

Determining the systematic errors in fits of dust thermal emission

The role of laboratory data in upcoming
models

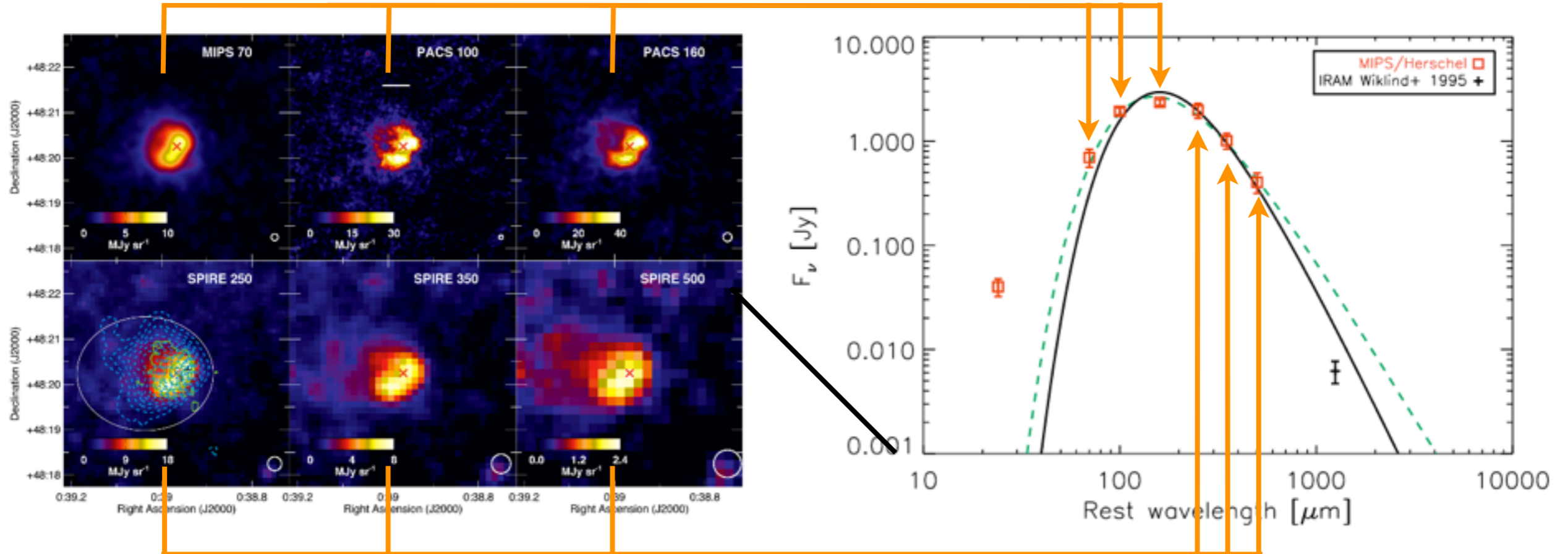


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Seoul, Feb 1st 2018

Context: Dust masses from the SED

NGC 185 (De Looze+16)



Dust emission SED
(multi- λ photometry)

FIR: 100 - 1000 μm



Interpretation
through model



Dust physical quantities:
Mass, temperature...

D/G: gas mass tracer?

Dust formation history at high z
(Dust budget crisis)

Context: Dust masses from the SED

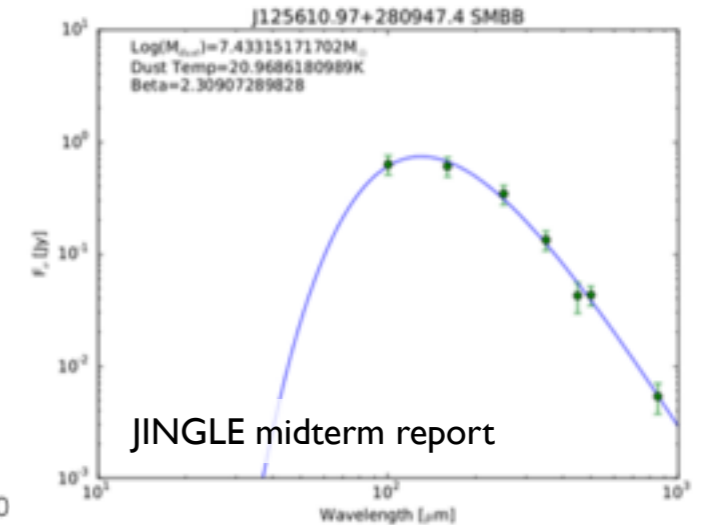
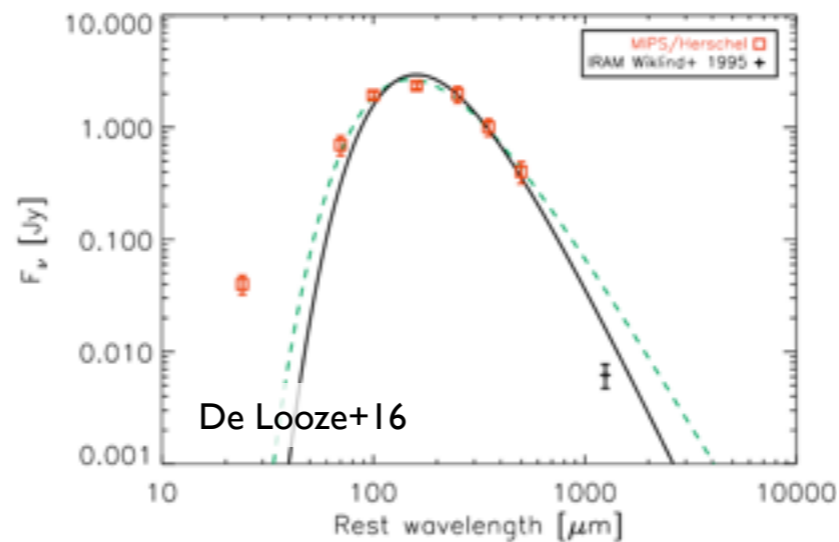
Modified blackbody (MBB)

Issues:

- Parameter degeneracy
- No single T
- No constant β / simple power law
- κ_λ depends on many things:
 - Temperature
 - Composition
 - Grain shape

Alternative forms of κ_λ :

- Observational (e.g., Draine)
 - No info on actual dust properties
- **Laboratory measurements**

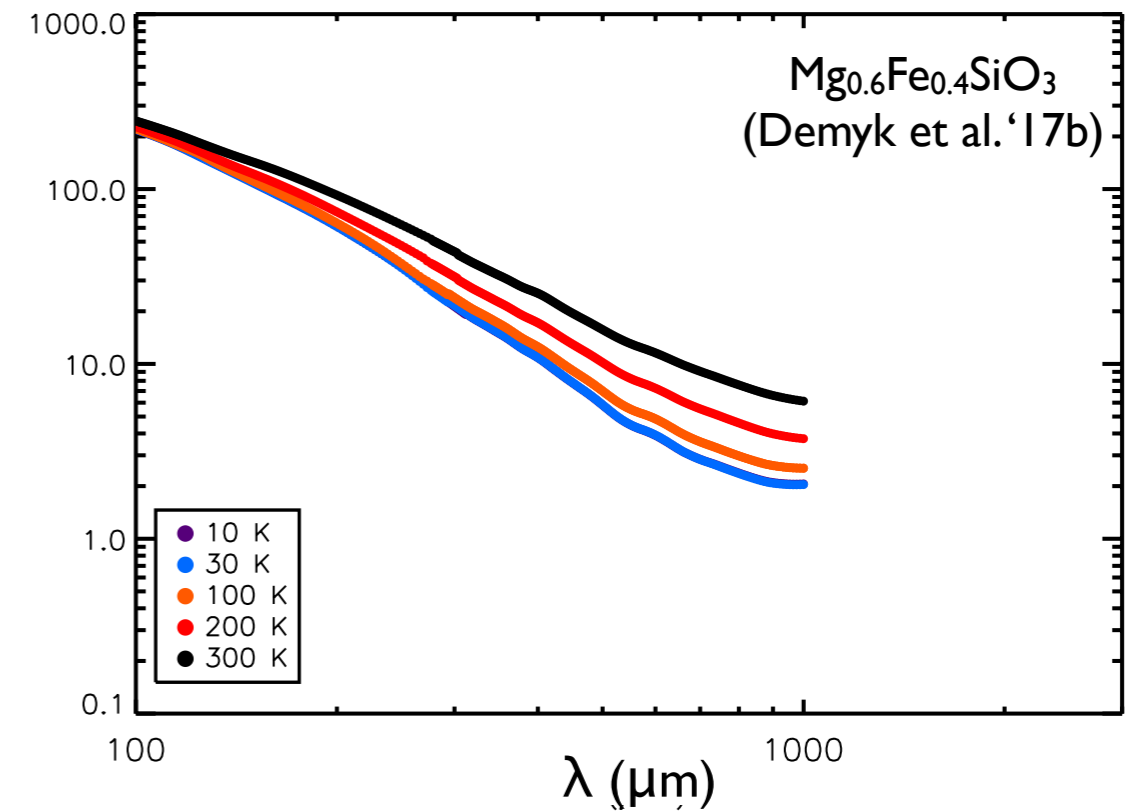
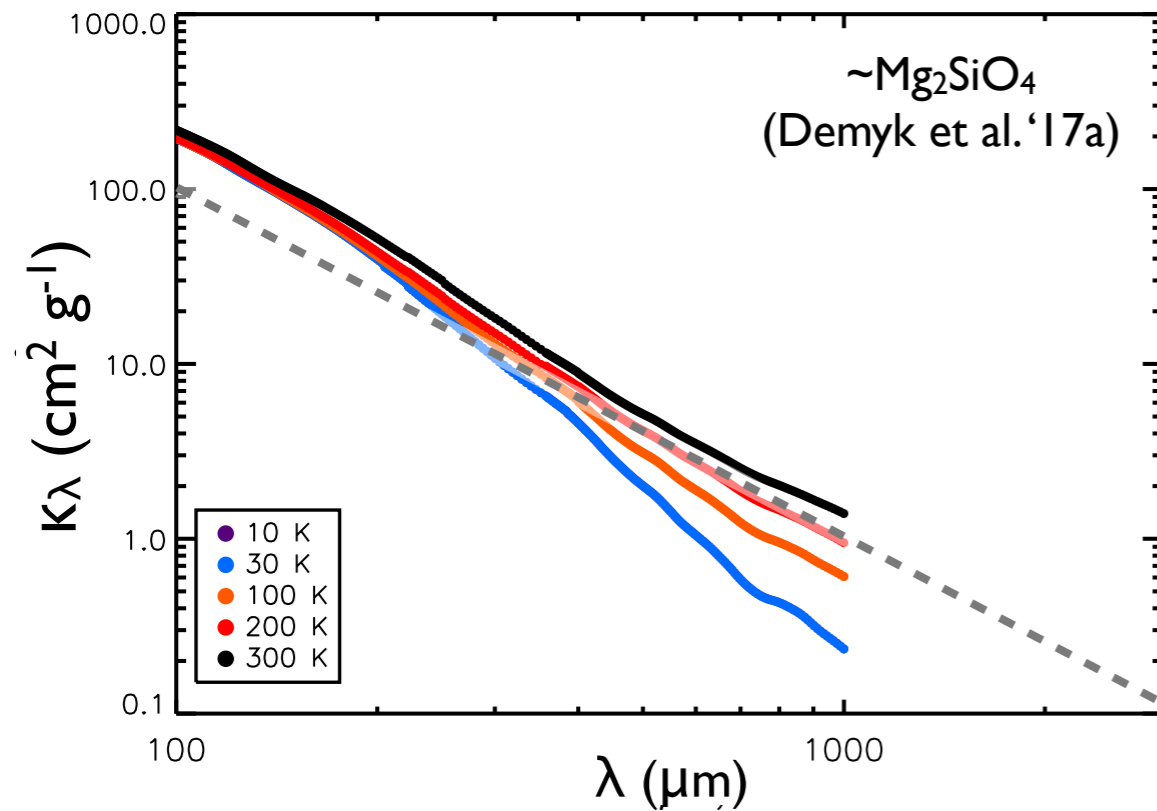
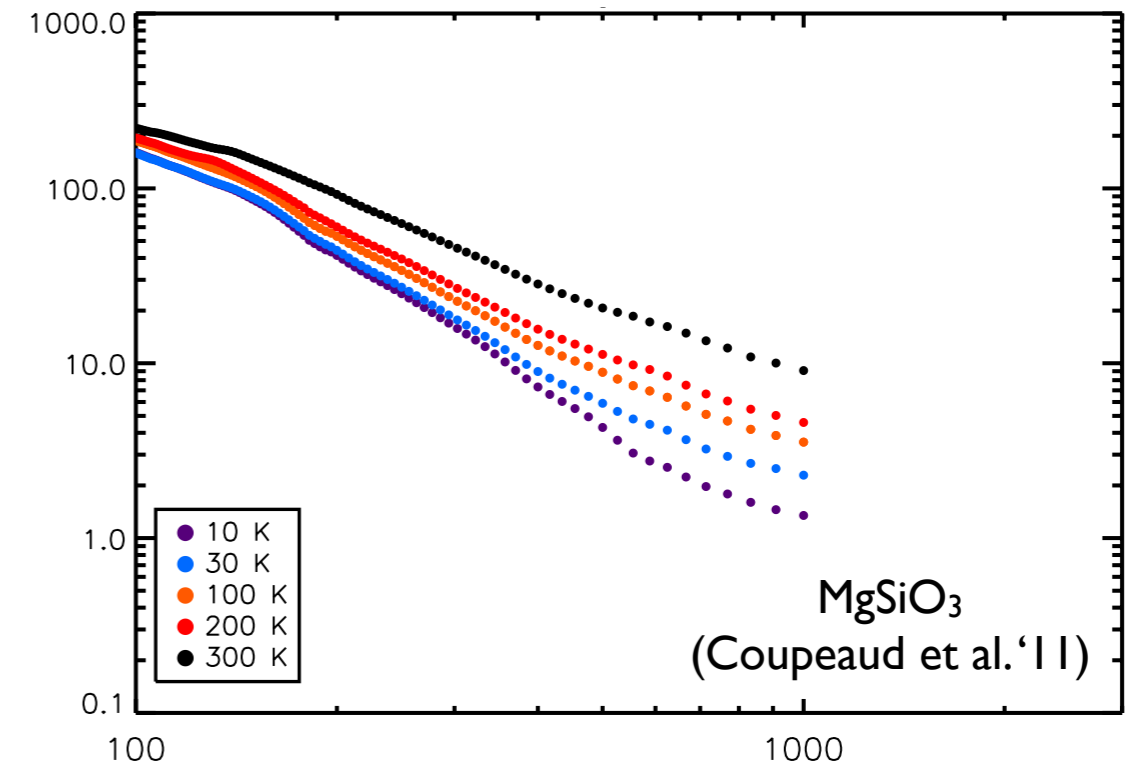
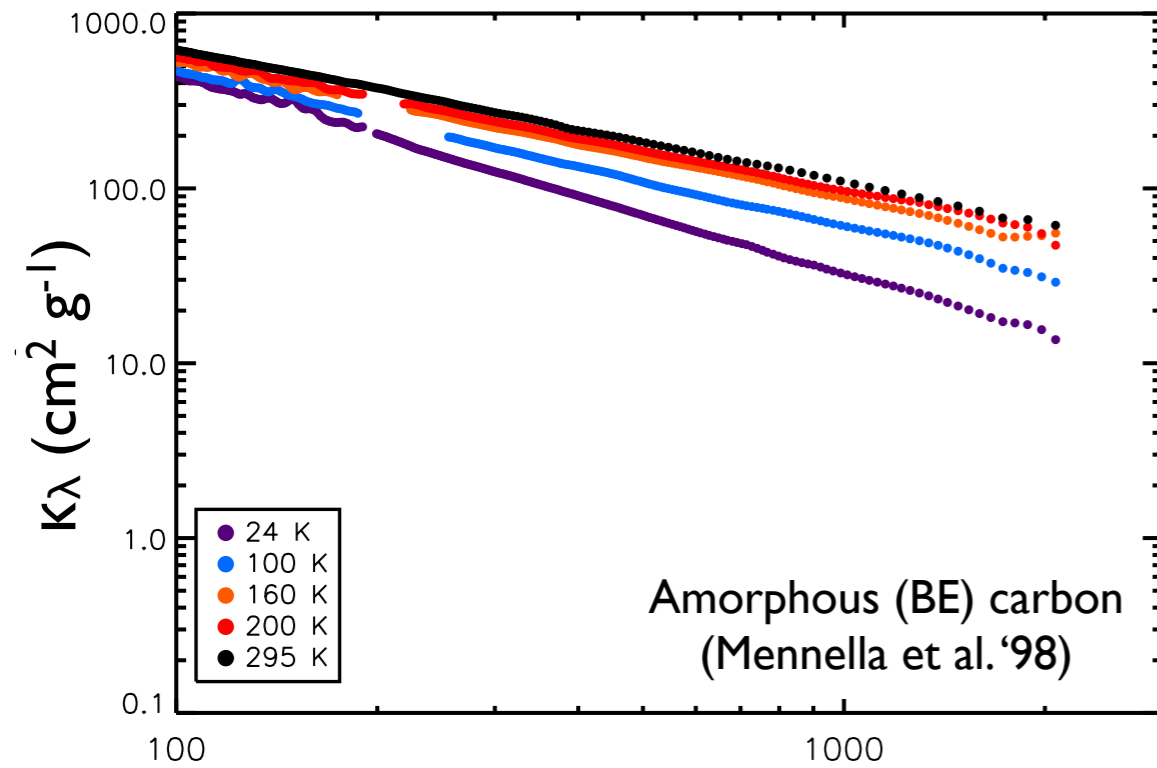


$$F_\lambda = \overbrace{M_{d,surf}}^{\tau_\lambda} \times \underbrace{\kappa_\lambda}_{\kappa_0 \times (\lambda/\lambda_0)^{-\beta}} \times B_\lambda(T)$$

↓

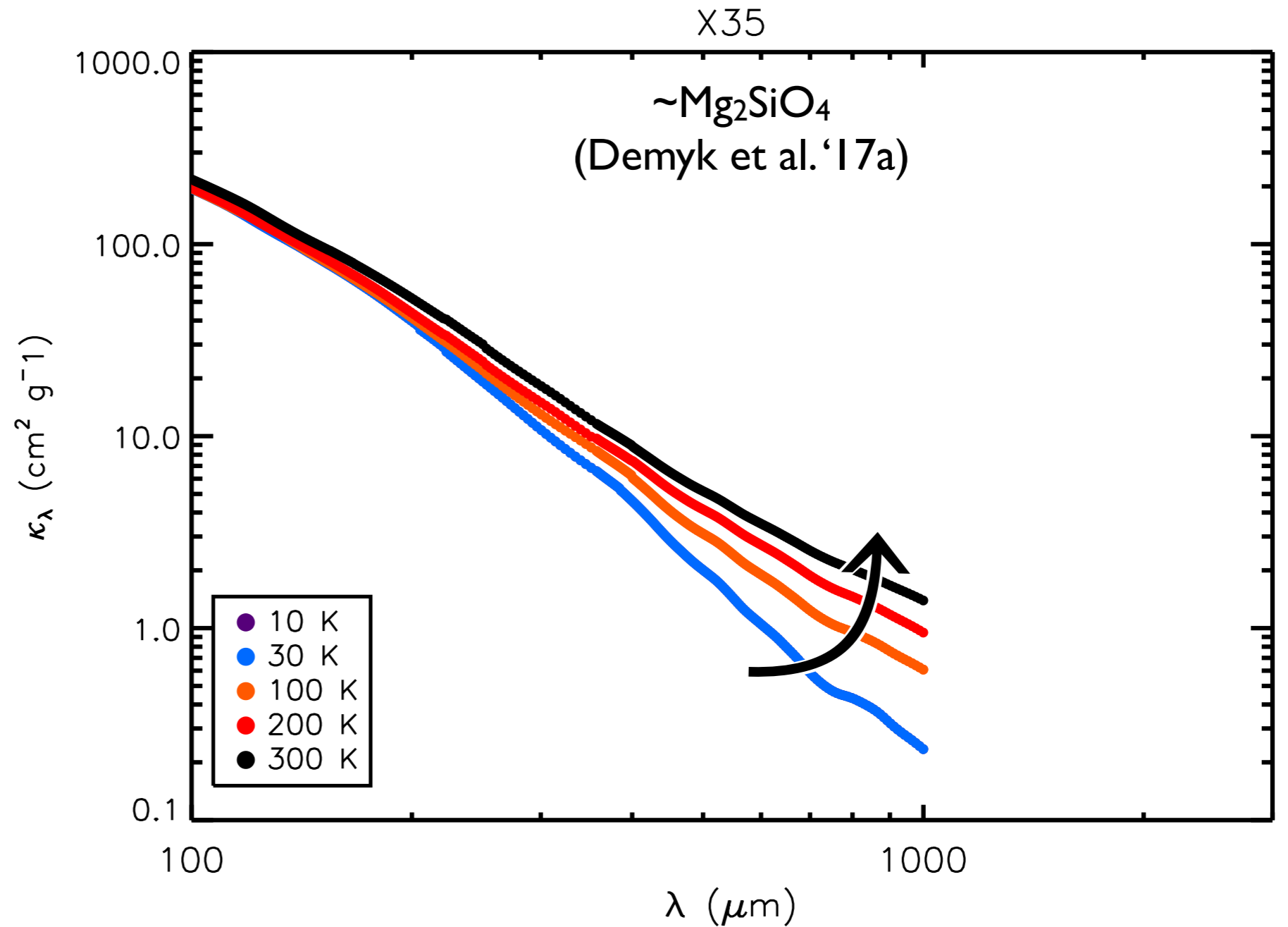
$T, \beta, \tau_0 [\kappa_0]$

Laboratory dust opacities



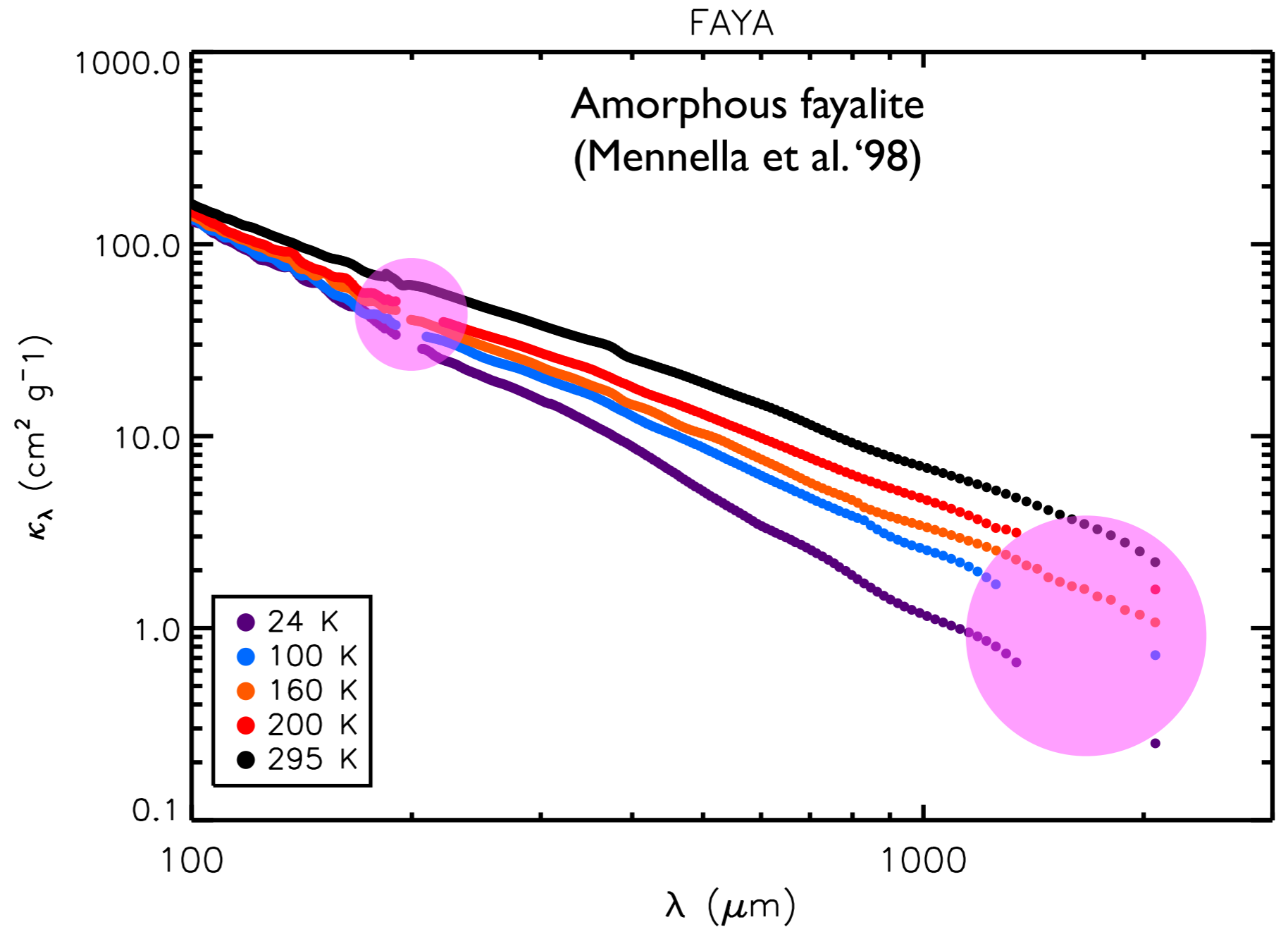
Laboratory dust opacities

- Interpolation on T



Laboratory dust opacities

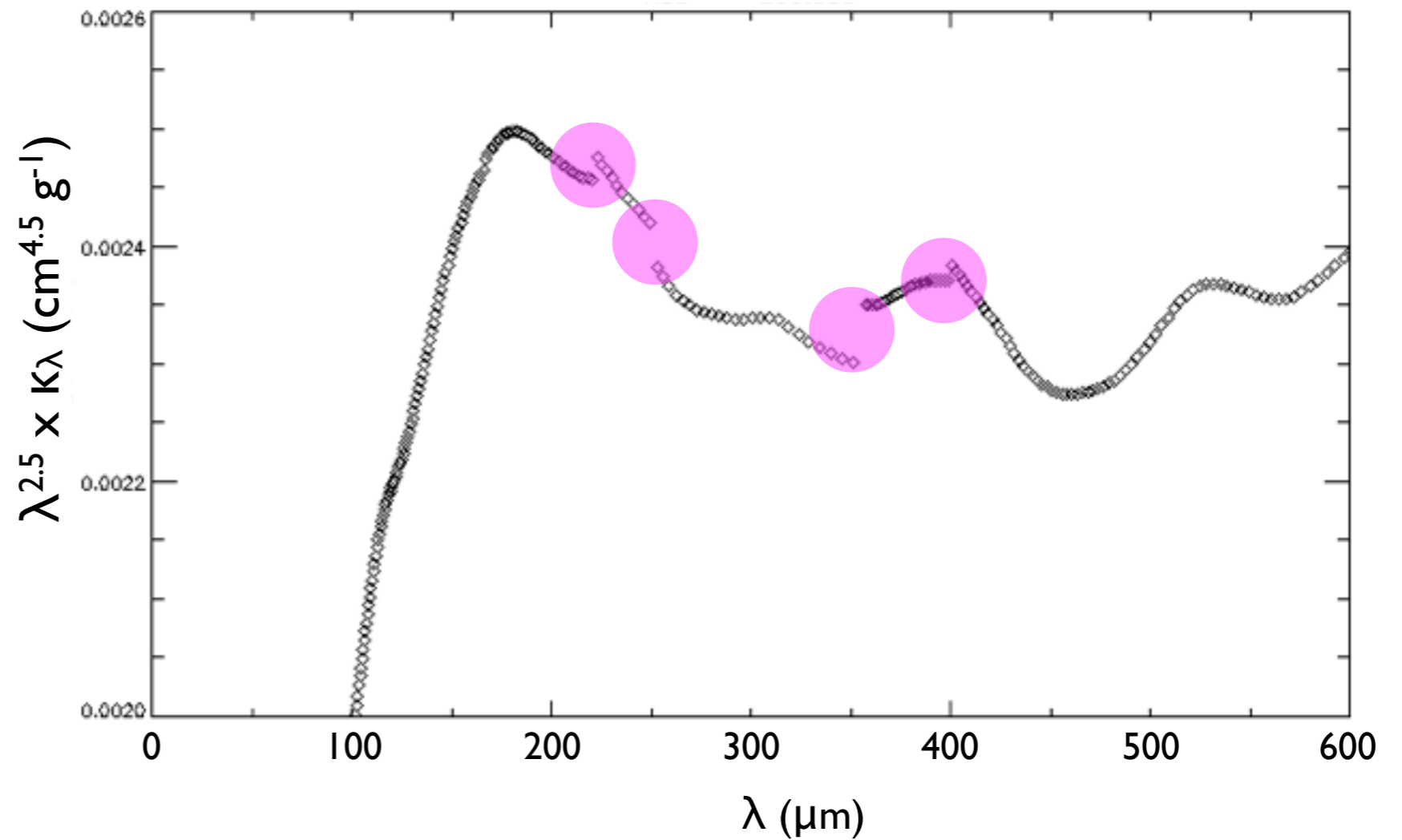
- Interpolation on T
- Interpolation on λ
- 2D interpolation



Laboratory dust opacities

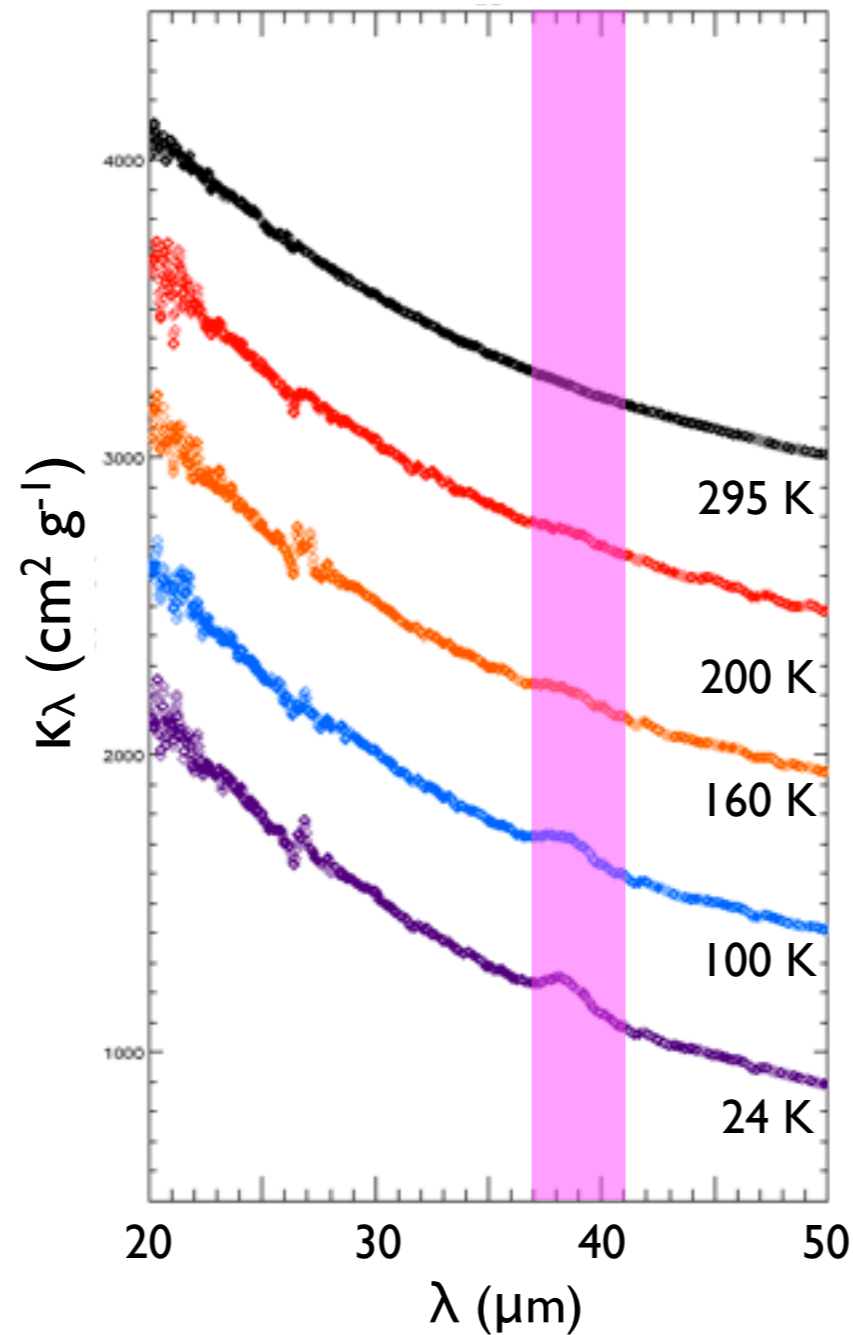
$\sim\text{Mg}_2\text{SiO}_4$ (Demyk et al. '17a), $T = 200$ K

- Interpolation on T
- Interpolation on λ
- 2D interpolation
- Smooth (if necessary)
- Correct for artifacts

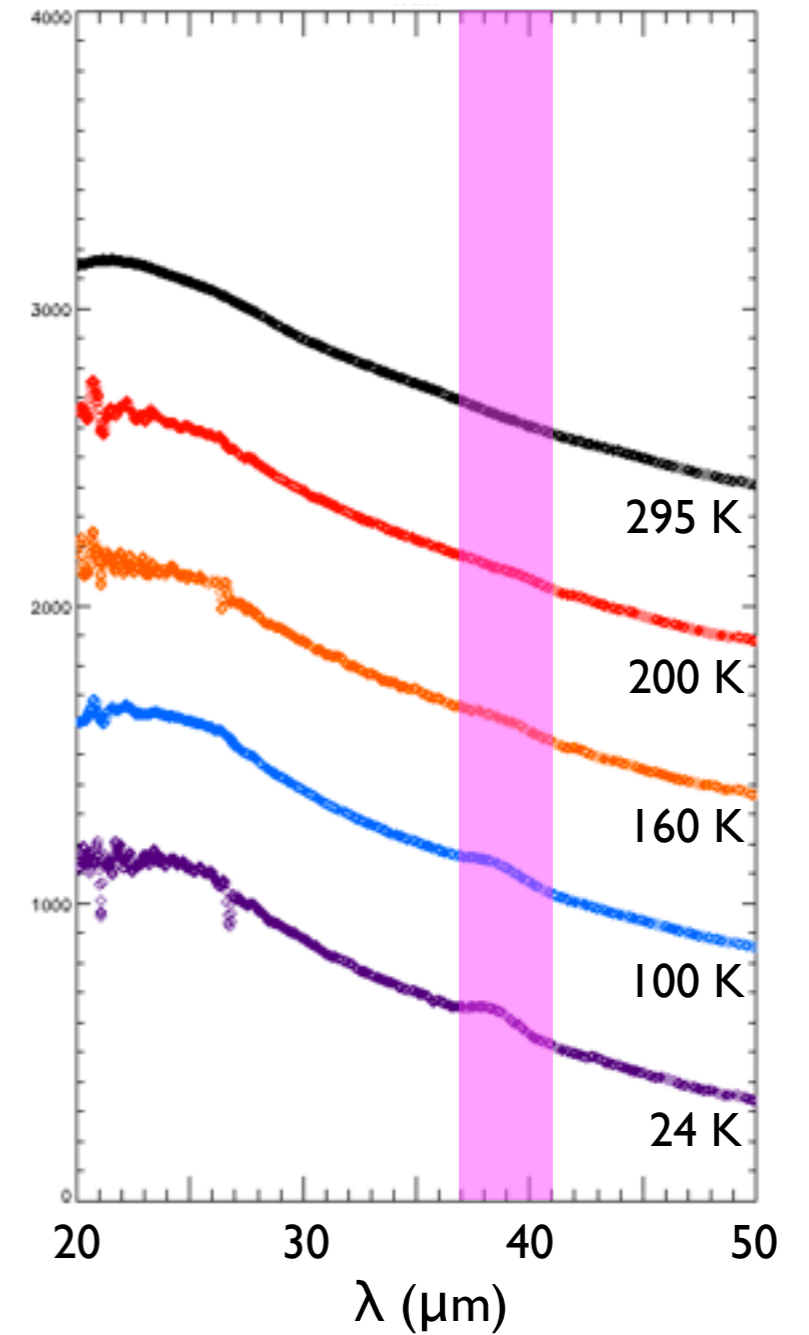


Laboratory dust opacities

BE Carbon (Mennella '98)

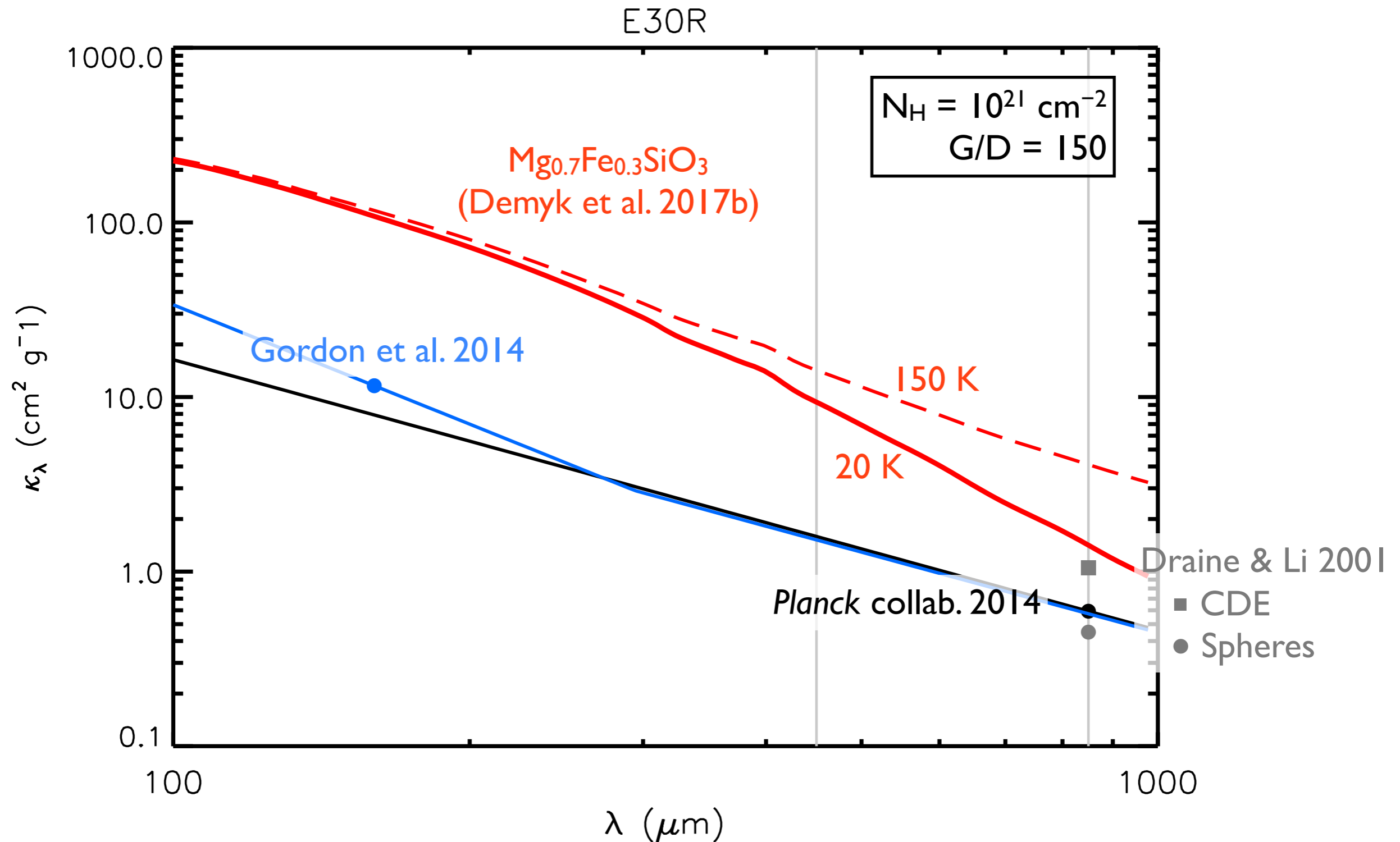


Am. Fayalite (Mennella '98)

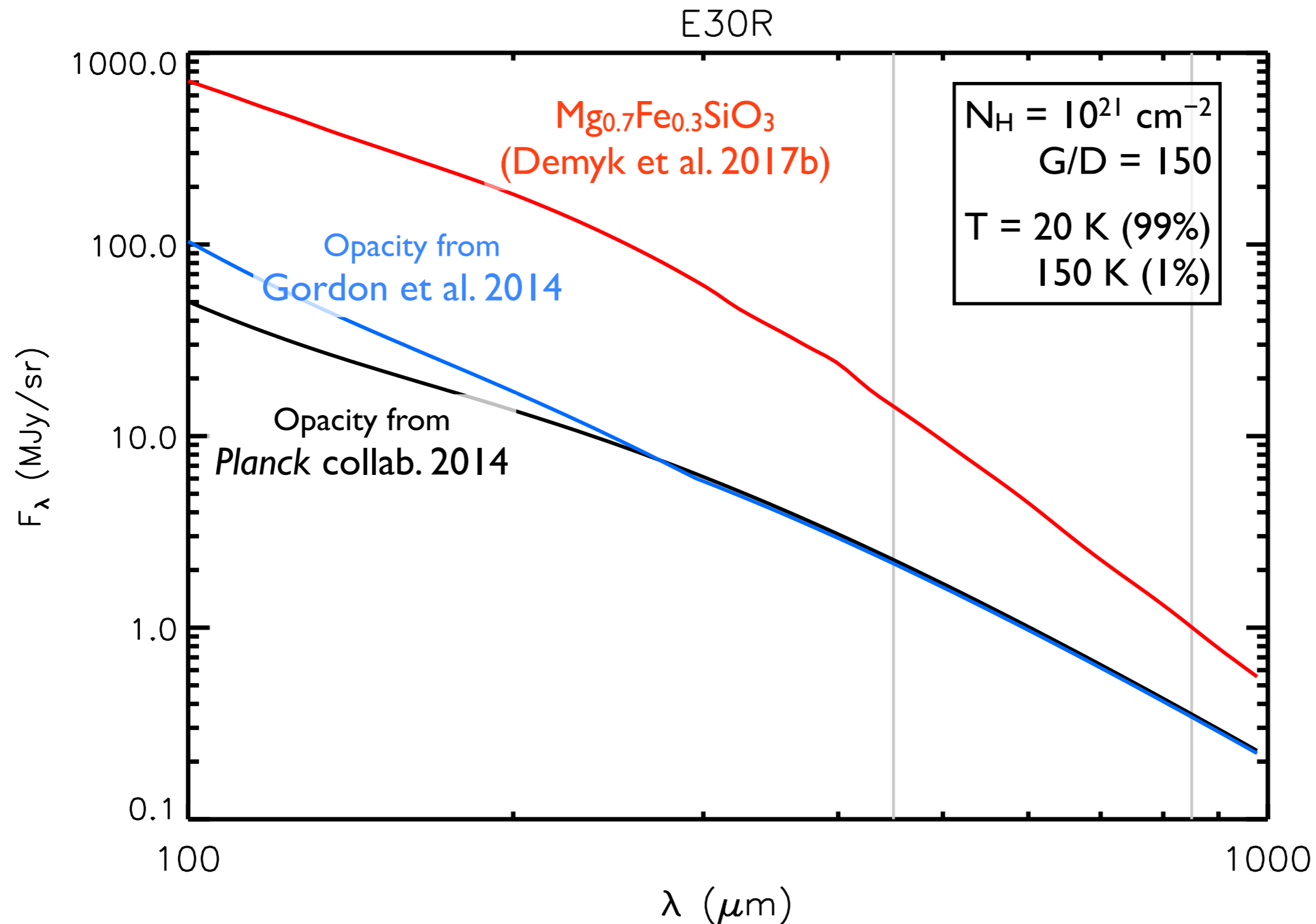


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- Interpolation on λ
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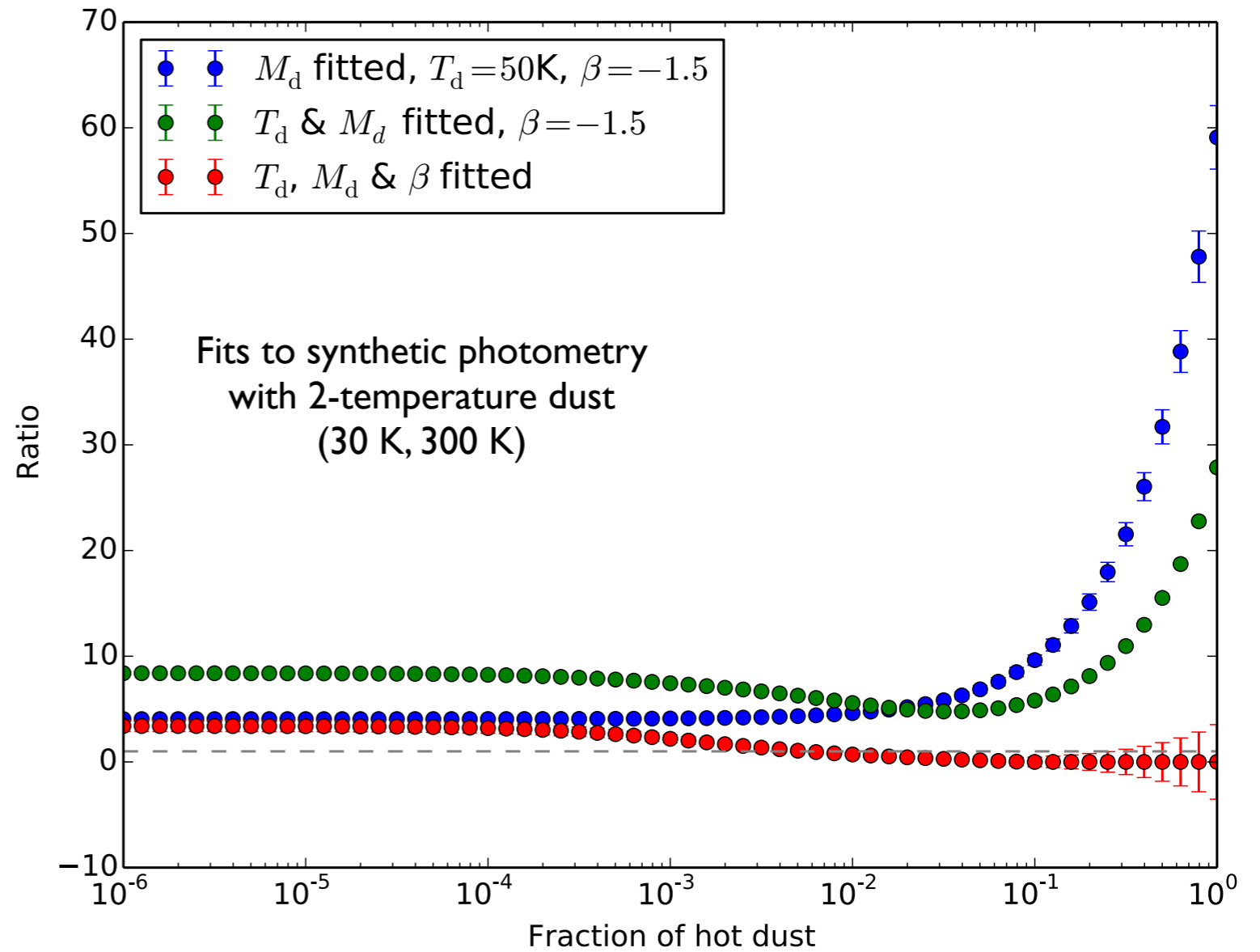
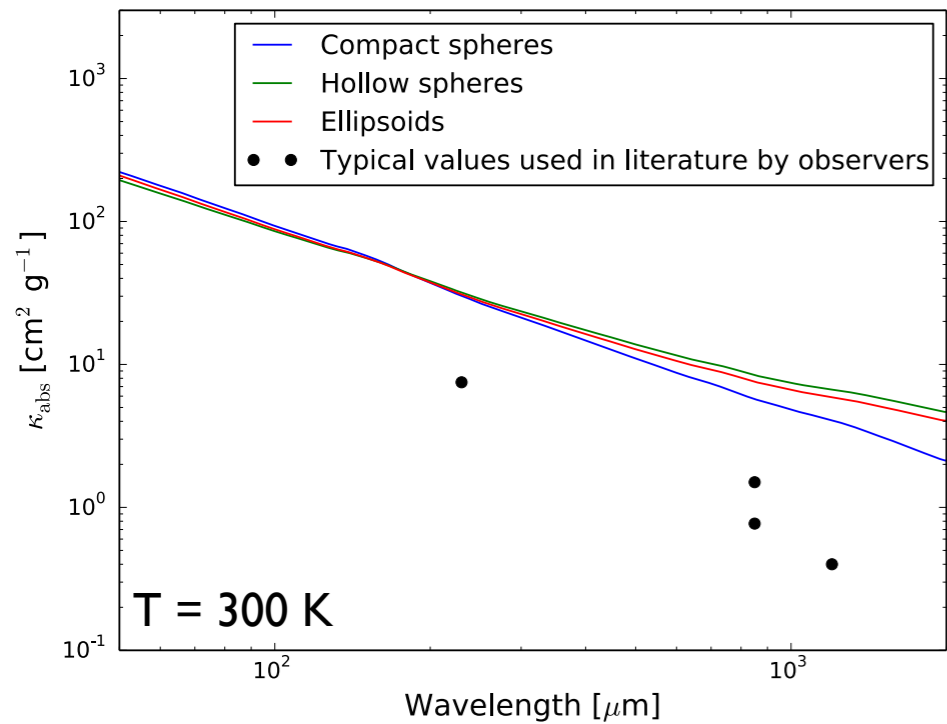
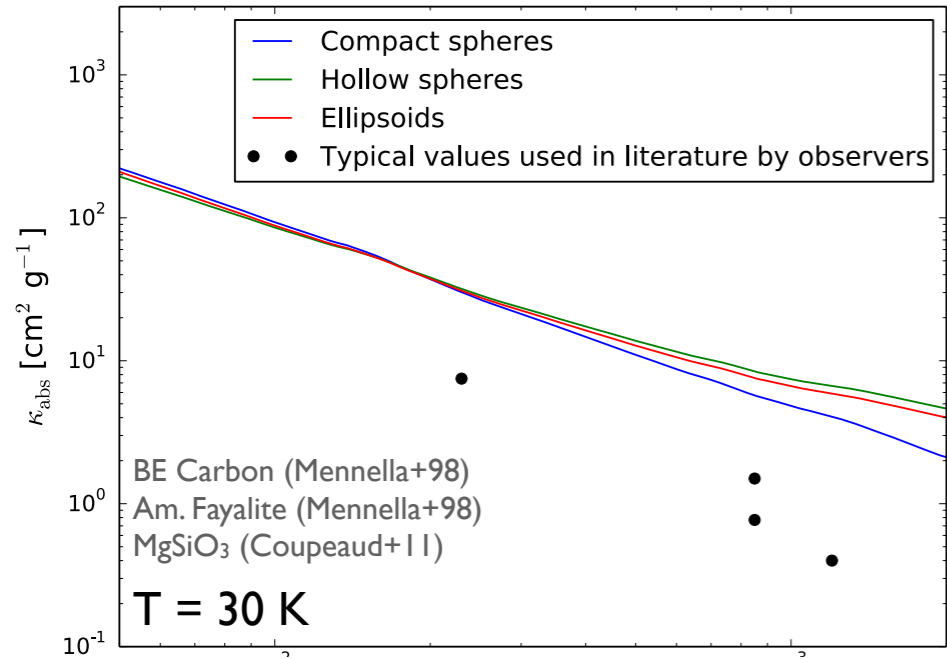
Lab results vs. observations



Lab results vs. (synthetic) observations



Bias estimation (Work by Peter Scicluna)



Conclusions

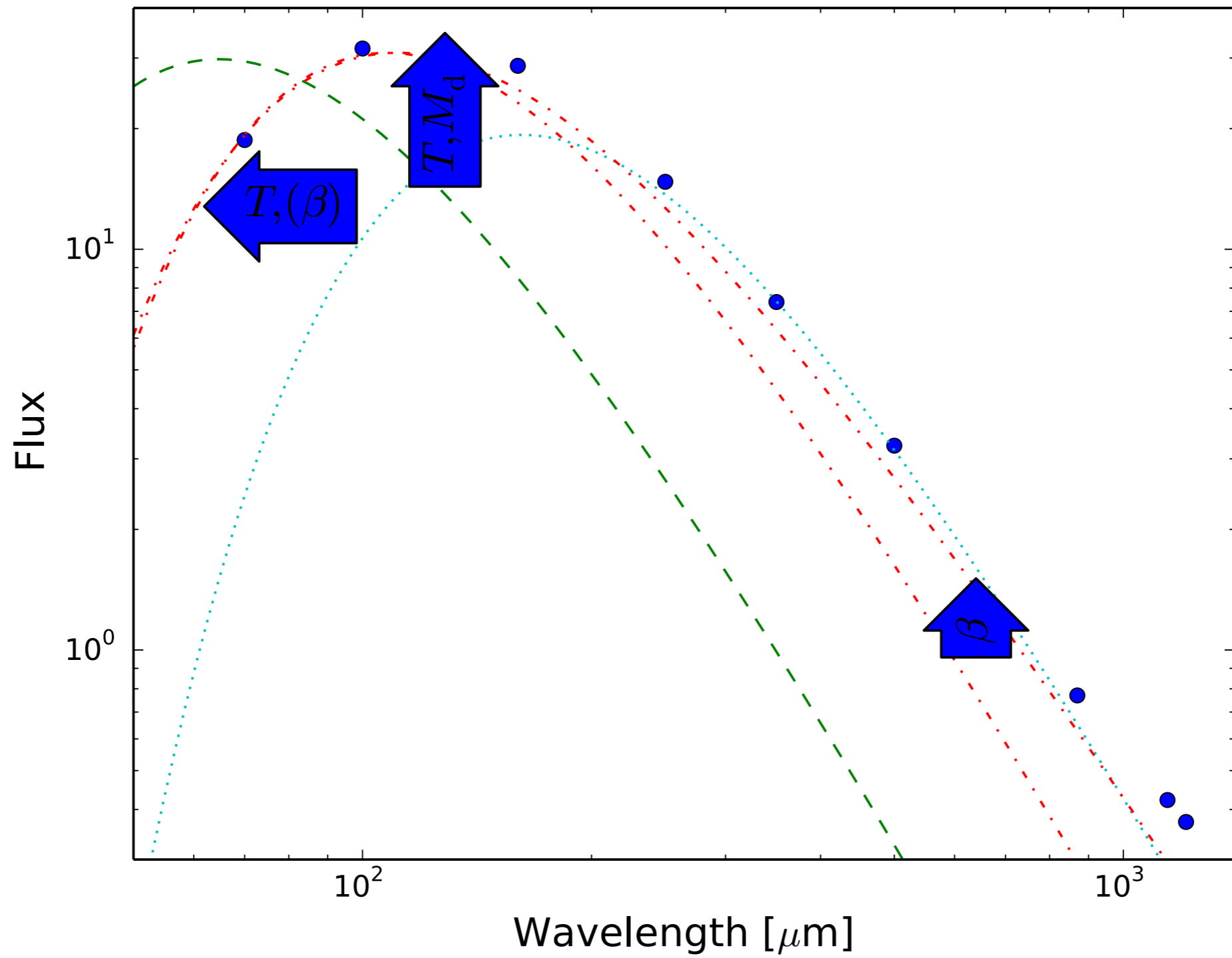
- Dust mass determination depends on choice of opacity
- Large differences between lab-derived and observation-derived opacities
 - Power law (single- β) model inadequate
 - Dependence on T
- Fits of synthetic photometry
 - ⇒ Mass overestimated by up to $\sim 10x$

Conclusions

- Dust mass determination depends on choice of opacity
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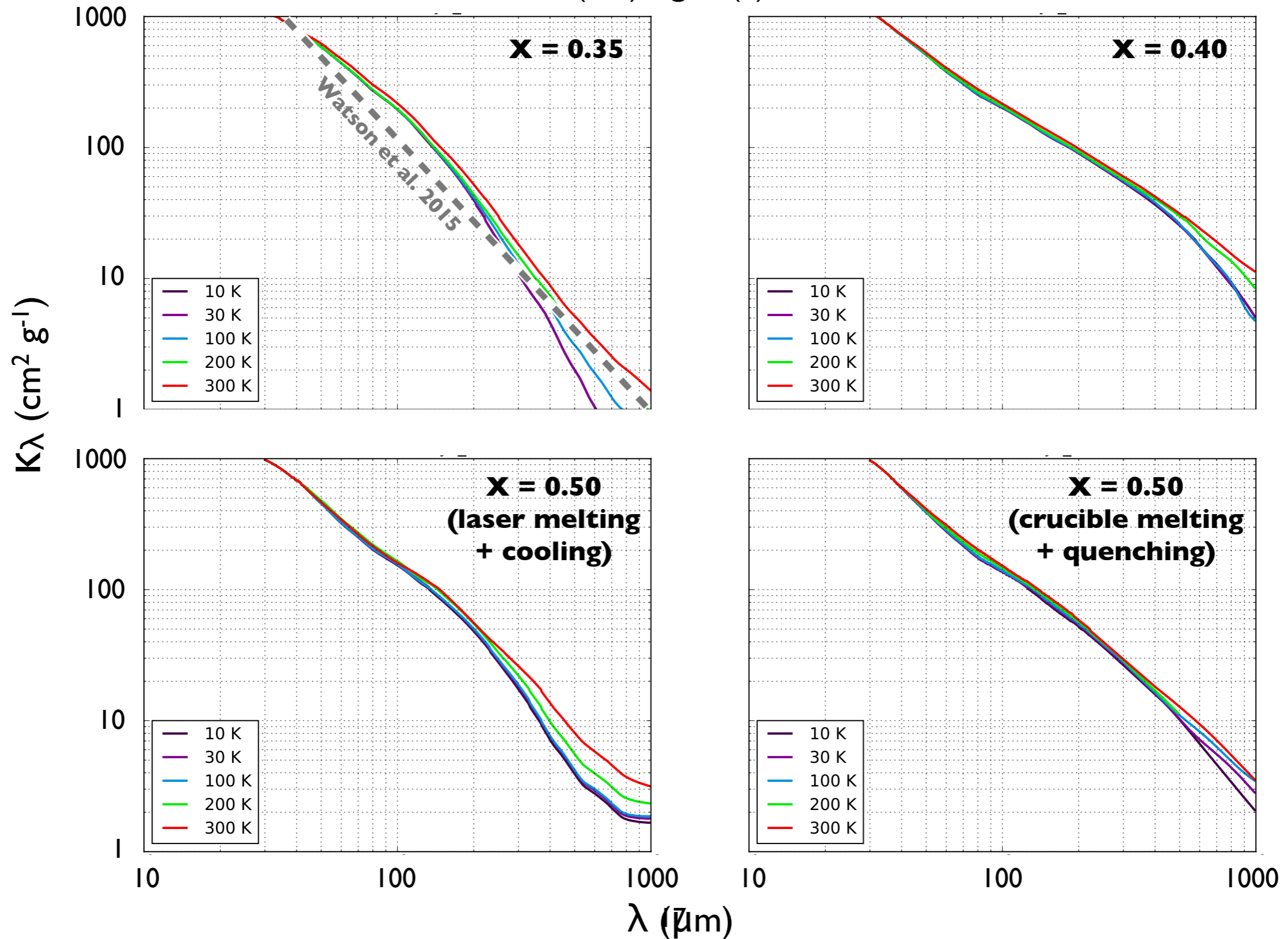
PRELIMINARY

Thank you for your
attention!

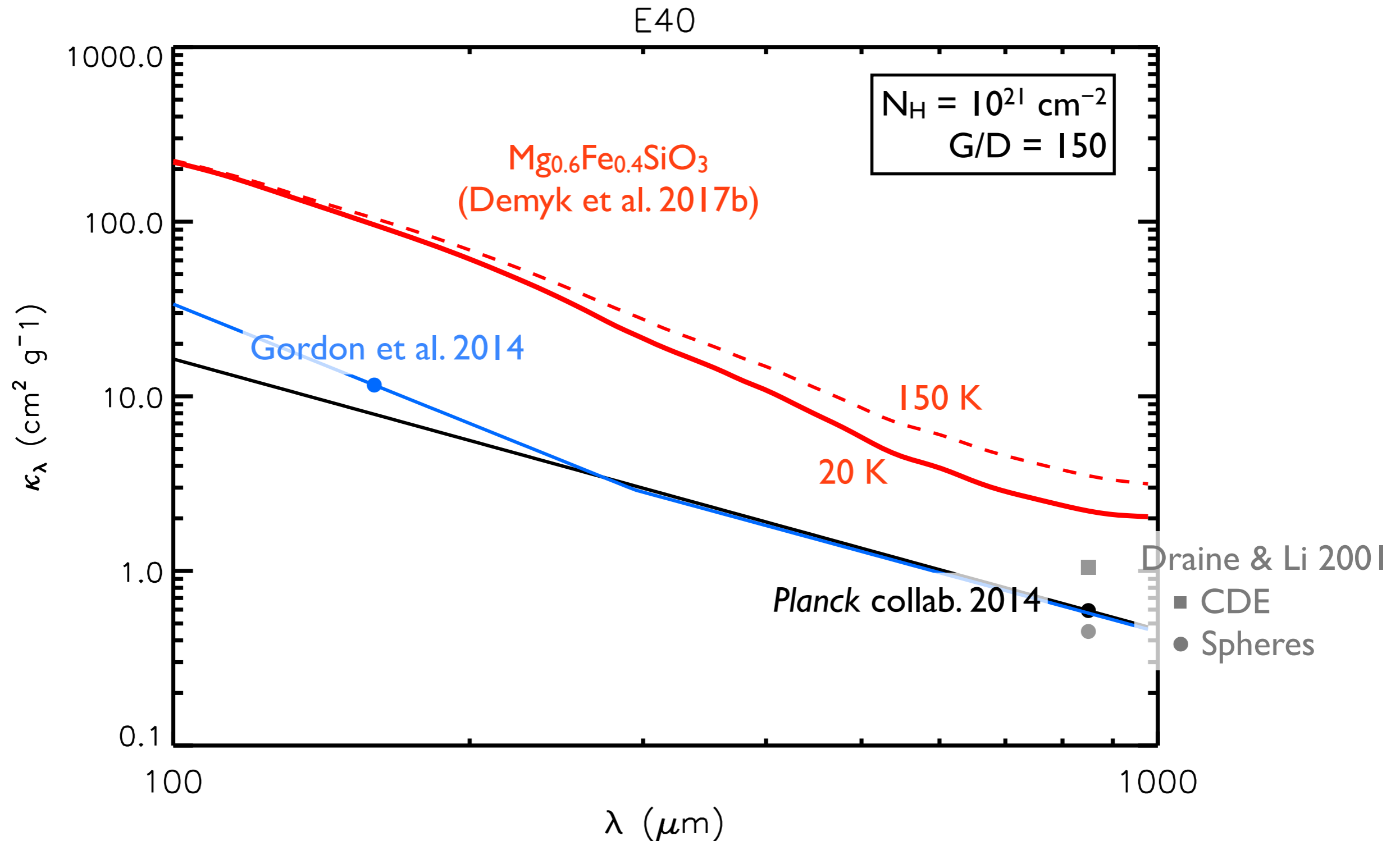


κ_λ : Demyk et al. 2017

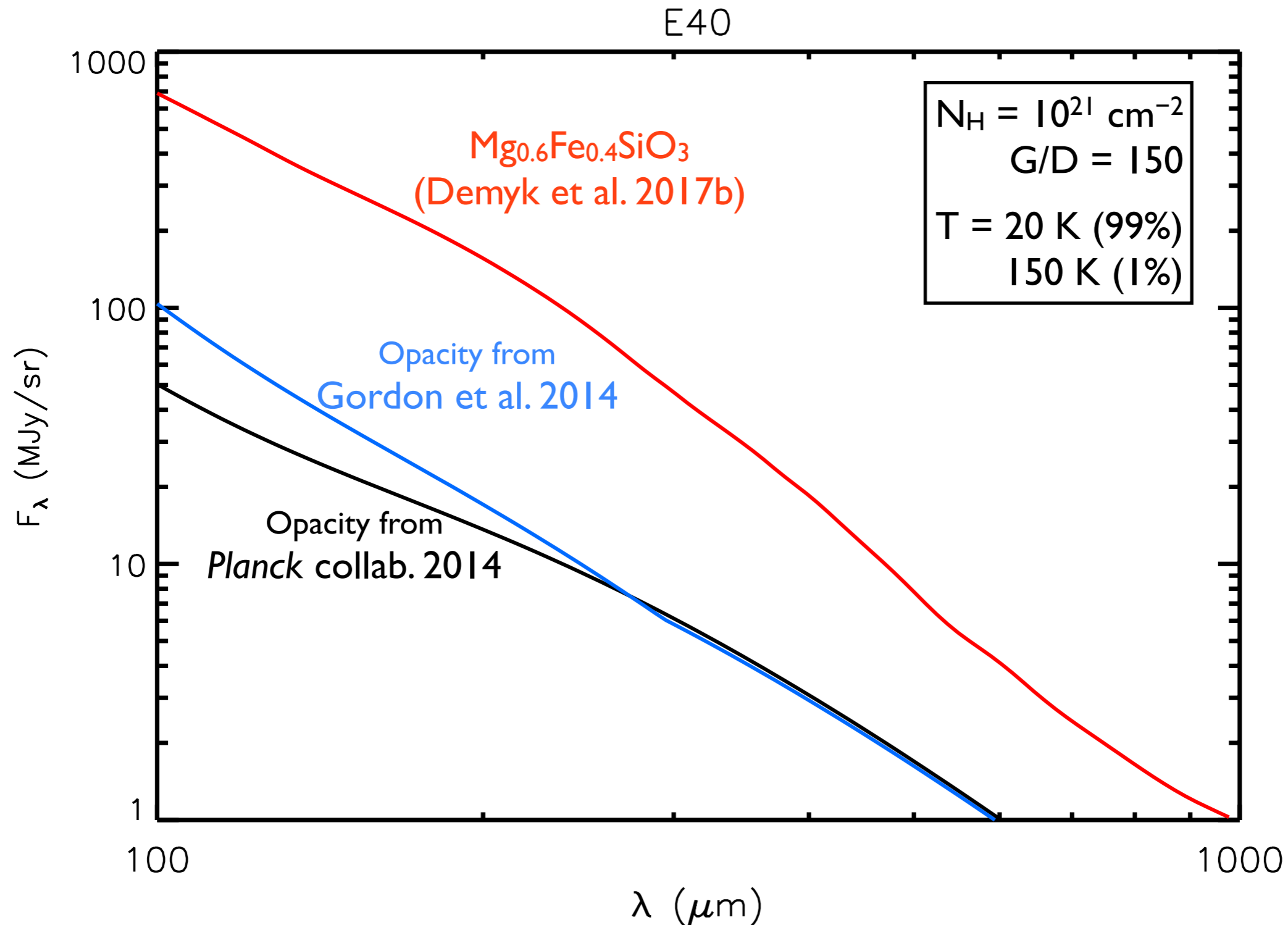
$(1-x)\text{MgO}-(x)\text{SiO}_2$



Lab results vs. observations



Lab results vs. (synthetic) observations



Work by Peter Scicluna, $z = 1$

