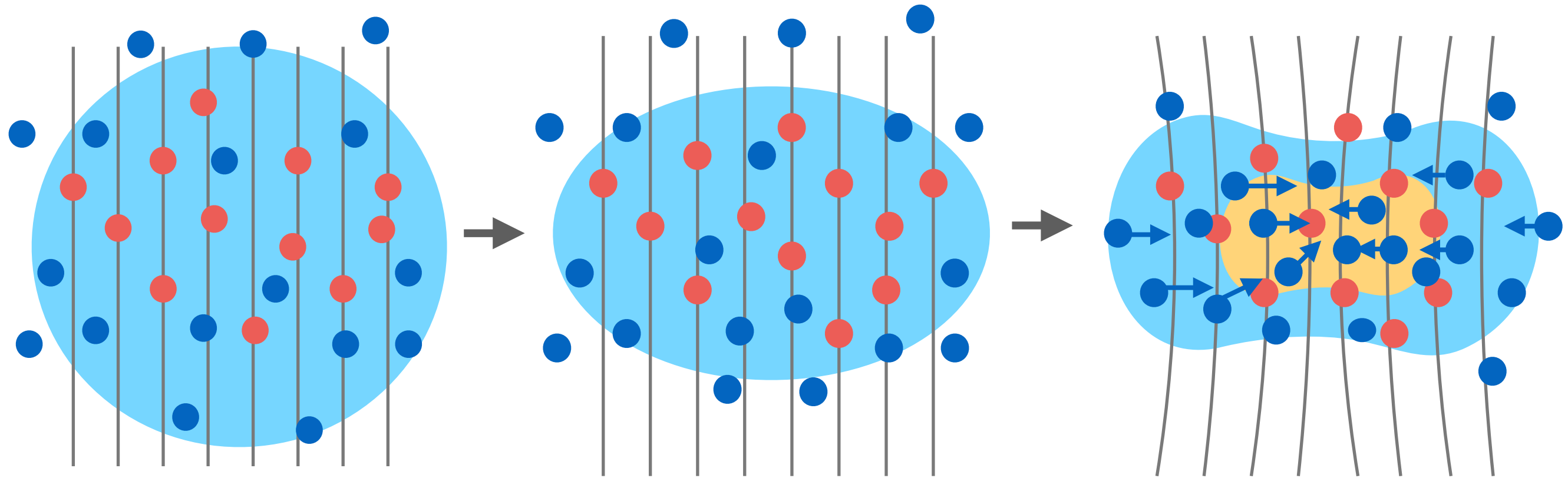
- Star formation scenario



# Strong field model - Ambipolar diffusion



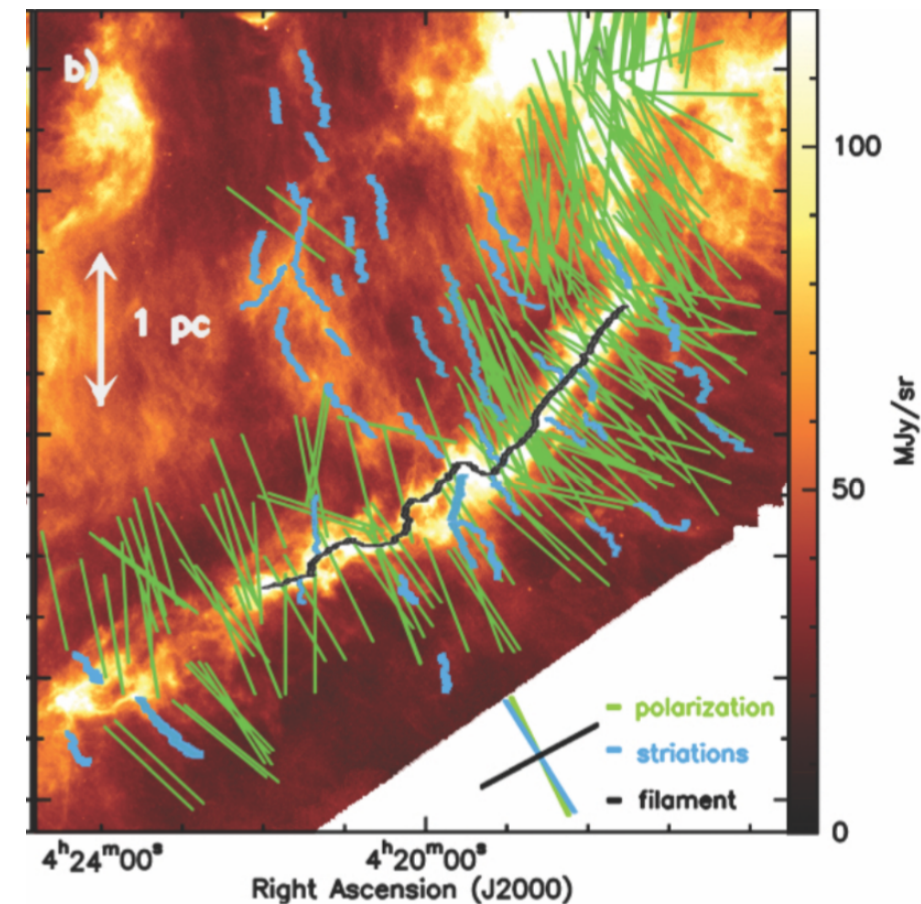
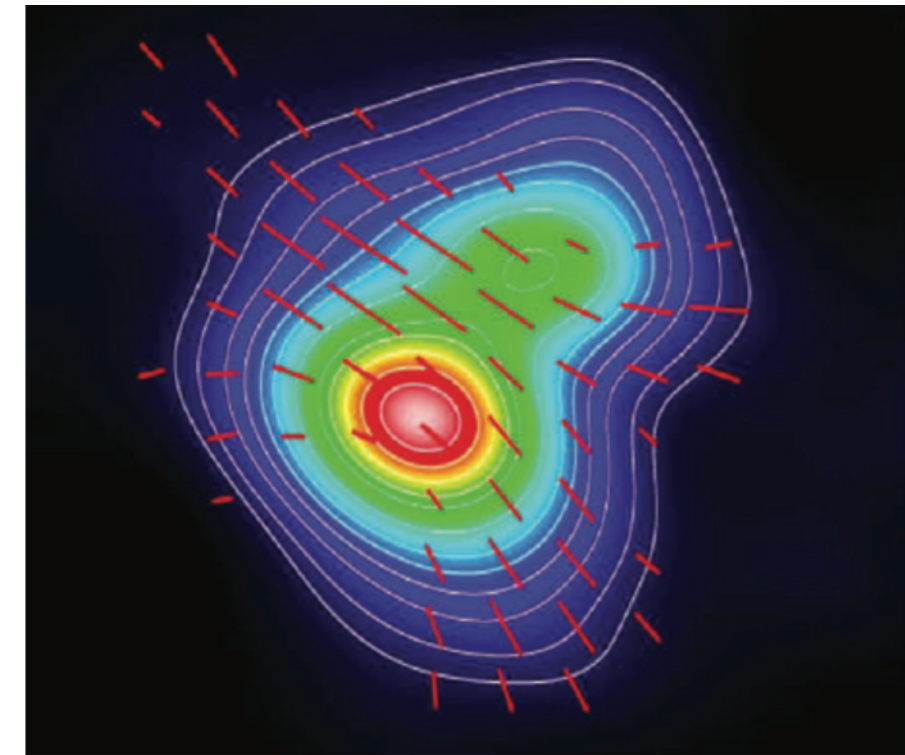
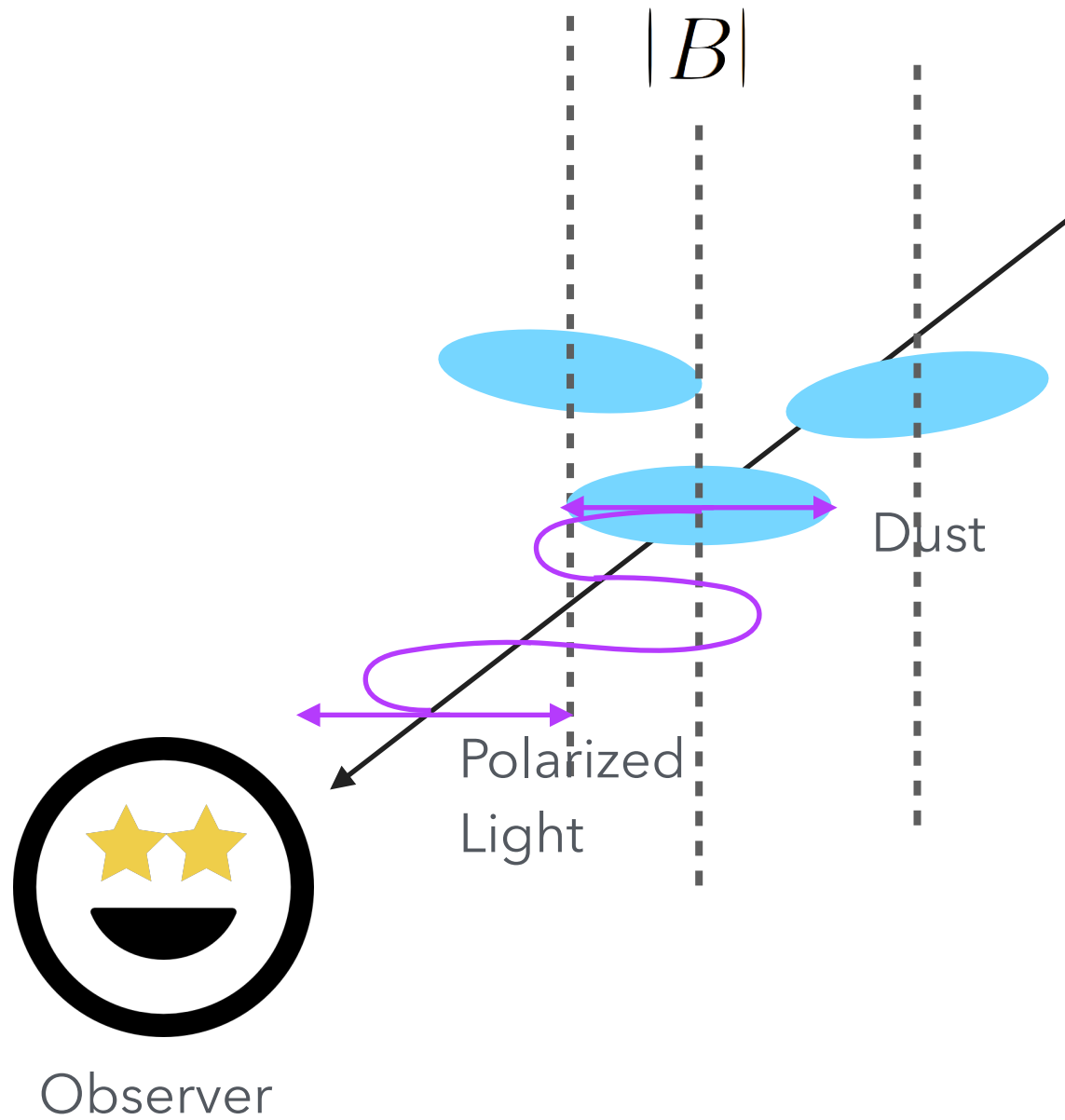
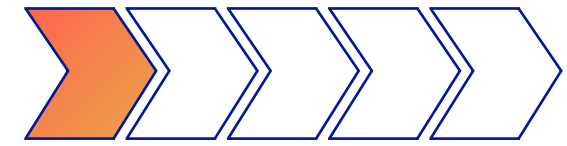
Mouschovias 1991, Mouschovias & Ciolek 1999



● Charged particle (ion, electron..)      ● Neutral particle

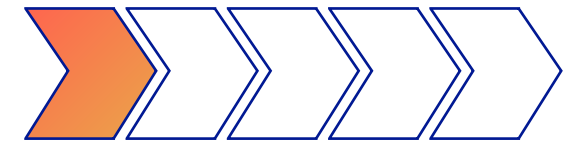


# Polarization



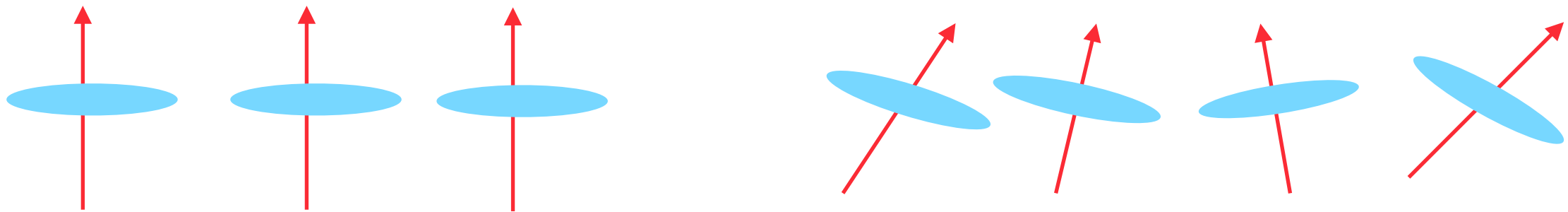
Girart, Rao & Marrone 2006  
P. Palmeirim et al 2013

## ● Chandrasekhar-Fermi (CF) method

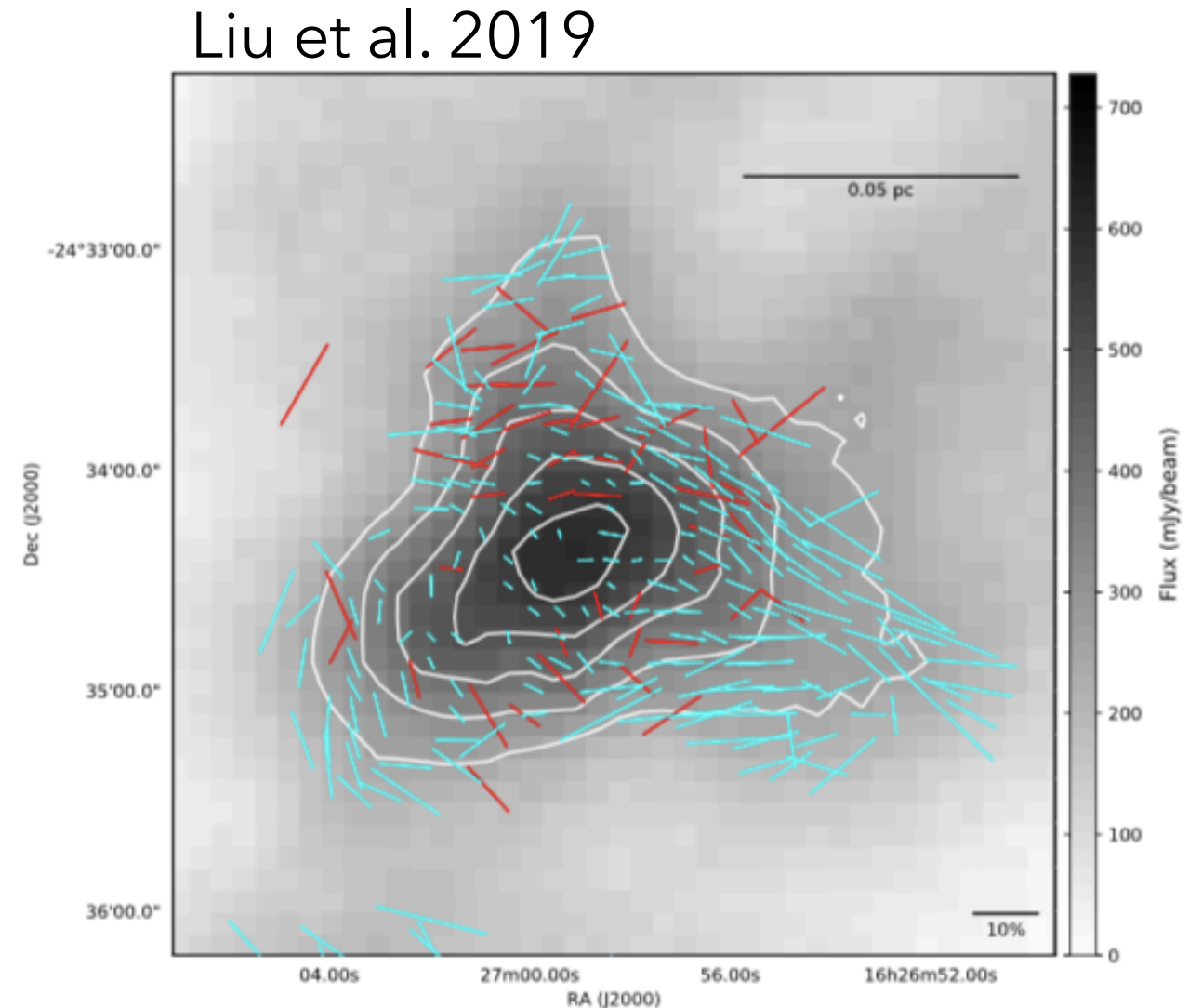
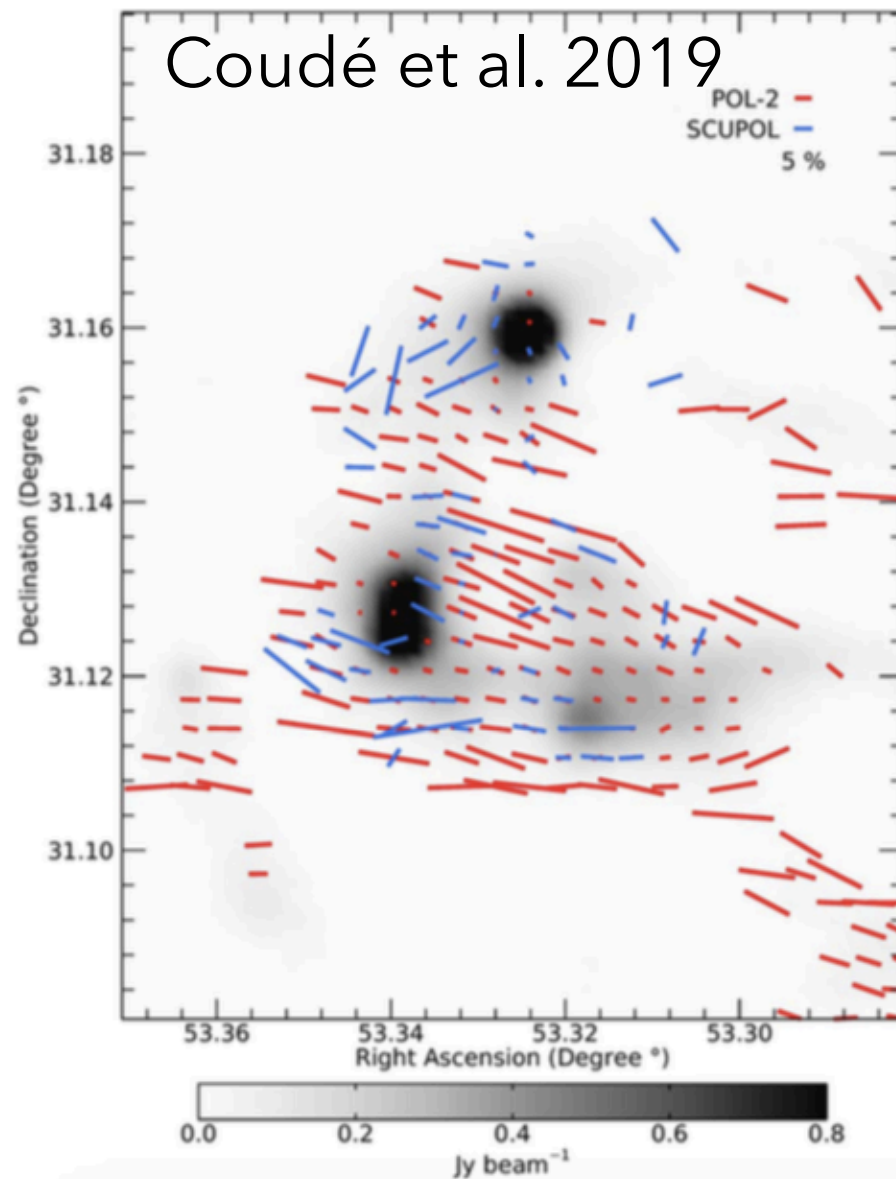
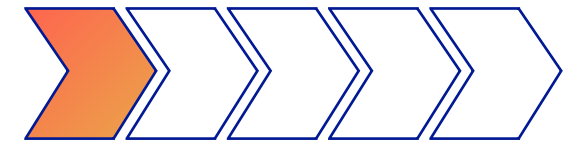


The dispersion of measured polarization angles is proportion to the distortion of magnetic field lines by turbulent motions of gas.

$$B_{pos} = 9.3 \frac{\sqrt{n_{H_2}} \Delta V}{\delta \phi}$$



# Chandrasekhar-Fermi (CF) method

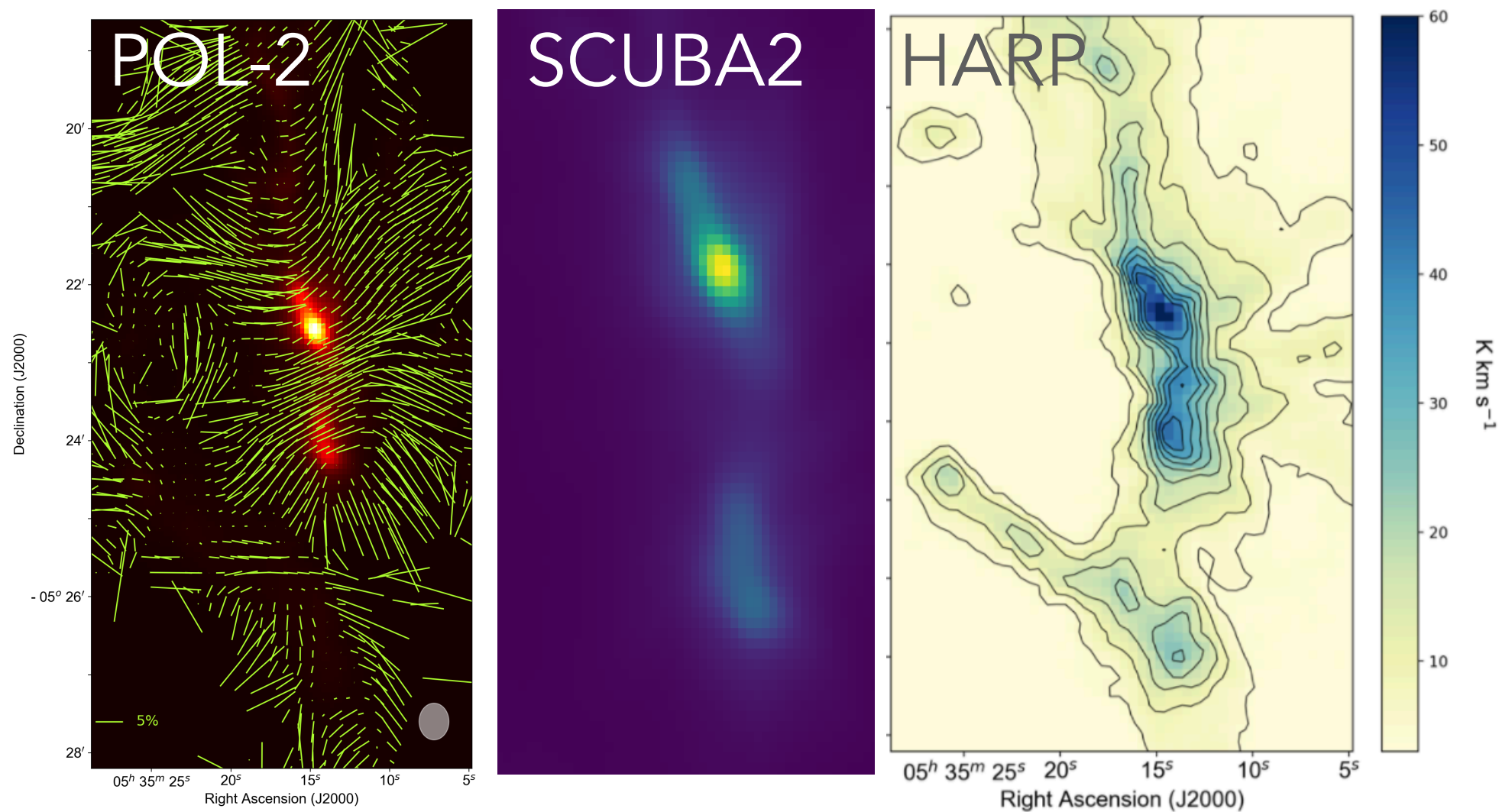


Previous studies obtained magnetic field strengths over the scales of molecular clouds or cores.

We will estimate the distribution of magnetic field strength within a molecular cloud, OMC 1.

# OBSERVATIONS

- JCMT (James Clerk Maxwell Telescope)



BISTRO (B-field In STar-forming RegiOns), Gould Belt Survey , POL2 commissioning data

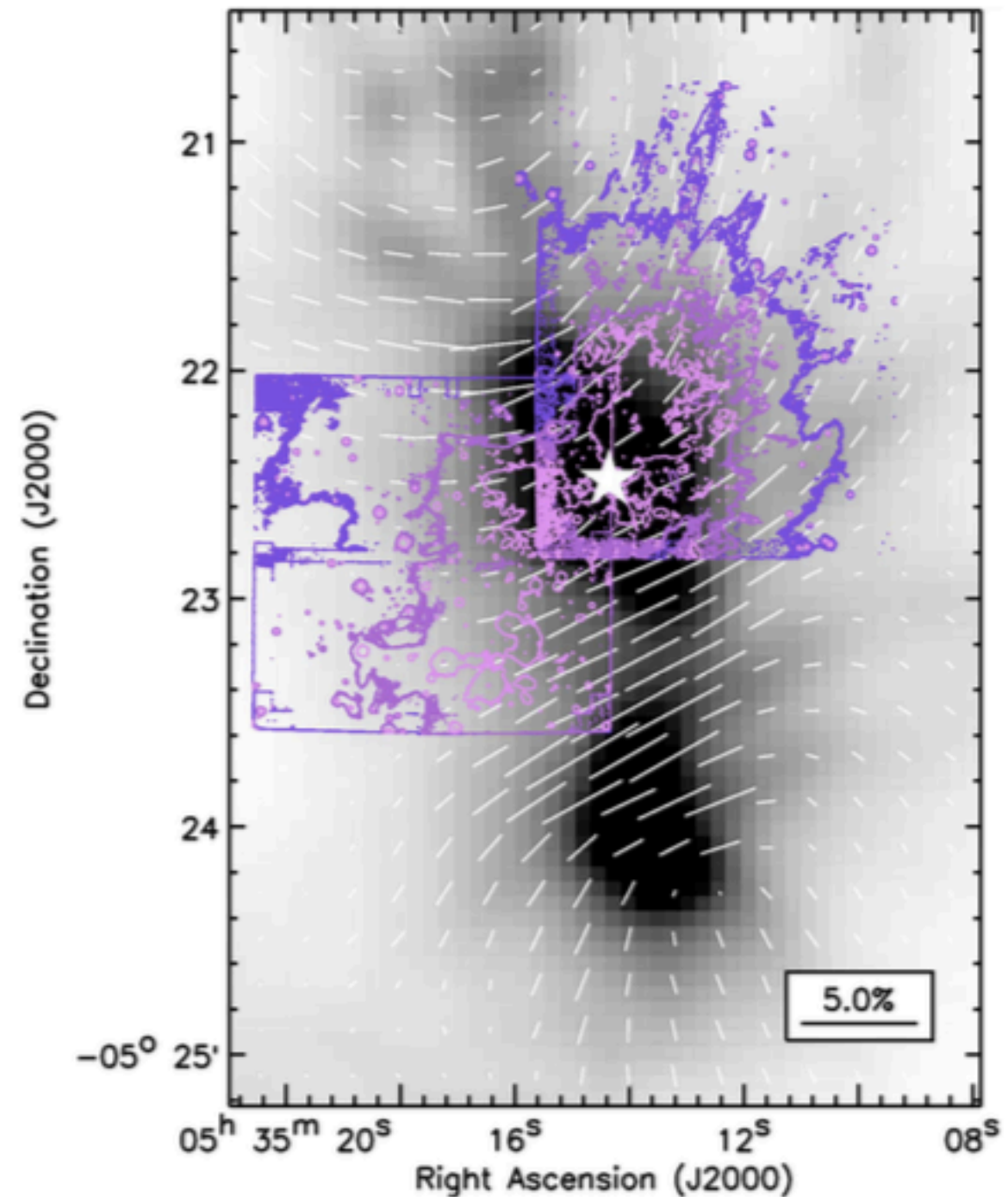
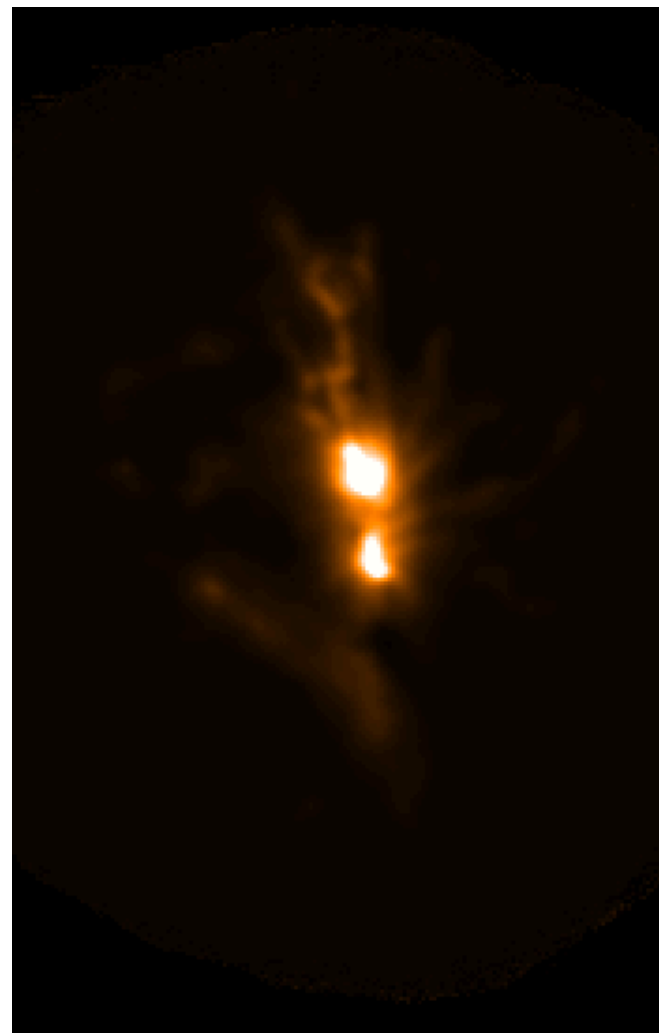


# Orion A region - OMC 1



$388 \pm 5$  pc (Kounkel et al. 2017)  
High-mass star-forming region  
BN/KL & S clumps  
Explosive outflow

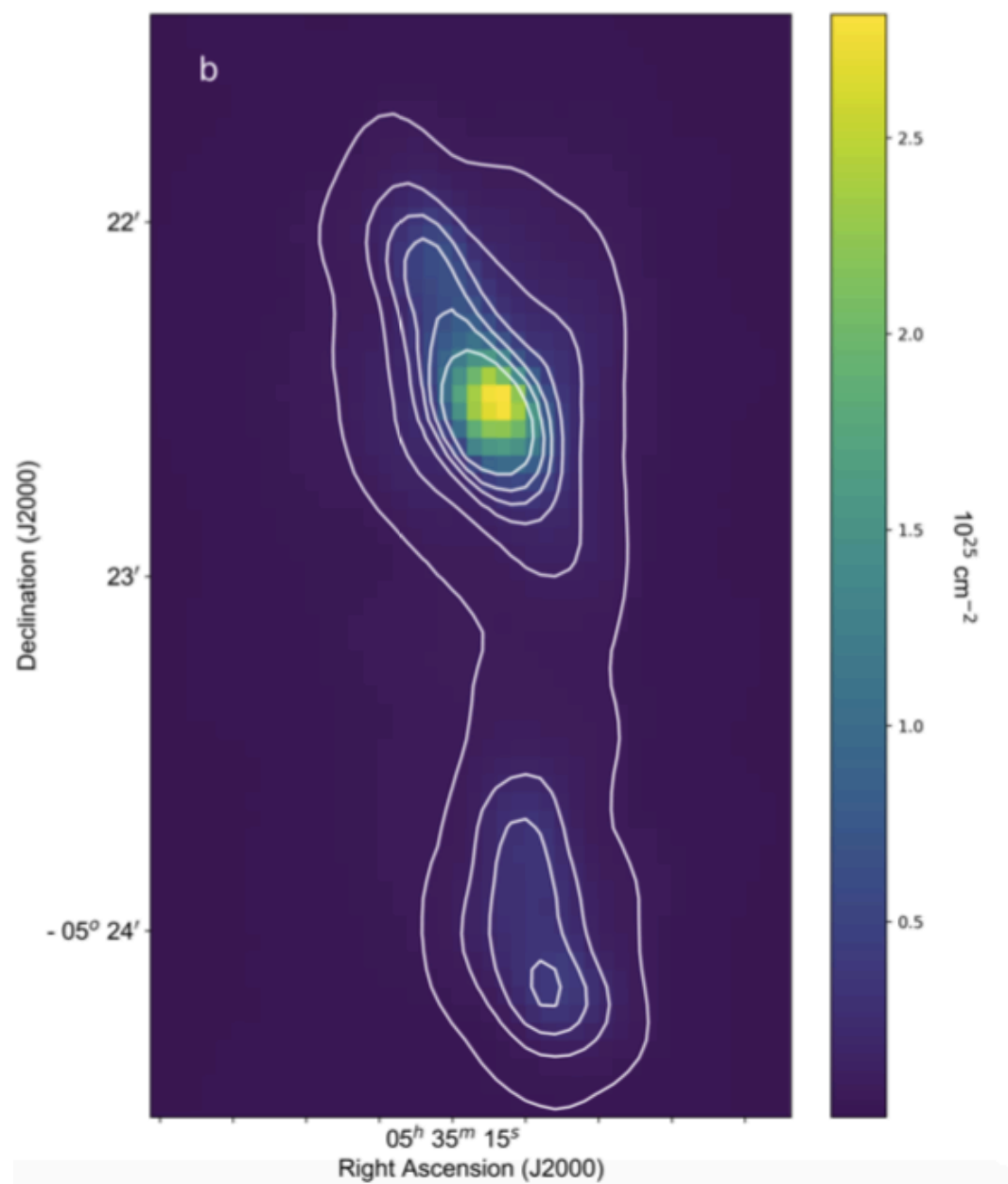
JCMT 850  $\mu\text{m}$   
dust continuum



Pattle et al. 2017

# RESULTS

## ● Column density map



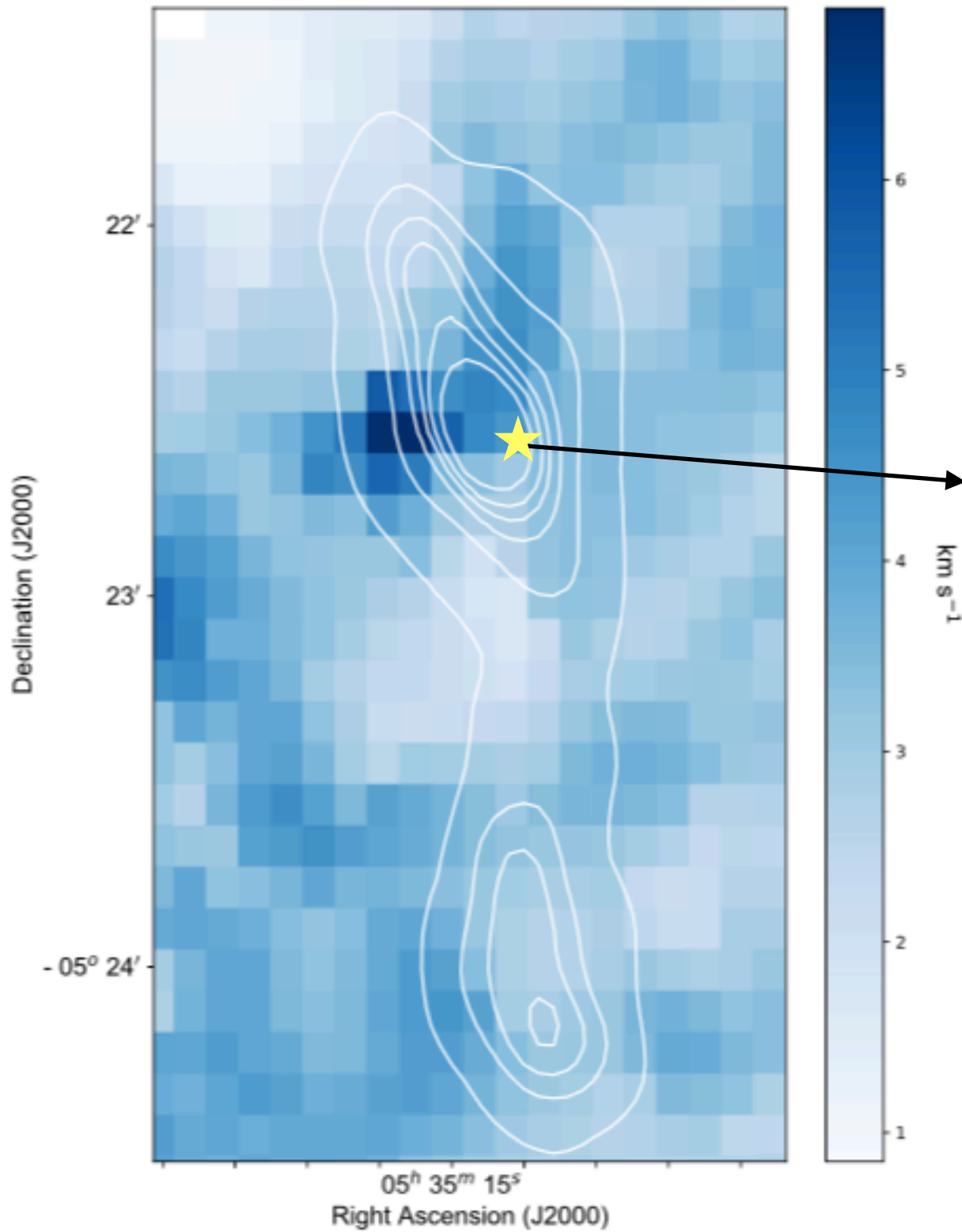
$$\frac{I_{850}}{I_{450}} = \left( \frac{\nu_{850}}{\nu_{450}} \right)^{3+\beta} \times \frac{e^{\frac{h\nu_{450}}{k_B T}} - 1}{e^{\frac{h\nu_{850}}{k_B T}} - 1} \quad \beta = 2$$

$$I_{\nu} = \mu m_H \kappa(\nu) N(\text{H}_2) B_{\nu}(T)$$

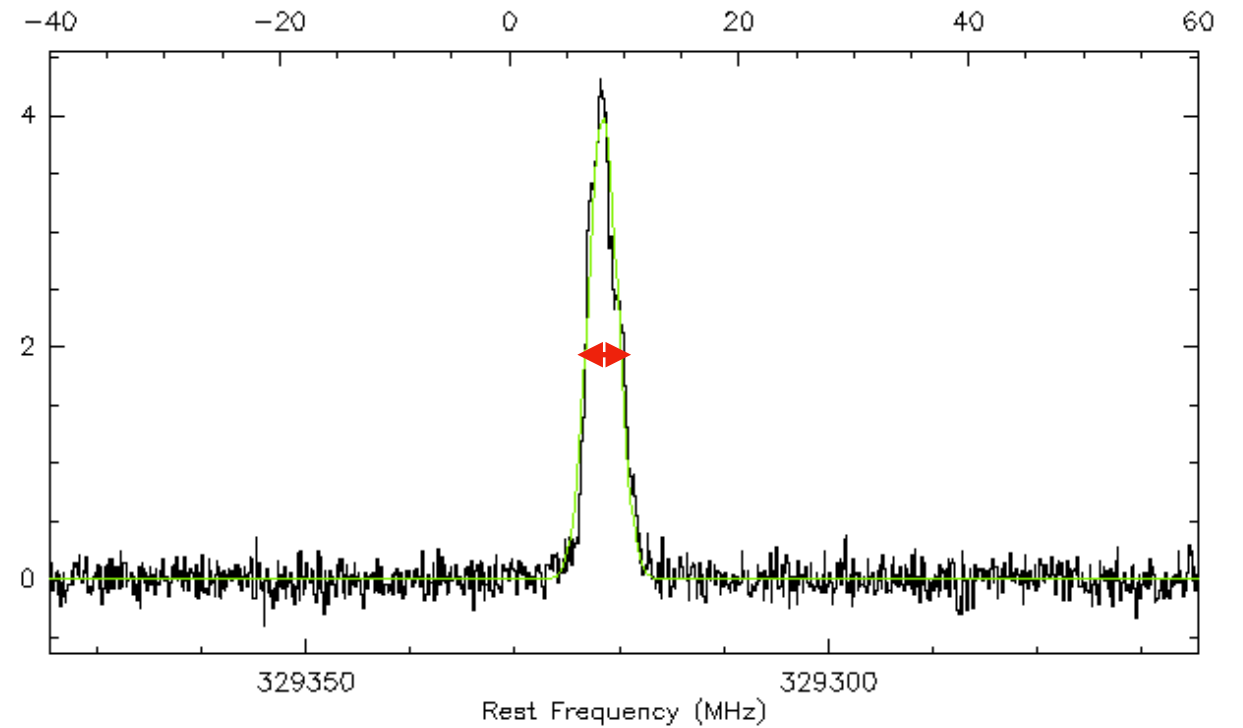
$$\kappa(\nu) = \kappa_{\nu_0} \left( \frac{\nu}{\nu_0} \right)^{\beta}$$

Salji et al. 2015; Pattle et al. 2017

# Velocity dispersion map

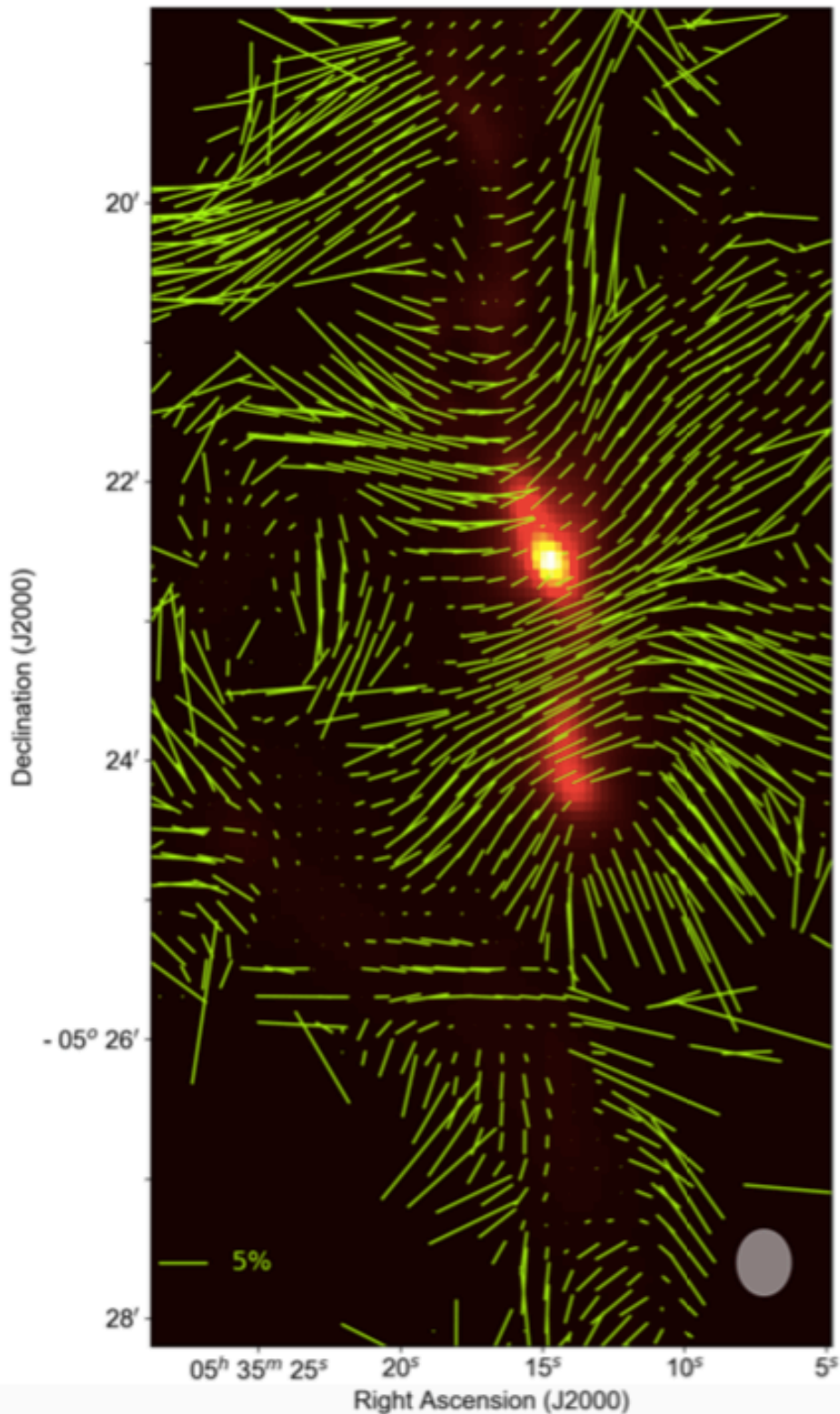


HARP



Measure the Full Width at Half Maximum of  $C^{18}O$

# Polarization angle dispersion

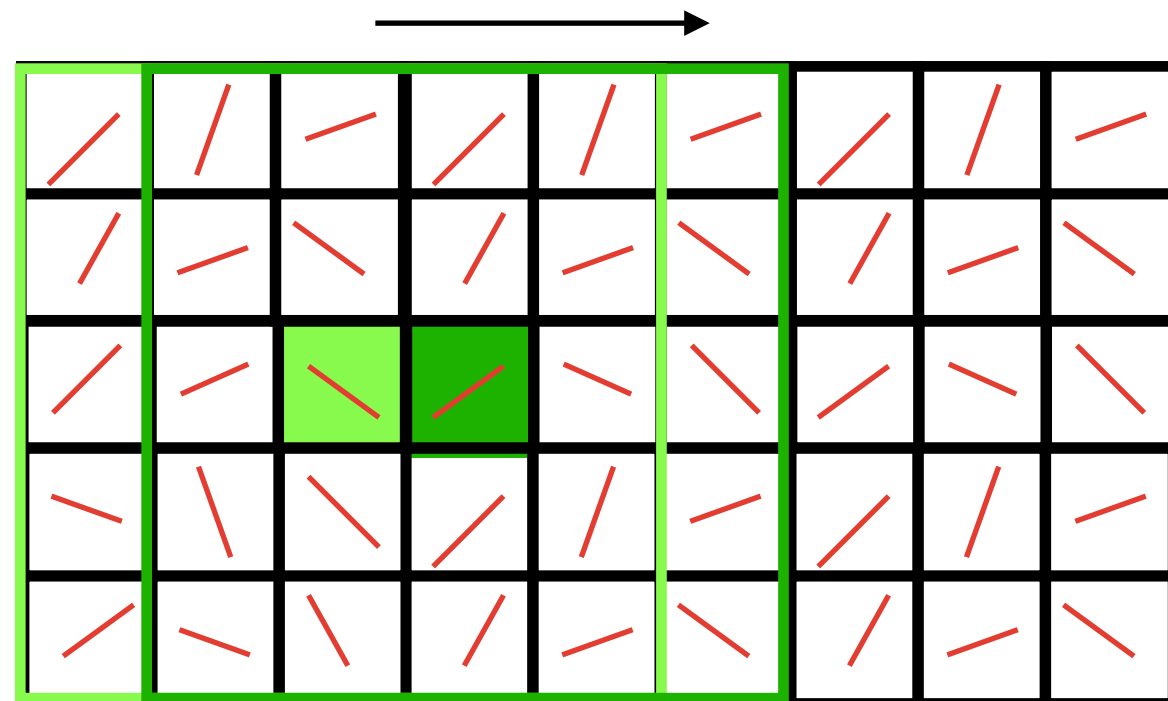


1 pixel = 4 arcsec.

Make moving boxes  
(4x4, 5x5, 6x6, 7x7)

Determine angle dispersion at the  
center of a box

$$AD = \sqrt{\frac{\sum(\theta - \bar{\theta})^2}{N}}$$





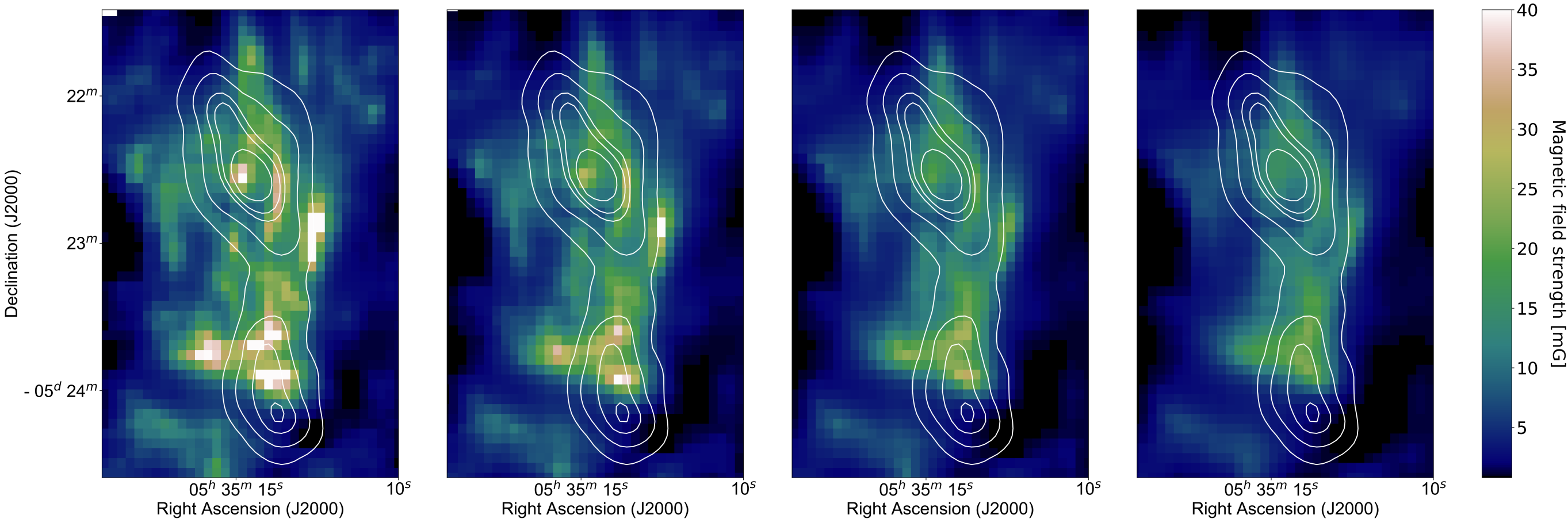
# DISCUSSIONS

4X4

5X5

6X6

7X7



$14.8 \pm 6.7$

$12.1 \pm 5.6$

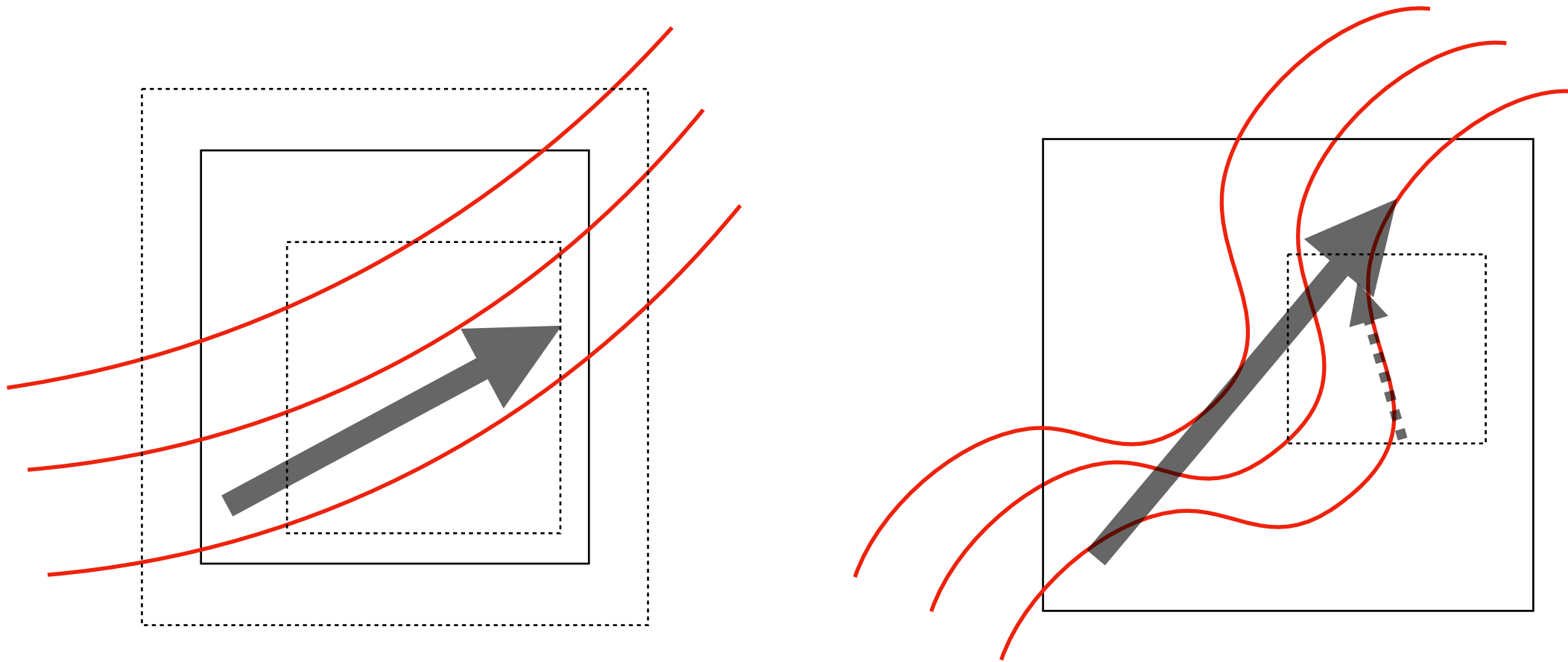
$10.4 \pm 5.0$

$9.3 \pm 4.7$  [mG]

## ● How to measure curvature

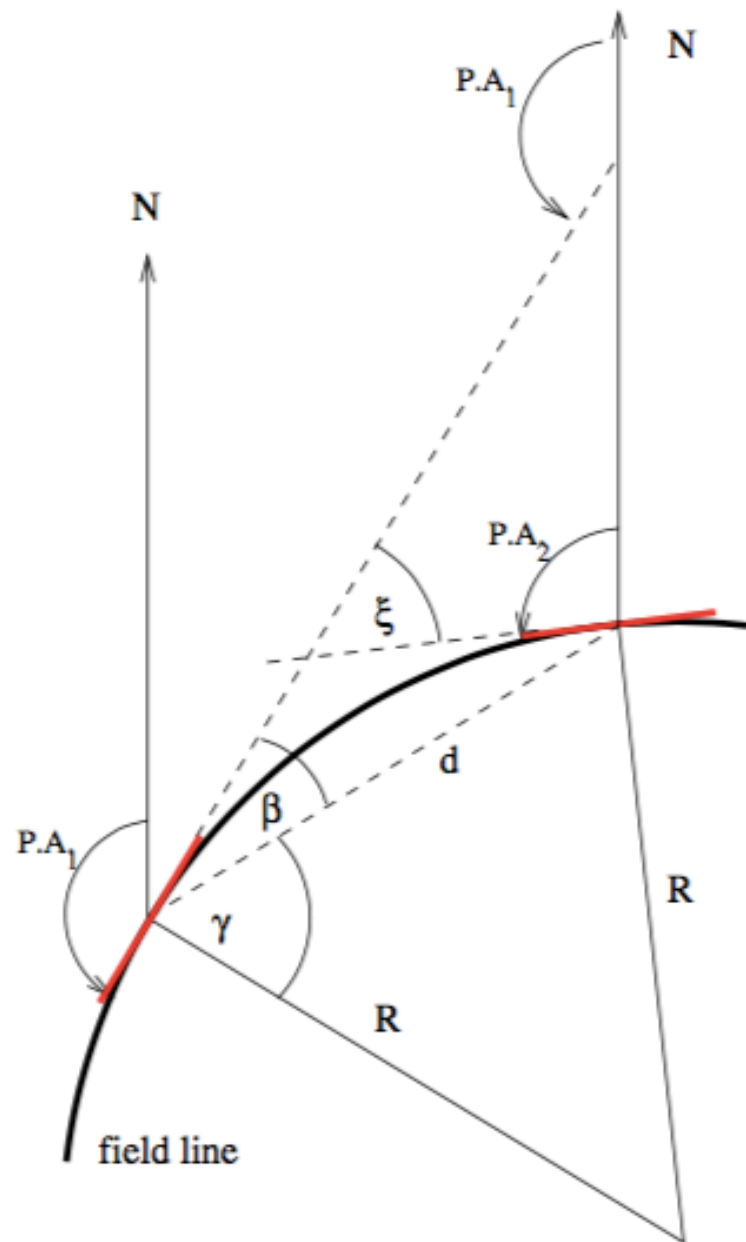


Assumption : A mean field direction in a box is uniform.



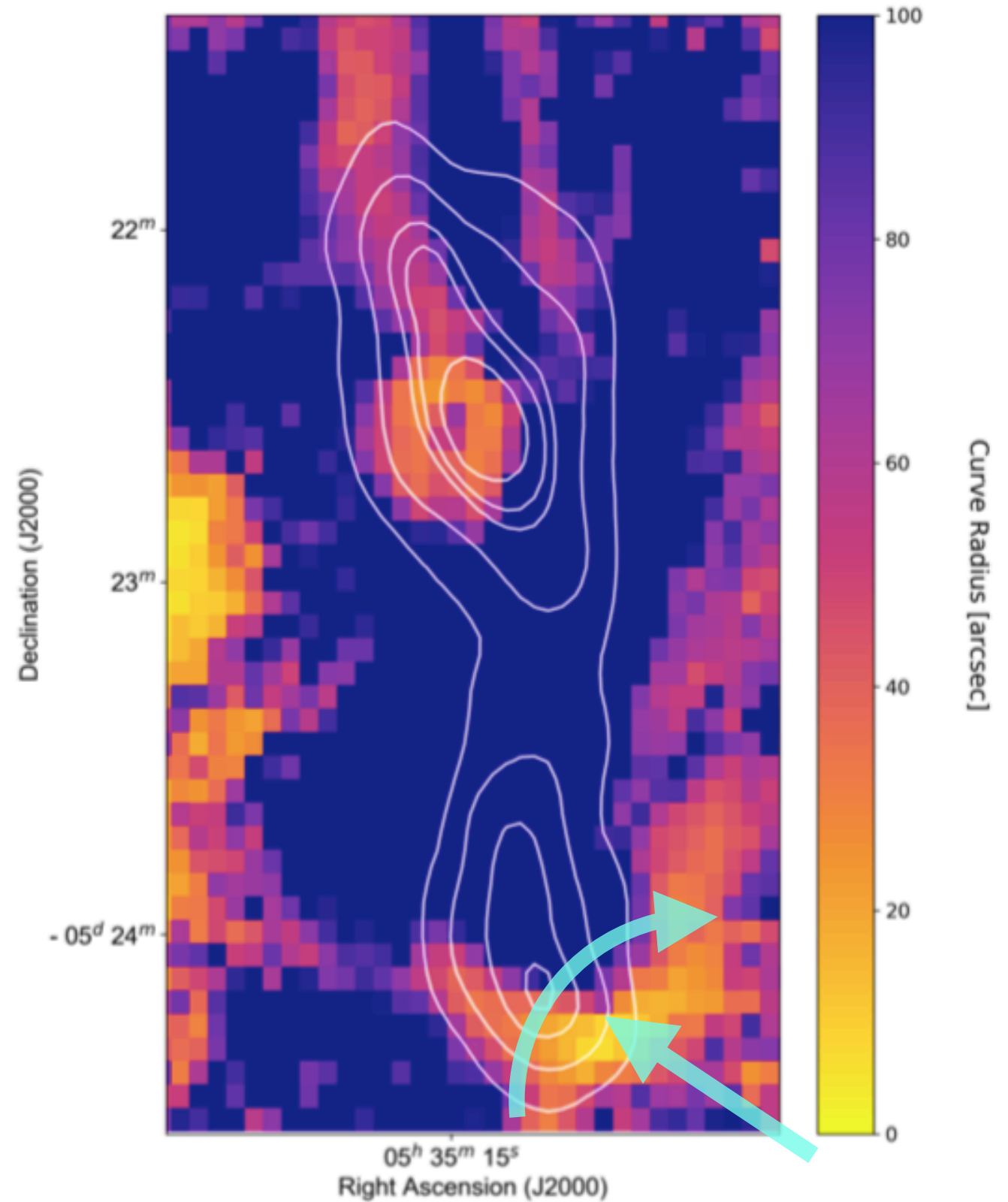
If we use a moving box whose size is larger than curvatures of magnetic field lines, then an angle dispersion of polarization segments (a magnetic field strength) within the box will be overestimated (underestimated).

# ● Curvature Radius



$$C = \frac{1}{R} = \frac{2}{d} \cos\left(\frac{1}{2}[\pi - \Delta P.A.]\right)$$

Koch et al. 2012

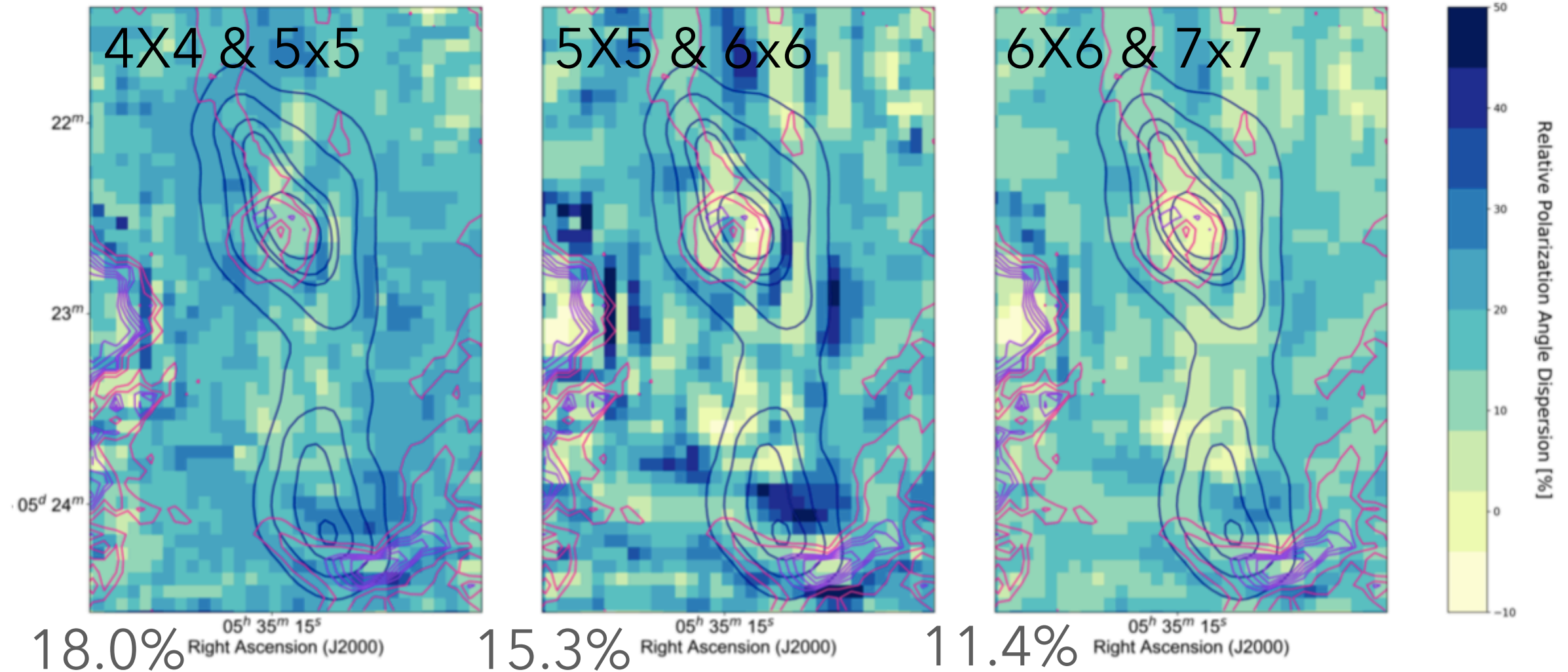


# Relative polarization angle dispersion



$$\Delta AD = \frac{AD_L - AD_S}{AD_L} \times 100[\%]$$

$AD_L$  angle dispersion in a large box  
 $AD_S$  angle dispersion in a small box



The relative angle dispersion between 6x6 and 7x7 boxes is the smallest and uniform. The central part of BN/KL shows smaller curvature than 7x7 box, so we consider 6x6 box is appropriate to estimate the distribution of magnetic field strength in OMC 1.

# SUMMARY



- 1) We developed a way to find a distribution of magnetic field strength using the CF method.
- 2) We obtained a map of magnetic field distribution with a pixel size of 4" in OMC 1 region.  
The maximum, minimum and mean field strengths obtained by 6x6 box are 0.5, 30.3, and, 10.4, respectively.
- 3) The size of a moving box should be determined in comparison with curvature of field lines.

**Thank You**