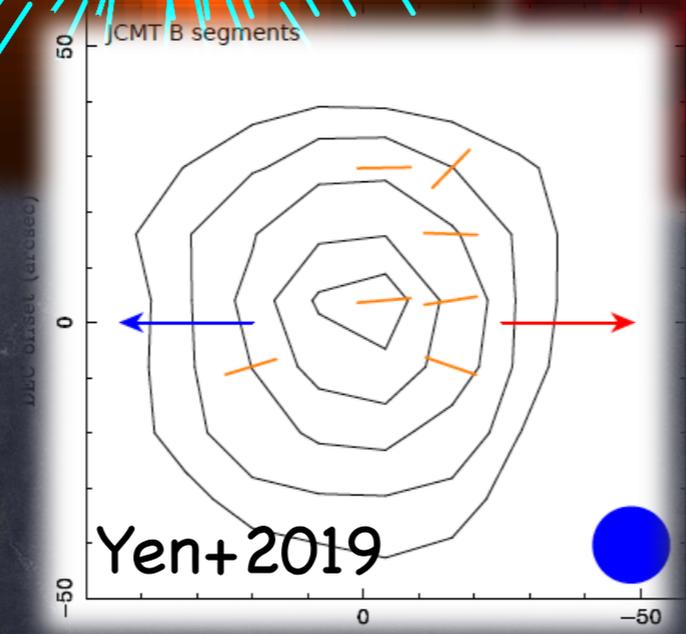
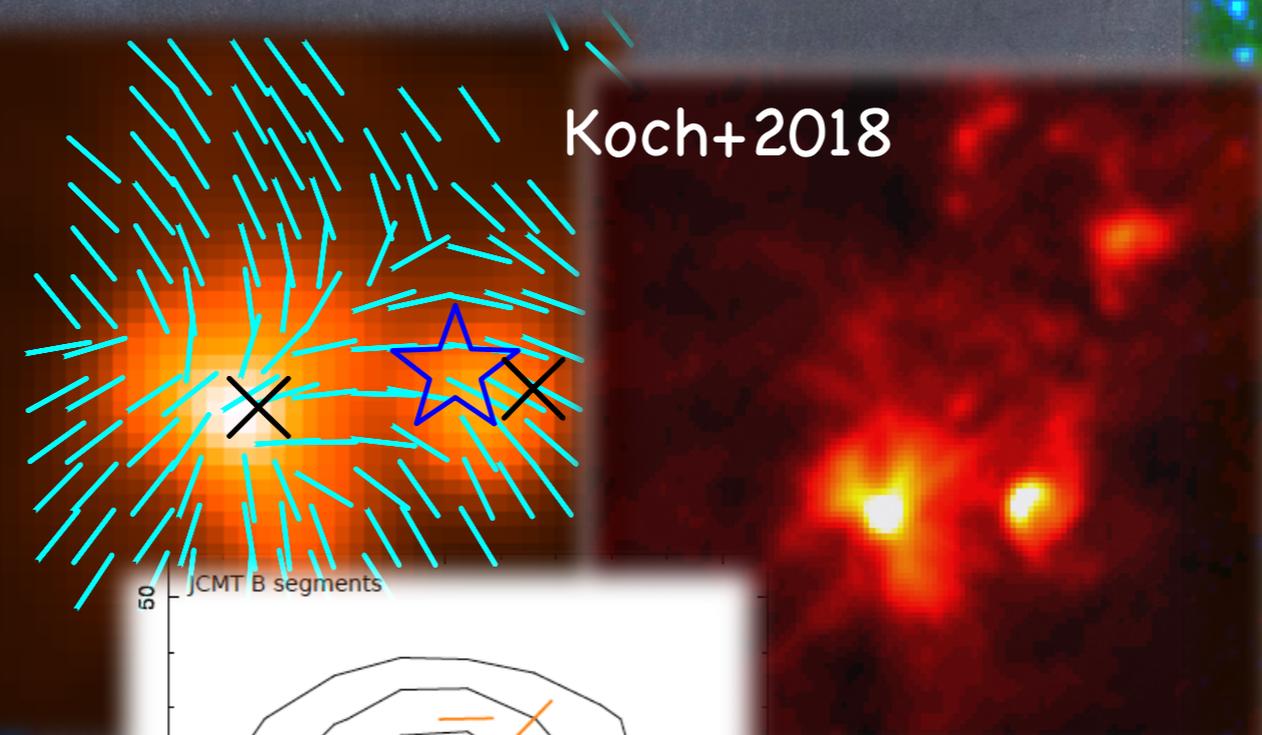
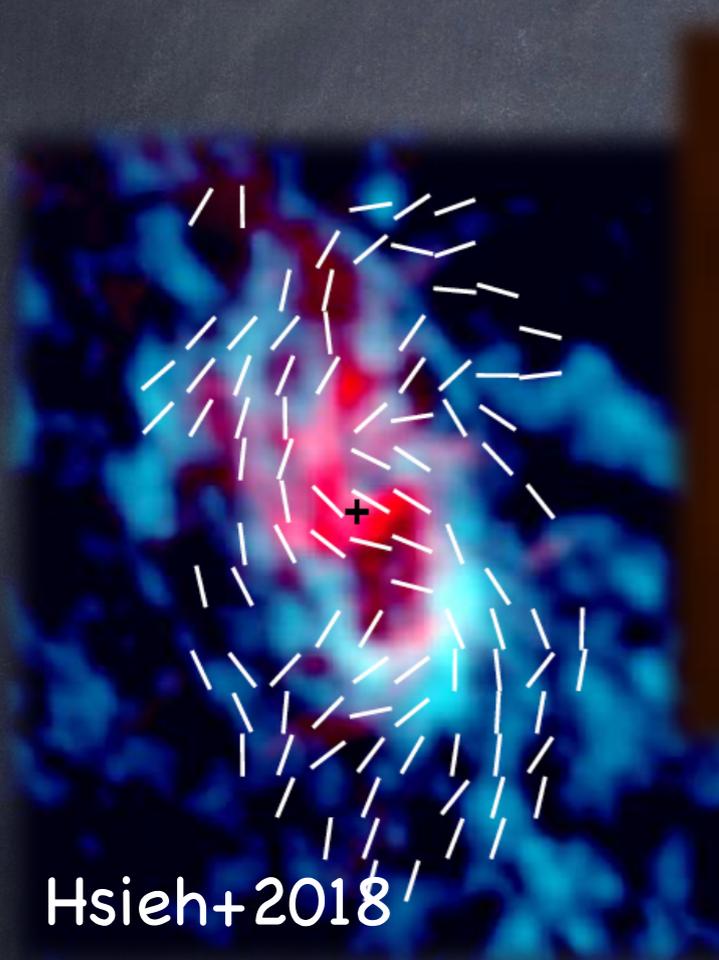


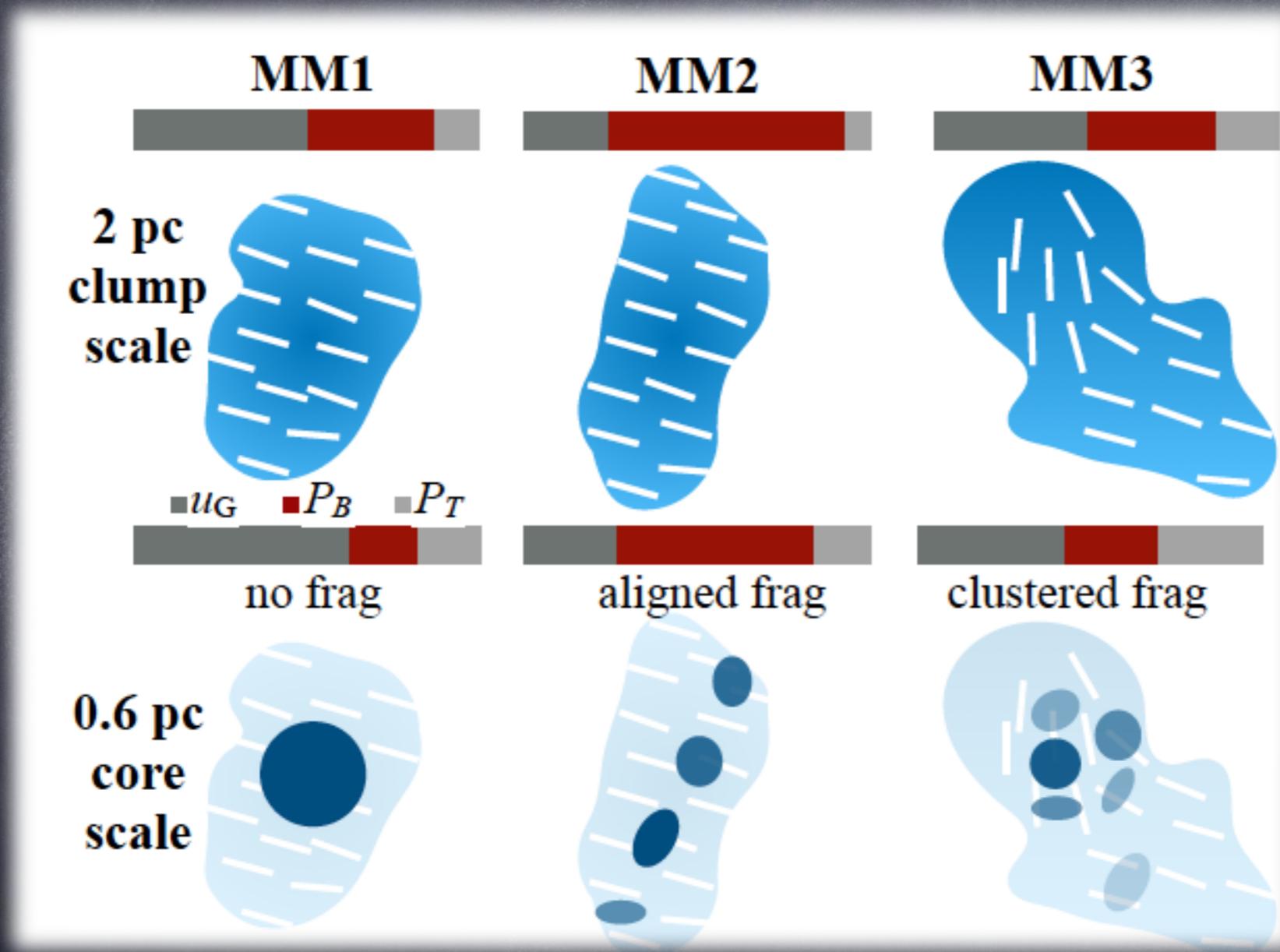
# Fragmentation and Relative Importance between Gravity, Magnetic Field, and Turbulence in IRDC G34

Patrick Koch (ASIAA)

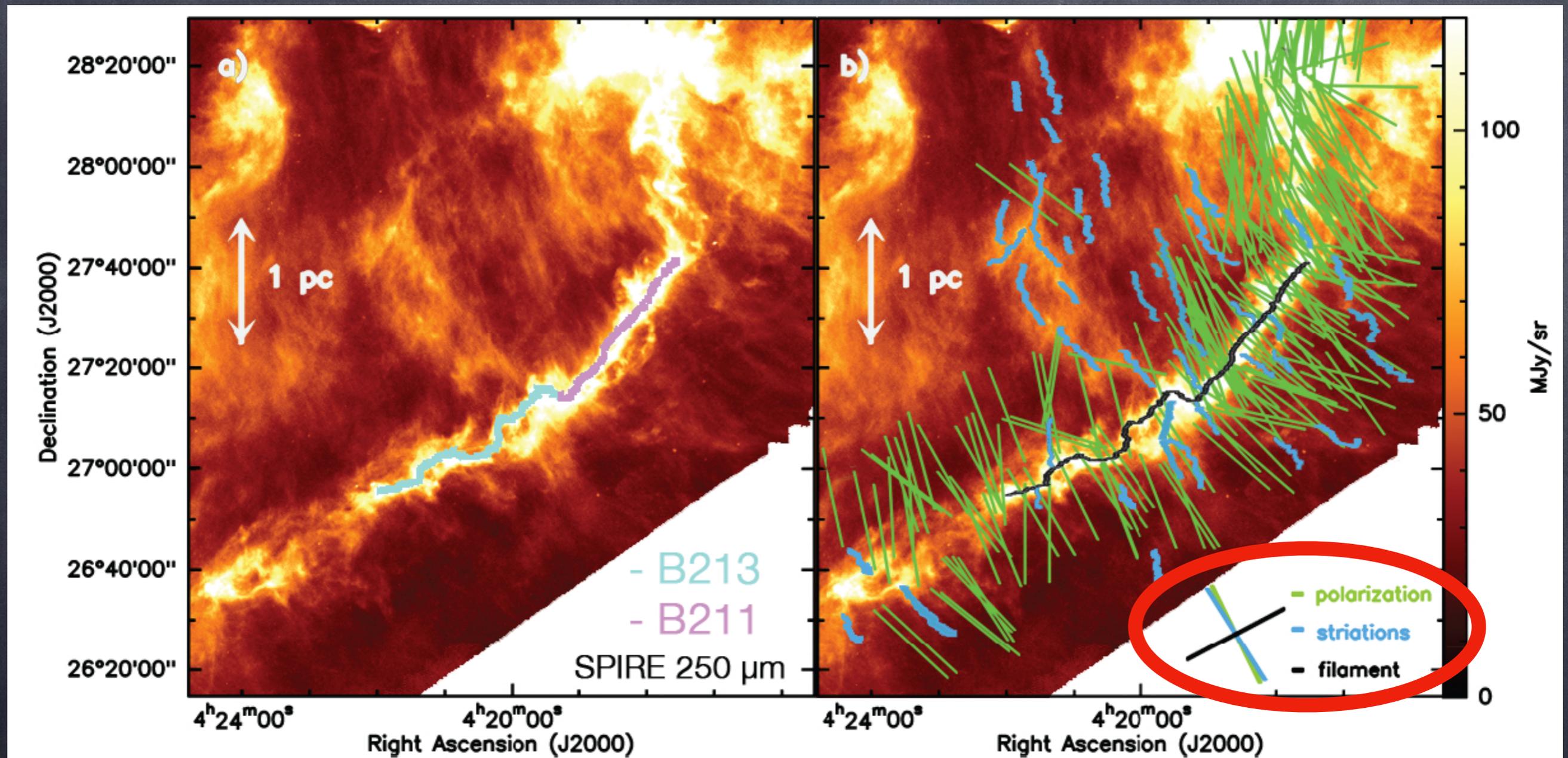
and Y.-W. Tang, N. Peretto,  
G. Novak, A. Duarte-Cabral,  
N. Chapman, P.-Y. Hsieh, H.-W. Yen



# Fragmentation and Relative Importance between Gravity, Magnetic Field, and Turbulence in IRDC G34

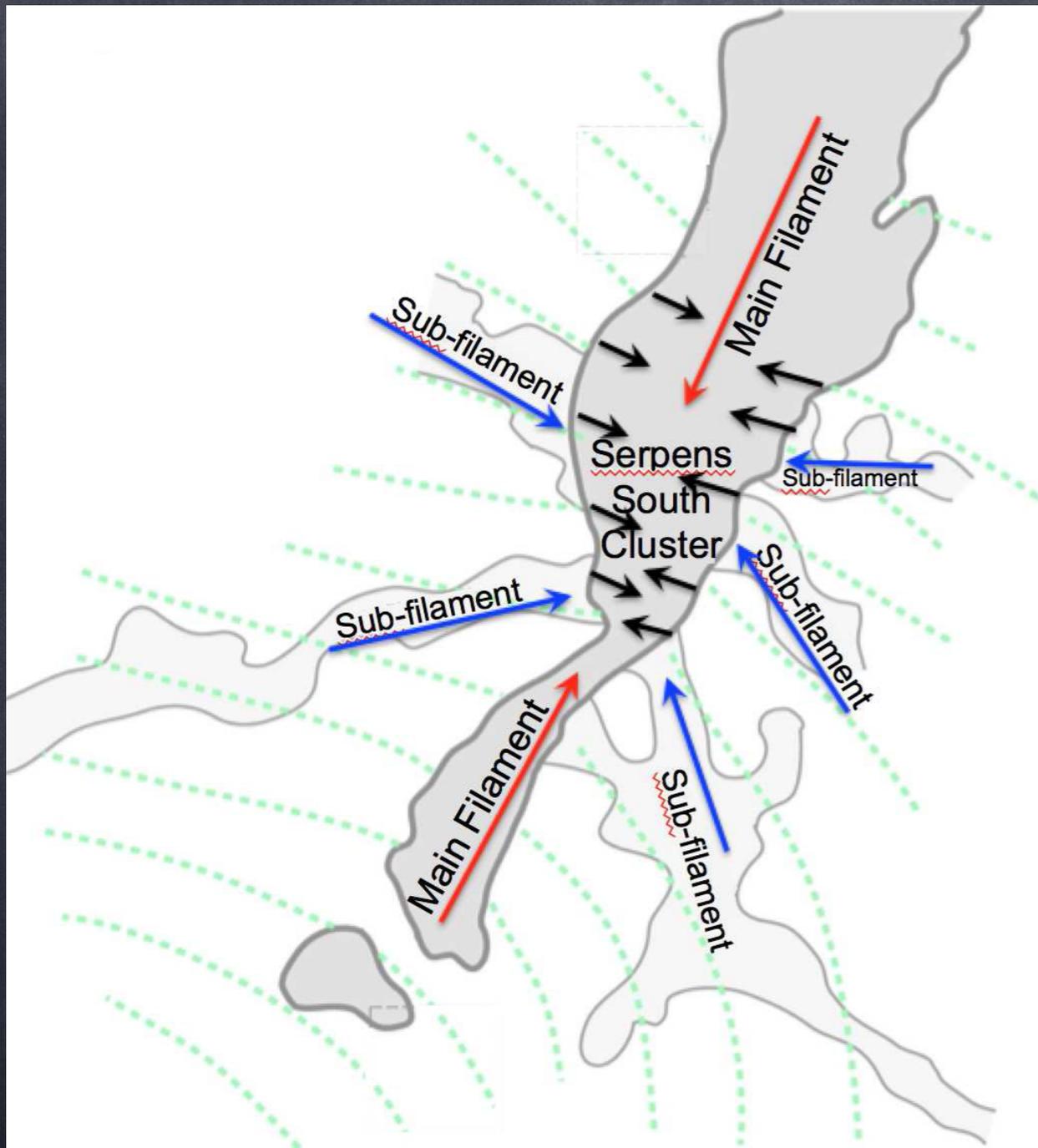


# Filaments and Magnetic Fields



- *Taurus molecular cloud; Herschel / SPIRE 250 μm (~ 18" resolution; Palmeirim+2013)*
- *polarization: optical (Heyer+2008; Heiles+2000) and infrared (Chapman+2013)*
- *B-field  $\parallel$  striations  $\perp$  filament*

# Filaments and Magnetic Fields: Proposed Scenario

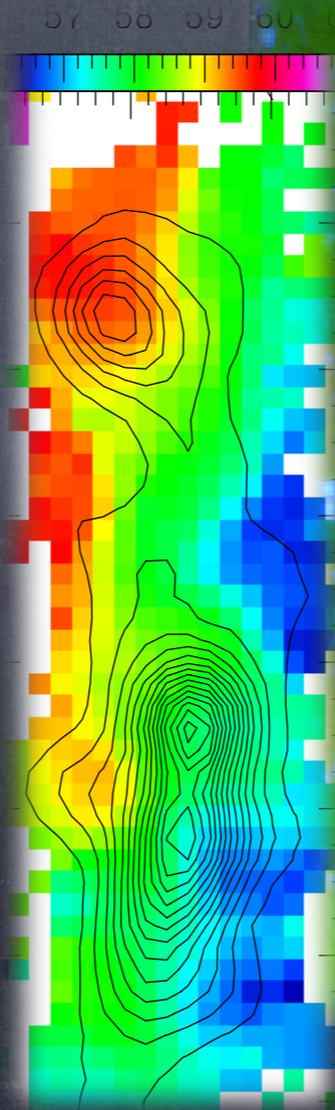


- accretion of background material through striations or sub-filaments along B-field lines
- radial contraction within main filament
- longitudinal infall along main filament
- important ancillary line data: coherent velocity structures
- formation steps: (1) large-scale MHD flows (turbulent?) lead to filamentary network with universal filament width  $\sim 0.1$  pc; (2) **densest filaments fragment into prestellar cores** by gravitational instability (critical line mass  $\sim 16 M_{\text{solar}}/\text{pc}$ , critical density  $\sim 2 \times 10^4 \text{ cm}^{-3}$ )

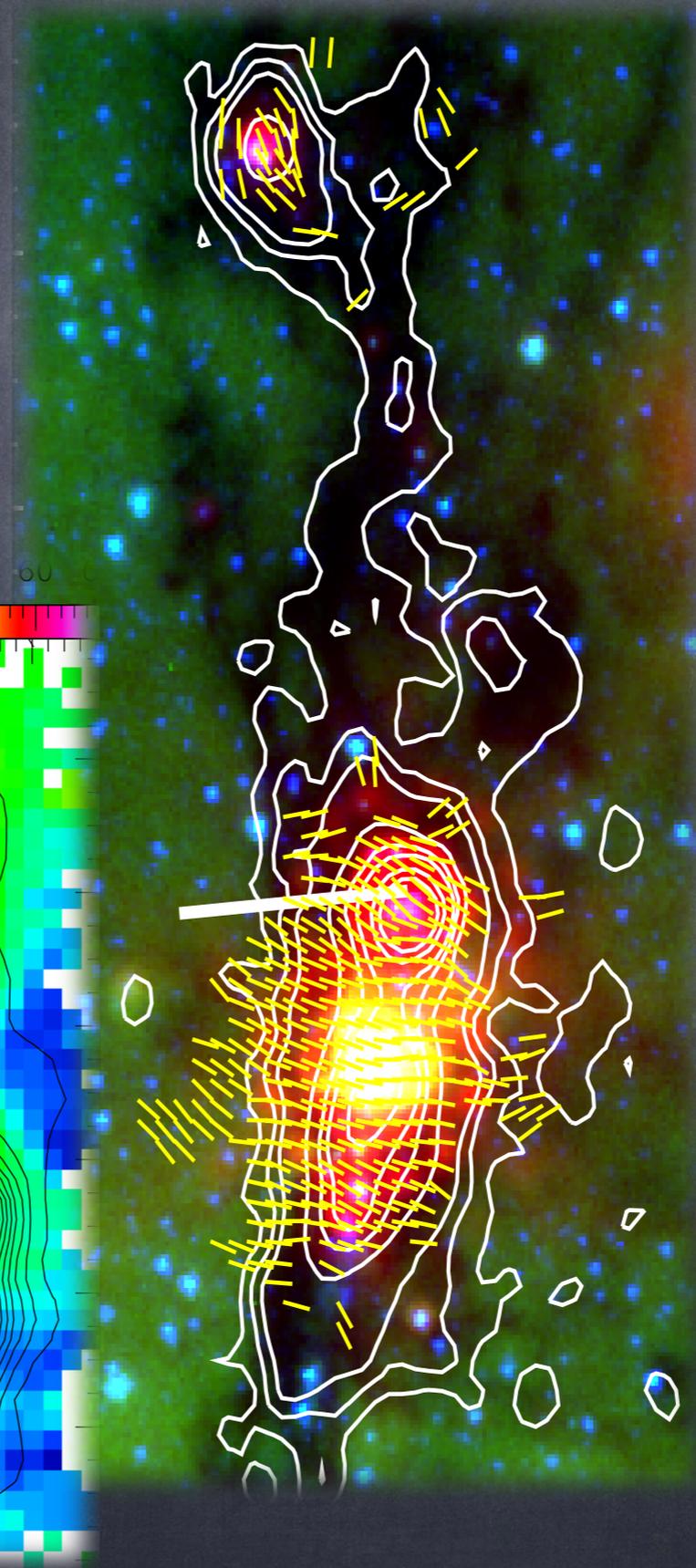
(e.g. André+2010; Mouschikov+2010; Molinari+2010; Sugitani+2011; Kirk+2013; Arzoumanian+2013; André+2014; )

# IRDC G34.43

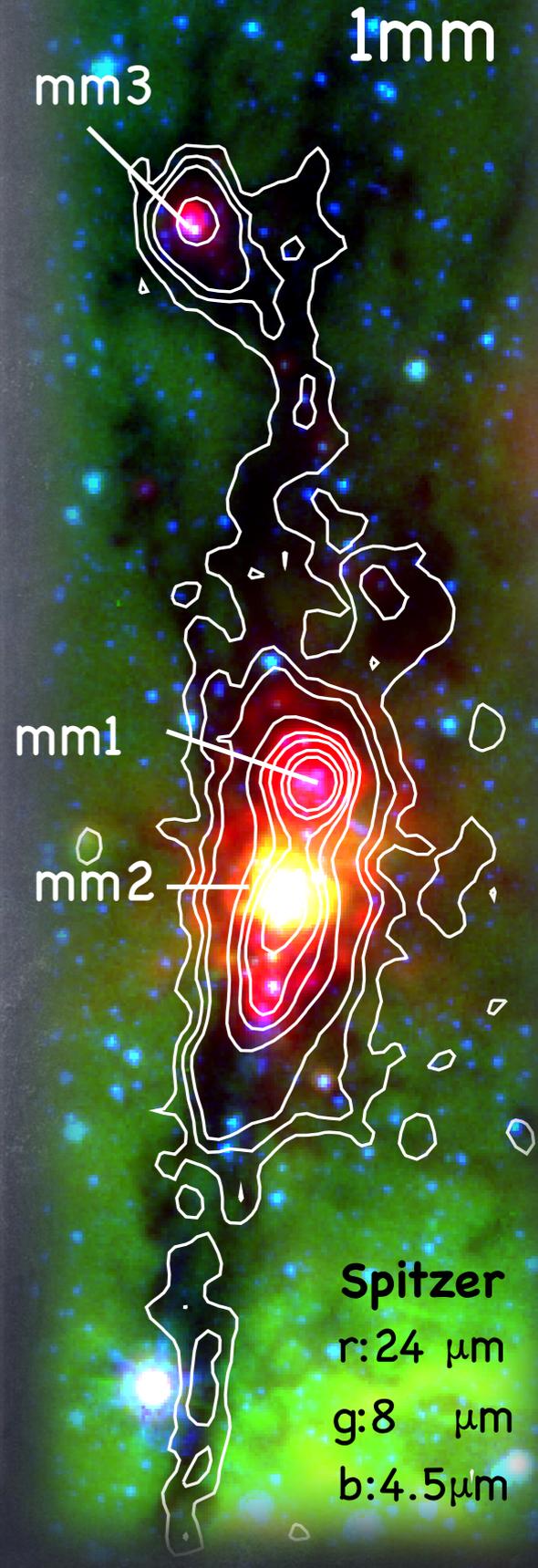
- distance: 3.7 kpc, elongated length  $\sim 8$  pc
- mass: 1200  $M_{\text{sol}}$  (mm1), 1300  $M_{\text{sol}}$  (mm2)  
300  $M_{\text{sol}}$  (mm3)
- overall, very small virial parameter ( $\alpha \sim 0.2$ ), system gravitationally bound, but SF efficiency only  $\sim 7\%$ .  
additional support from B-field ?
- observed with the CSO/SHARP (350  $\mu\text{m}$ , resolution 10")
- polarization percentage 0.4 - 10%
- B-field clearly organized perpendicular to longer axis around mm1/mm2; more aligned with longer axis on mm3, small dispersion
- add line kinematics:  
 $\text{N}_2\text{H}^+$  (1-0) from IRAM-30m ( $\theta \sim 28''$ ), clear large-scale gradient



Peretto+2017



Tang, Koch+2019

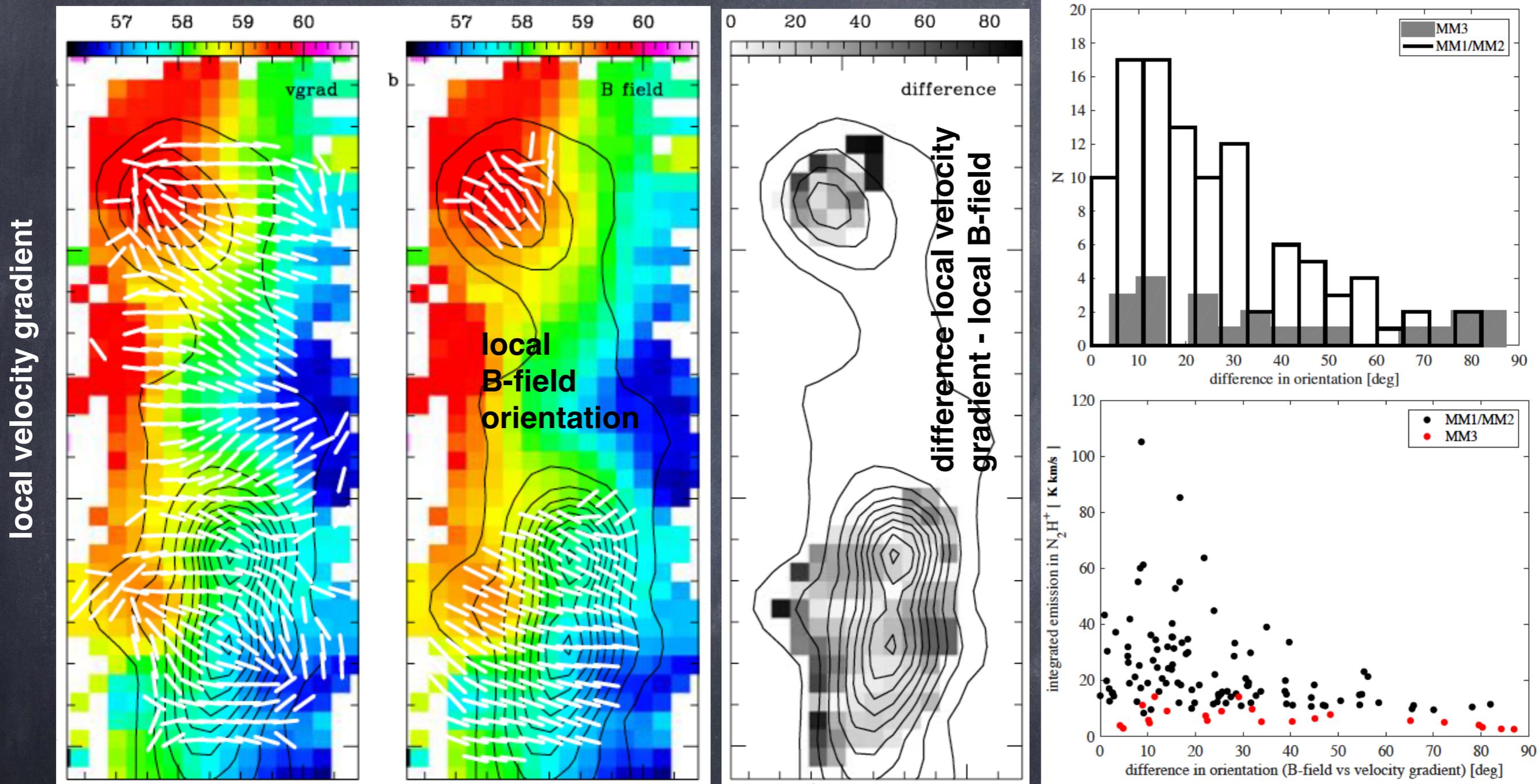


Rathborne+2006

Spitzer  
r: 24  $\mu\text{m}$   
g: 8  $\mu\text{m}$   
b: 4.5  $\mu\text{m}$

# B-field, velocity gradient, turbulence & gravity

*which component is dominant? negligible? - benchmark analysis*

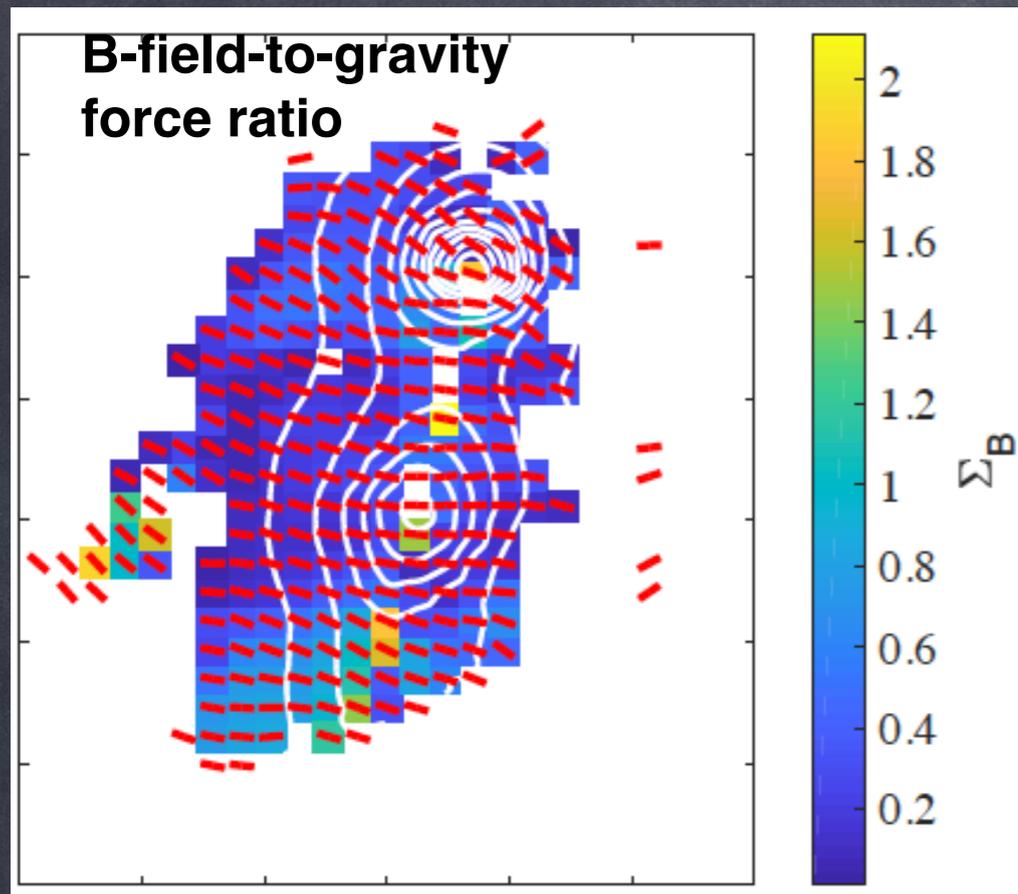


*B vs v: small differences and spatially not random, but organized*

(Tang, Koch+2019)

# B-field, velocity gradient, turbulence & gravity

*which component is dominant? negligible? - benchmark analysis*



## Polarization - Intensity gradient technique

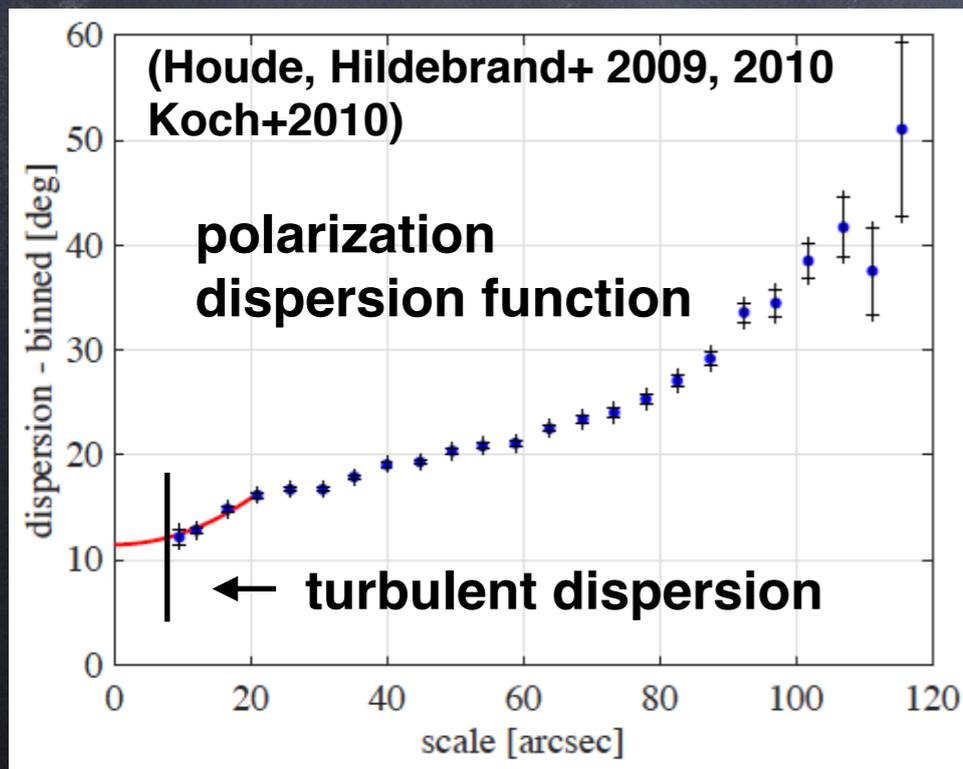
$$\Sigma_B \equiv \frac{\sin \psi}{\sin \left( \frac{\pi}{2} - |\delta| \right)} = \frac{F_B}{|F_G + F_P|}$$

$$B = \sqrt{\frac{\sin \psi}{\sin \left( \frac{\pi}{2} - |\delta| \right)} (\nabla P + \rho \nabla \phi) 4\pi R}$$

Koch+2012ab, 2013

- competition gravity vs B-field
- map of local field strength, local force ratio

*B vs G: spatially not random, but organized*



*and also DCF, virial parameters, mass-to-flux ratio, energy comparison*

# Analysis — Gravity vs Turbulence vs B-field

## 2 pc clump area

Object	$\frac{\sqrt{\langle B_t^2 \rangle}}{B_0}$	$\frac{\sqrt{\langle B_t^2 \rangle}}{B_0} N$	N	$\Delta\phi_B$ ( $^\circ$ )	$B_\perp$ (mG)	$\langle \Sigma_B \rangle$	$\lambda_{\text{obs}}$	$\alpha_{\text{vir}}$	$\alpha_{B,\text{vir}}$	$P_T$ ( $10^{-9}$ dyn/cm $^2$ )	$P_B$ ( $10^{-9}$ dyn/cm $^2$ )	$u_G$	Relative importance
MM1	$0.21 \pm 0.02$	$0.59 \pm 0.18$	$9 \pm 6$	$17.3 \pm 0.8$	$0.19_{-0.04}^{+0.05}$	$0.42 \pm 0.02$	$1.2_{-0.2}^{+0.3}$	$0.4_{-0.2}^{+0.1}$	$0.9_{-0.3}^{+0.5}$	$2.0_{-0.7}^{+0.8}$	$5.7_{-2.1}^{+3.5}$	7.7	G~B>T
MM2	$0.11 \pm 0.01$	$0.33 \pm 0.12$	$10 \pm 8$	$9.0 \pm 0.5$	$0.34_{-0.09}^{+0.10}$	$0.89 \pm 0.02$	$0.6_{-0.1}^{+0.3}$	$0.4_{-0.2}^{+0.2}$	$2.4_{-1.1}^{+1.6}$	$2.0_{-0.9}^{+1.0}$	$18.4_{-8.5}^{+12.4}$	6.5	B>G>T
MM3	$0.35 \pm 0.03$	$0.71 \pm 0.38$	$5 \pm 5$	$20 \pm 5$	$0.06_{-0.03}^{+0.04}$	$0.63 \pm 0.04$	$1.3_{-0.9}^{+1.0}$	$0.6_{-0.3}^{+0.2}$	$1.2_{-0.7}^{+1.4}$	$0.3_{-0.2}^{+0.2}$	$0.6_{-0.5}^{+1.0}$	0.7	G~B~T

## 0.6 pc core area

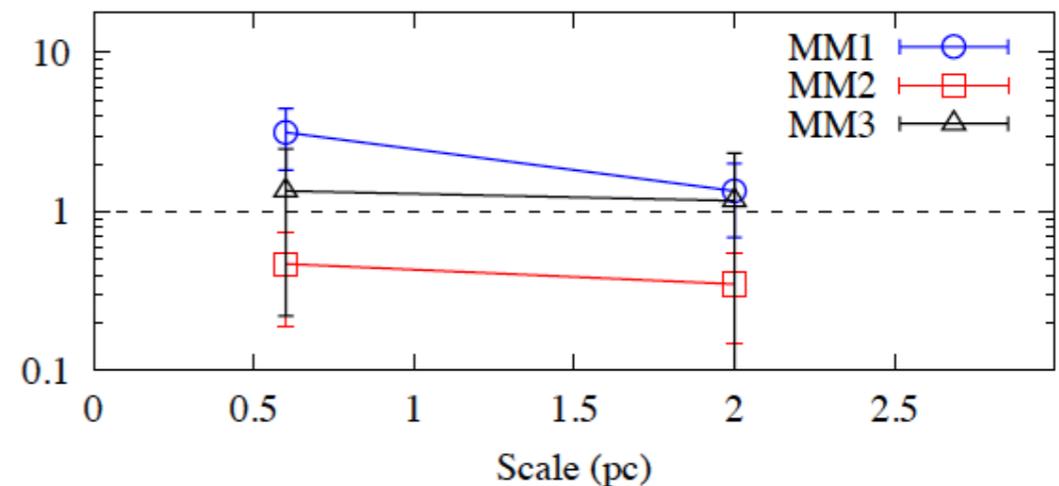
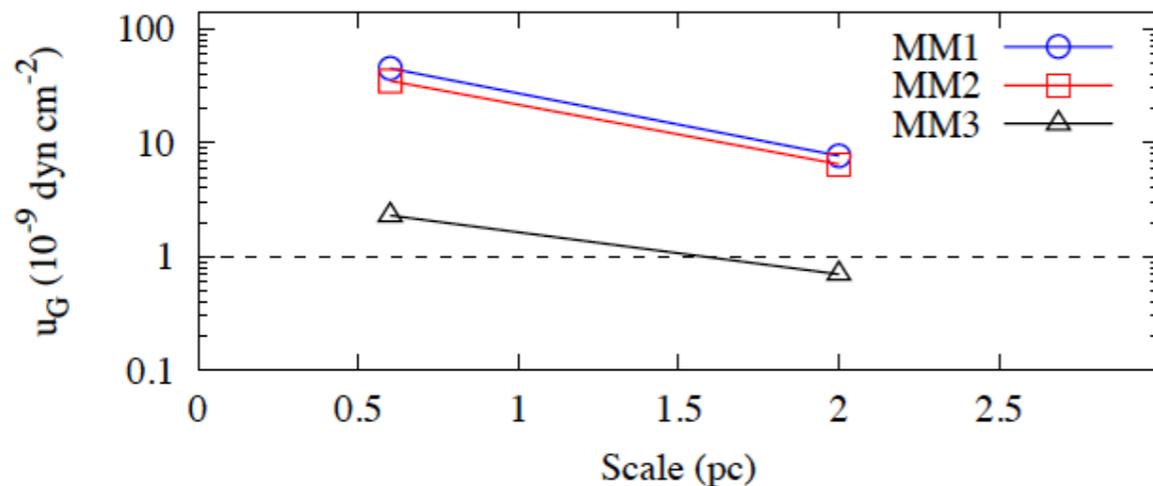
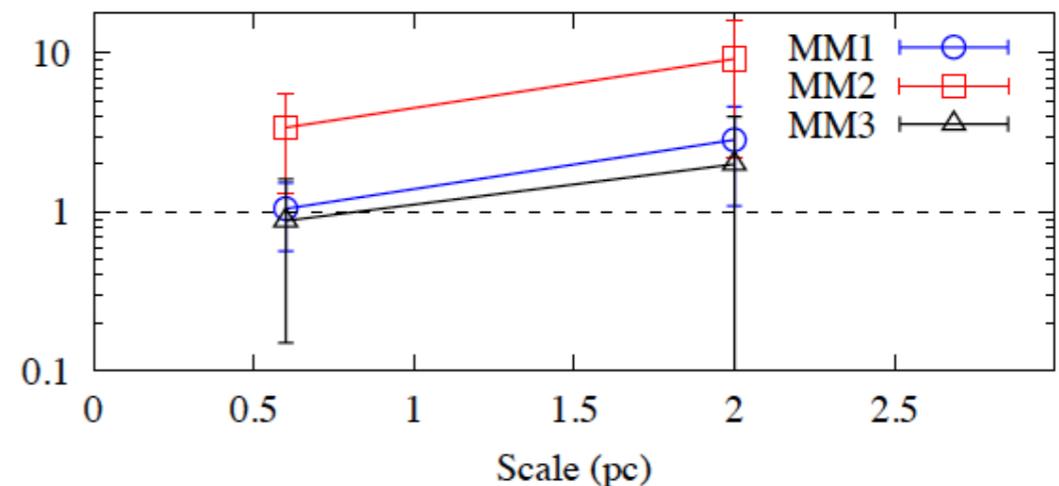
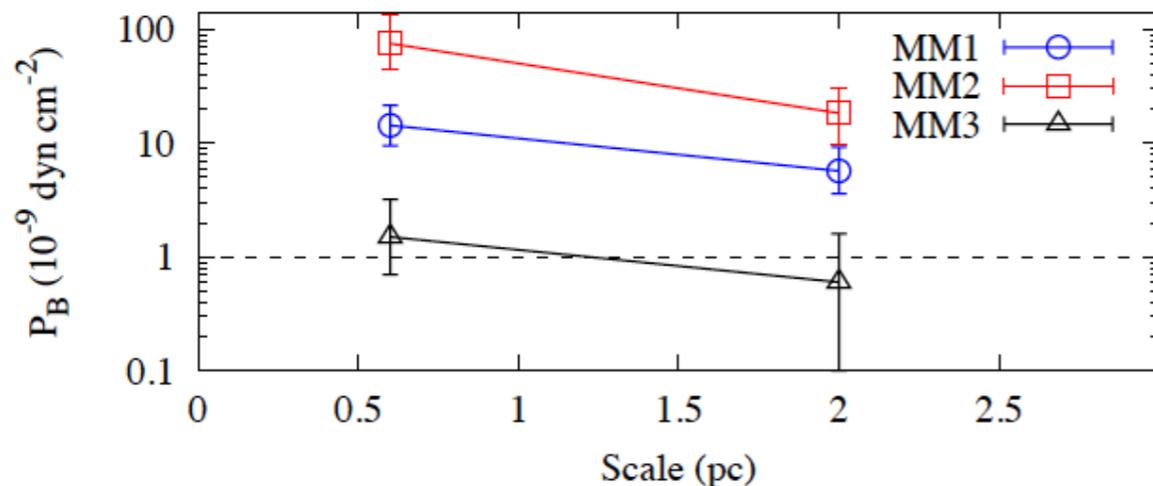
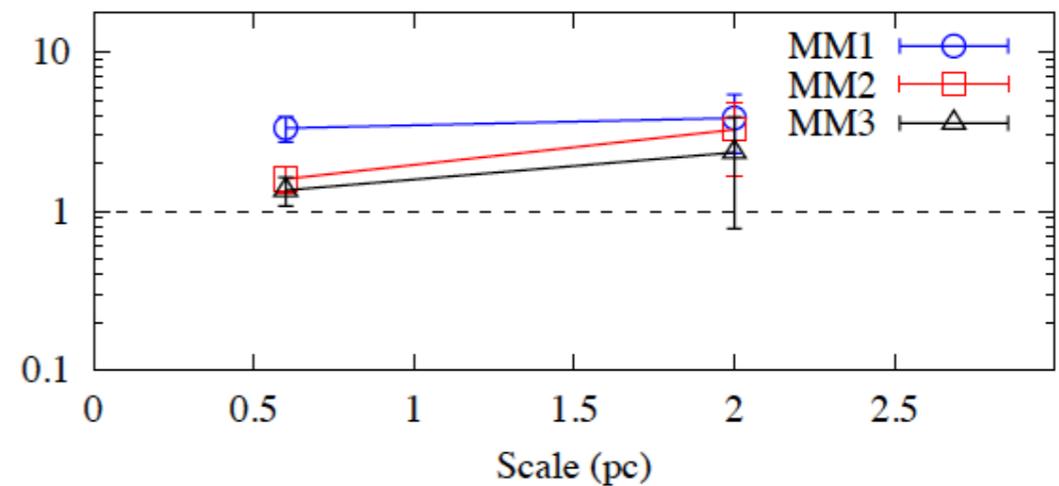
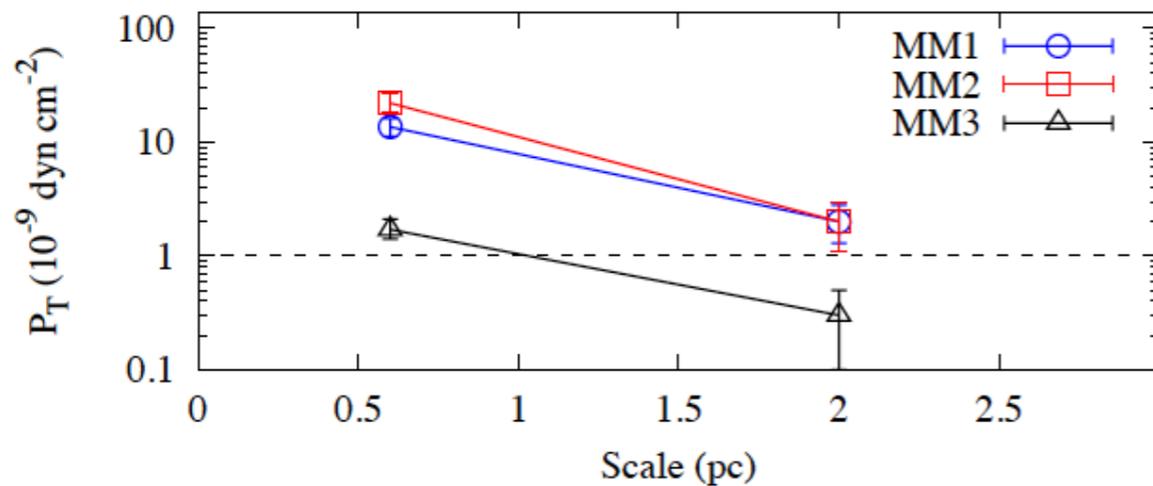
Object	$\frac{\sqrt{\langle B_t^2 \rangle}}{B_0}$	$\frac{\sqrt{\langle B_t^2 \rangle}}{B_0} N$	N	$\Delta\phi_B$ ( $^\circ$ )	$B_\perp$ (mG)	$\langle \Sigma_B \rangle$	$\lambda_{\text{obs}}$	$\alpha_{\text{vir}}$	$\alpha_{B,\text{vir}}$	$P_T$ ( $10^{-9}$ dyn/cm $^2$ )	$P_B$ ( $10^{-9}$ dyn/cm $^2$ )	$u_G$	Relative importance
MM1	$0.20 \pm 0.03$	$0.98 \pm 0.22$	$25 \pm 13$	$18 \pm 2$	$0.49_{-0.09}^{+0.11}$	$0.55 \pm 0.04$	$1.1_{-0.2}^{+0.3}$	$0.5_{-0.1}^{+0.1}$	$1.1_{-0.3}^{+0.4}$	$13.6_{-2.4}^{+2.6}$	$14.3_{-4.8}^{+7.2}$	45.1	G>B~T
MM2	$0.15 \pm 0.03$	$0.54 \pm 0.17$	$14 \pm 10$	$10 \pm 2$	$1.12_{-0.26}^{+0.37}$	$0.68 \pm 0.05$	$0.5_{-0.2}^{+0.1}$	$0.9_{-0.1}^{+0.1}$	$4.9_{-1.7}^{+3.3}$	$22.0_{-3.9}^{+4.3}$	$74.9_{-30.8}^{+57.6}$	35.0	B>G>T
MM3	$0.32 \pm 0.03$	$1.06 \pm 0.46$	$13 \pm 11$	$20 \pm 5$	$0.16_{-0.05}^{+0.07}$	$0.66 \pm 0.05$	$0.9_{-0.3}^{+0.4}$	$0.9_{-0.2}^{+0.3}$	$2.1_{-0.8}^{+1.4}$	$1.7_{-0.3}^{+0.4}$	$1.5_{-0.8}^{+1.7}$	2.3	G>T~B

*isolated single numbers have limited information  
need for joint analyses*

# Analysis — Subtle Balance $G$ vs $T$ vs $B$

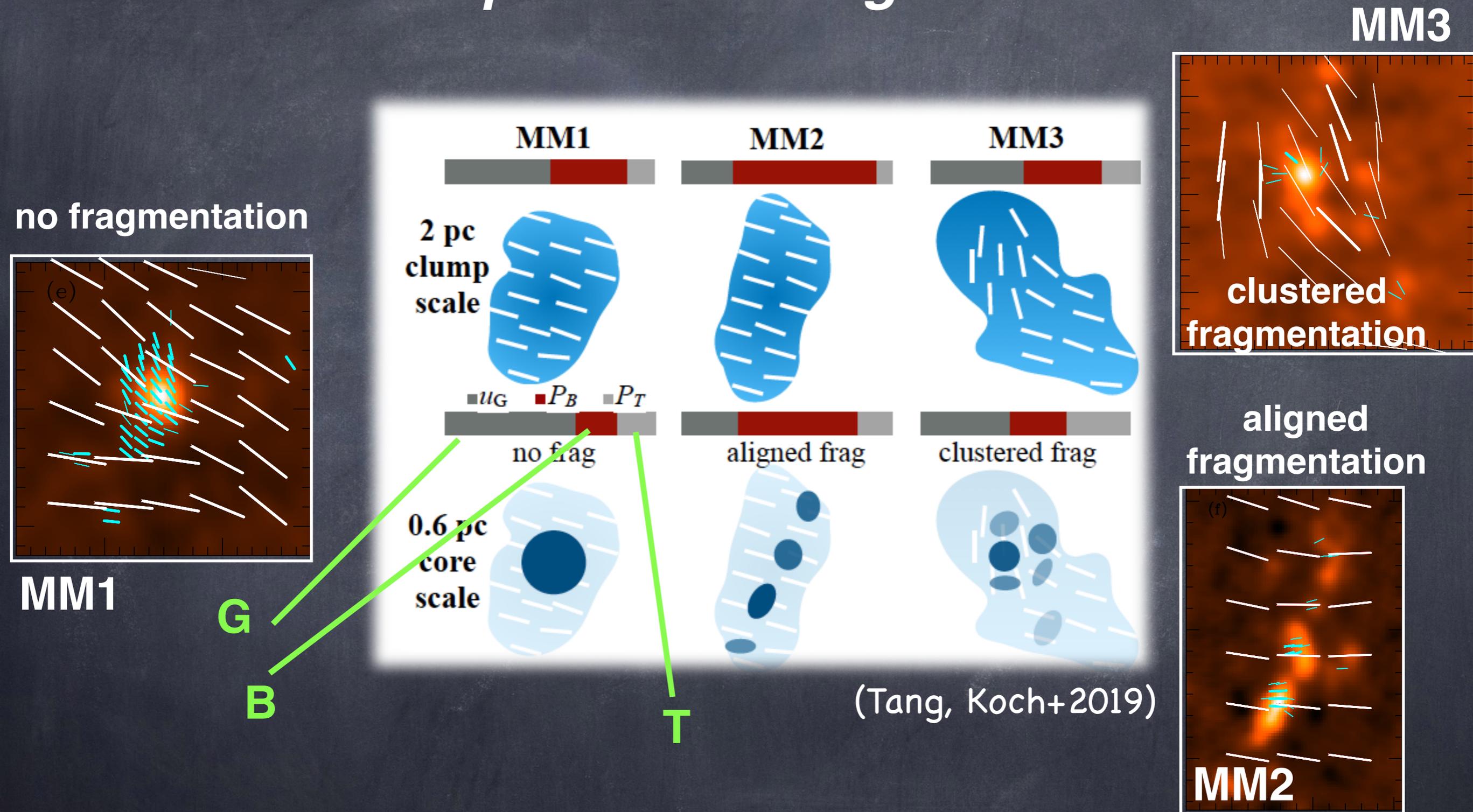
$u_G$   $P_T$   $P_B$

Ratios



# Analysis — Subtle Balance $G$ vs $T$ vs $B$

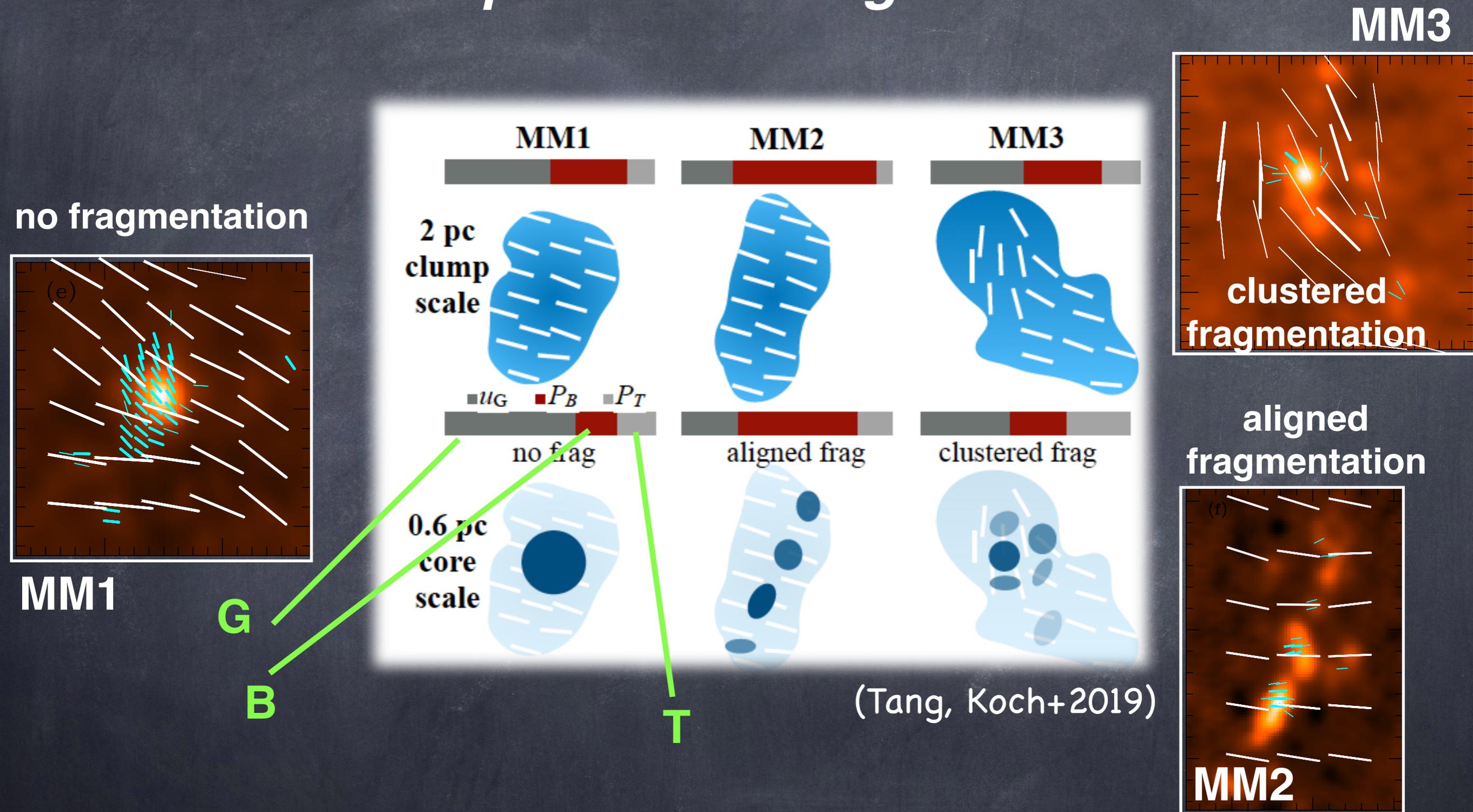
## Consequence for Fragmentation



*different relative importance between  $B$ -field, turbulence, and gravity seems to control fragmentation towards next smaller scale; also seen in simulation work by e.g., Seifried+2015*

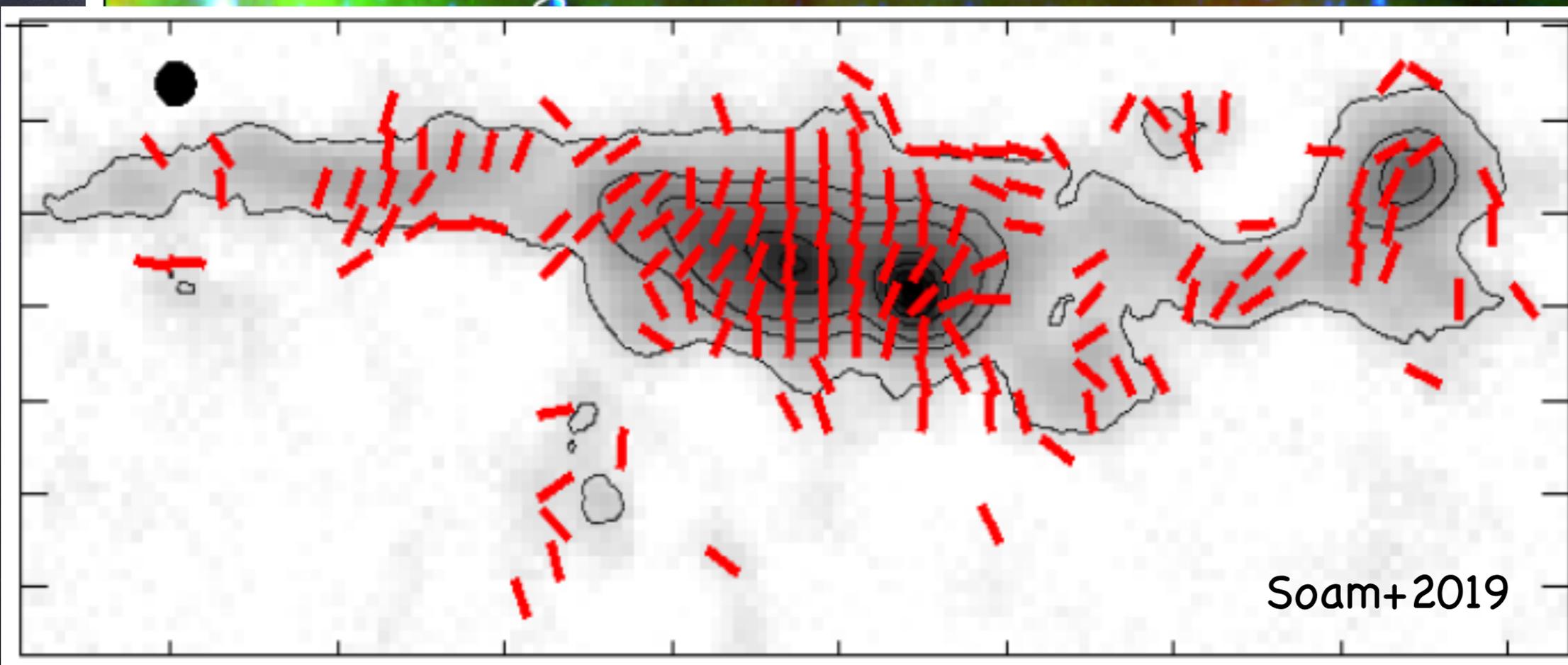
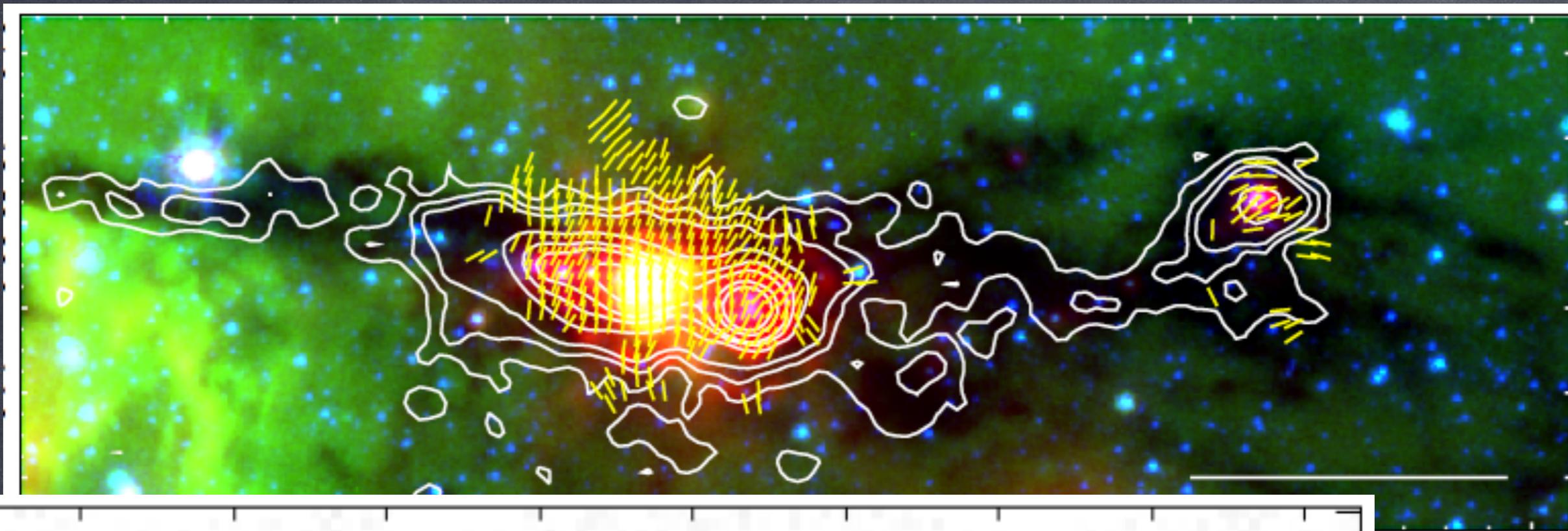
# Analysis – Subtle Balance $G$ vs $T$ vs $B$

## Consequence for Fragmentation



*JCMT / POL-2 science:  
ideally suited to probe “initial conditions” for fragmentation*

# CSO / SHARP 350 $\mu\text{m}$ vs JCMT / POL-2 850 $\mu\text{m}$



JCMT  
POL-2

# Conclusions

- role of B-field is variable (both in scales and locations)
- JCMT / POL-2 “intermediate” resolution between several-arcmin scales and (sub-)arcsec scales: “initial-condition” scale to map B-field morphologies
- generally: joint analyses is important  
B-field needs to be compared to turbulence, gravity (and more? feedback etc)
- need for ancillary (matching) data, in particular line kinematics;  
good and complete coverage is essential! smaller / larger areas give different results!
- example: initial-conditions for fragmentation in IRDC G34;  
balance between  $G$ ,  $B$ ,  $T$  derived for clump and core scales, trends across 2 different scales are different and can explain different fragmentation types
- need to develop analysis tools:  $\delta$  is a key observable, leading to local field strength measurement and local force ratio  $\Sigma_B$